Exploiting the Wealth of Theoretical Models in Astronomy with Virtual Observatory

Petr Škoda

Astronomical Institute Academy of Sciences
Ondřejov
Czech Republic

MHD Seminar, Ondřejov, Czech Republic, 4-th March 2010

Outline of the Talk

- VO – the hidden revolution in astronomy
- Data Avalanche in astronomy
- History of VO
- Basic principles of technology
- VO Tools
- VO and Society
- Theoretical Models in VO
- VO Science
- Demos

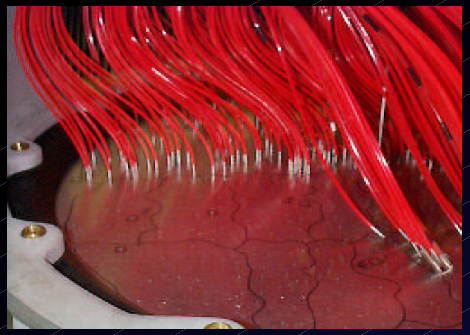
VO – The Hidden Revolution

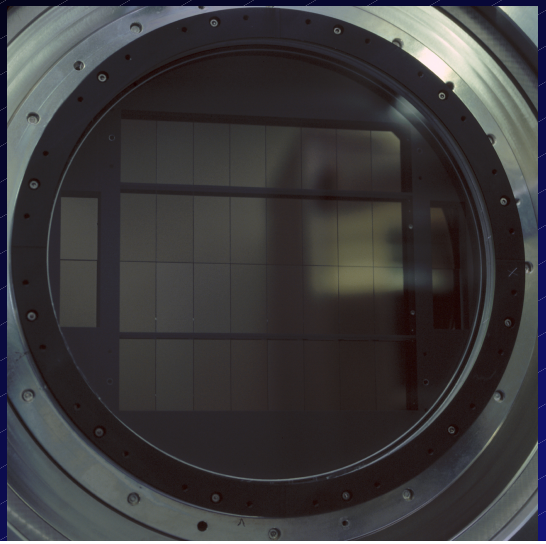
- VO is the radical change of the paradigma of the work of the scientists – effectiveness !!!
- Everyday question (what, where, format, units)
- Everyone is using it – but not stated (> 5 years)
  - CDS (Simbad, Aladin, Vizier), NASA, ESA archives
  - All looks like „ONLY“ another WEBs, client apps
- Scientists are conservative (don't like change)
  - The fear of buzzword VO (multispec, large scale)
- Computer literacy – obligatory (part of job)

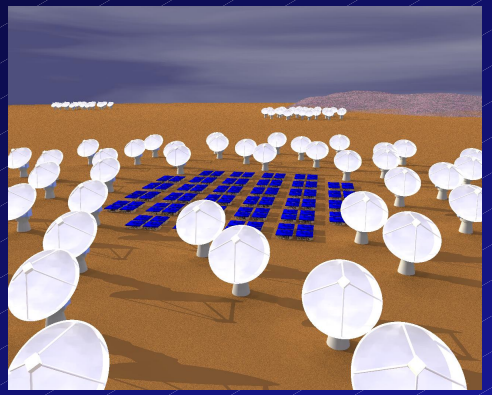
Analogy between VO and WWW

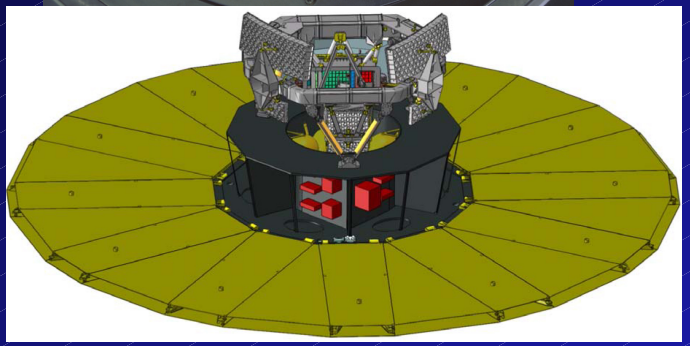
- Linking HYPERTEXT/DATA among servers
- Synergy effect of GLOBAL NET (Gopher,WAIS)
- Powerfull SEARCH (VERONICA – GOOGLE)
- DISTRIBUTED but CENTRAL Steering Organisation (W3C/IVOA)
- Recommendations = „Obligatory“ Standards
- Astronomers in forefront of development
- Scepticism (usefullness for my field ???)
- Steep Growth – average user can use it without knowledge of principles (effectivity, habits)

Data Avalanche

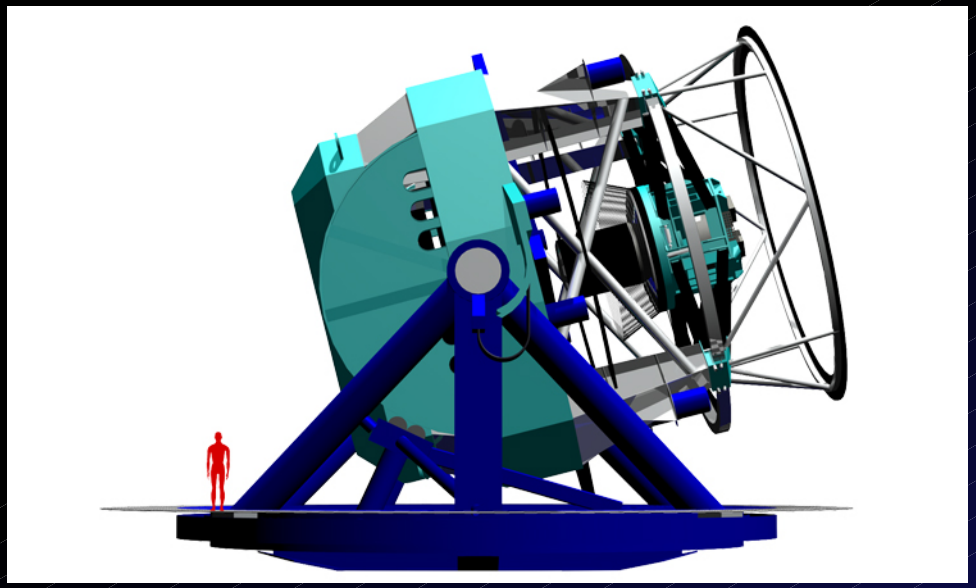


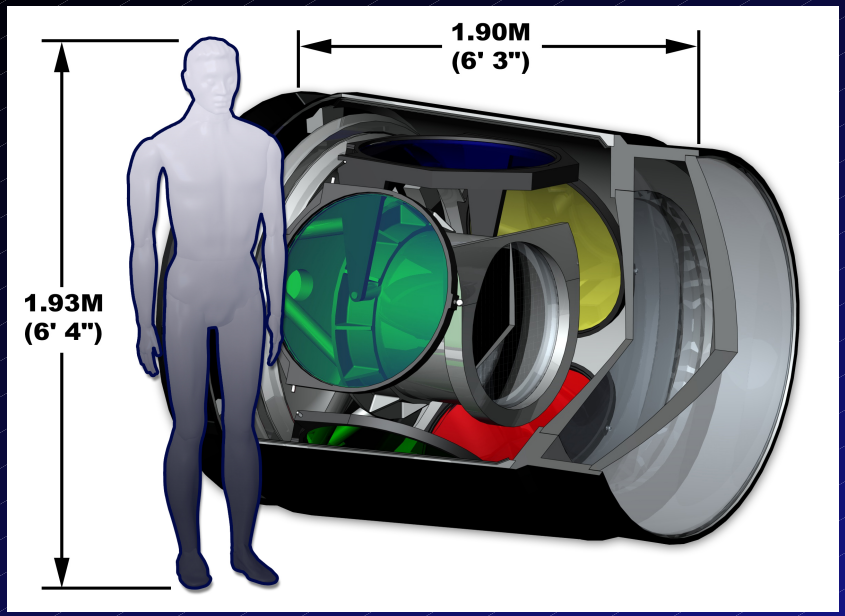


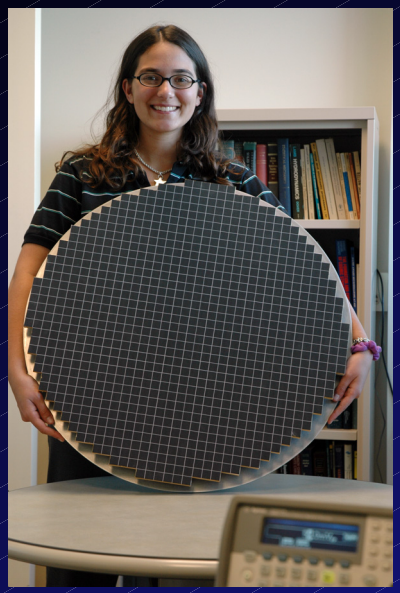




LSST (8.4m)





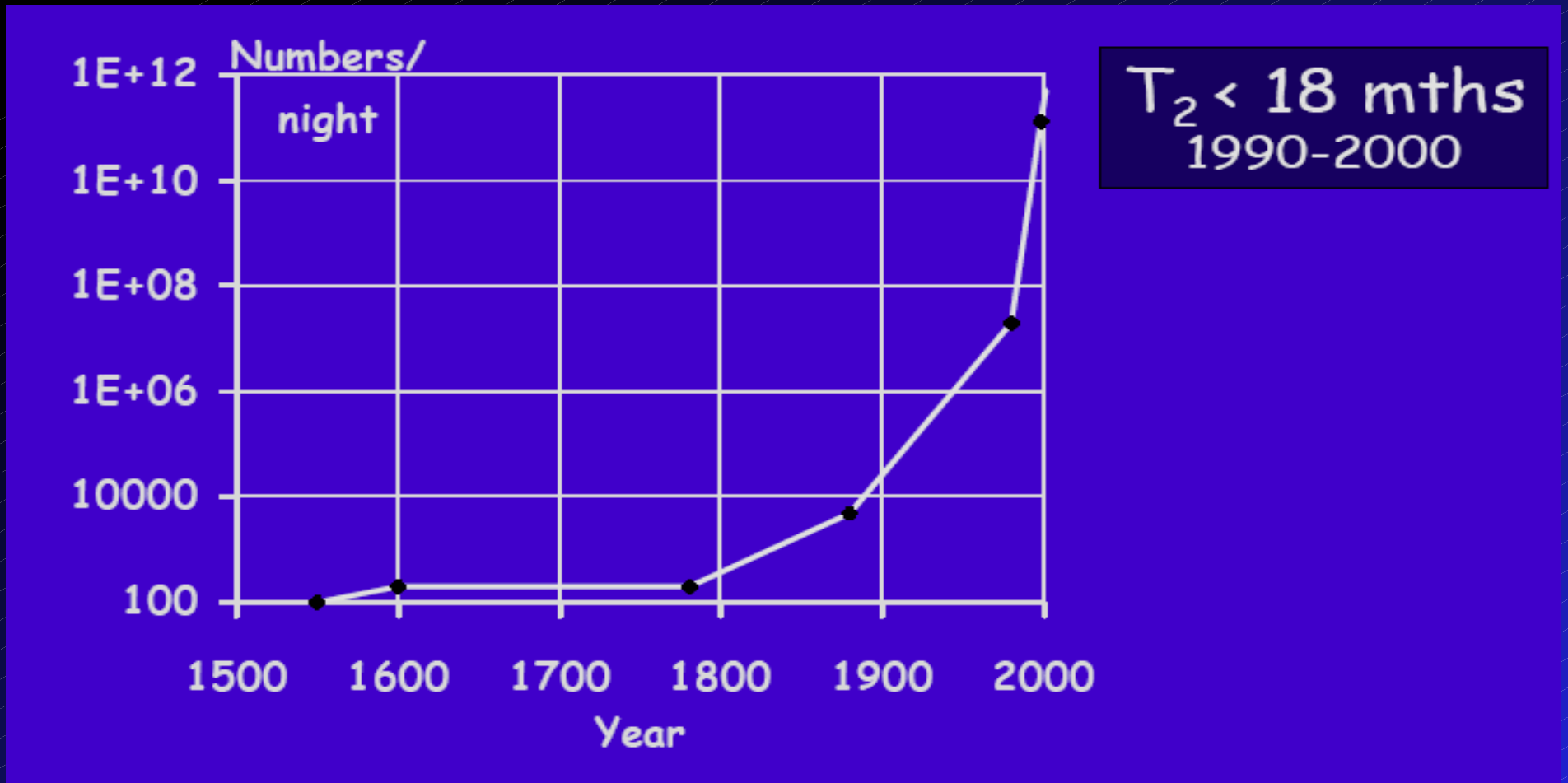


200 CCD 4kx4k,
32 channels (6400)
3.2 Gpix every 20 sec
64cm diameter
3.5 deg FOV
30 TB/night
2 TFLOPS
detection of changes
within 60sec

Data Avalanche

Moore law for chips –doubling 1.5 year

Data in astronomy – doubling < 1 yr ! (1000/10 yr)



Large Scale Data

- *Huge surveys*:100 million sources at < 3000 sources per night ⇒> 100 years to identify them
- *Huge data collections*: download and data analysis on desktop problematic/impossible.
- Example: downloading Sloan Digital Sky Survey (SDSS) DR6 data:
  - images (10 Terabytes) ⇒~ 3 months at 10 Mbps
  - catalogues (2 Terabytes) ⇒~ 3 weeks
  - on DVDs ⇒~ 2,100 of them
- And *data analysis*?? (similar size for MACHO, 2MASS etc)

History of VO

Success of IUE/HST archives

idea of the VO - end 2000

Federation of archives (MAST, NED)

- unified IF, data format for transport

Huge data – distributed processing

- GRID - started in HEP (accelerator science)

Multispectral research : radio---gamma

- Virtual Universe (UK), AstroVirTel (ESO)

Data for SDSS, SIMBAD, NED – key research

Virtual Observatory : Key Definitions

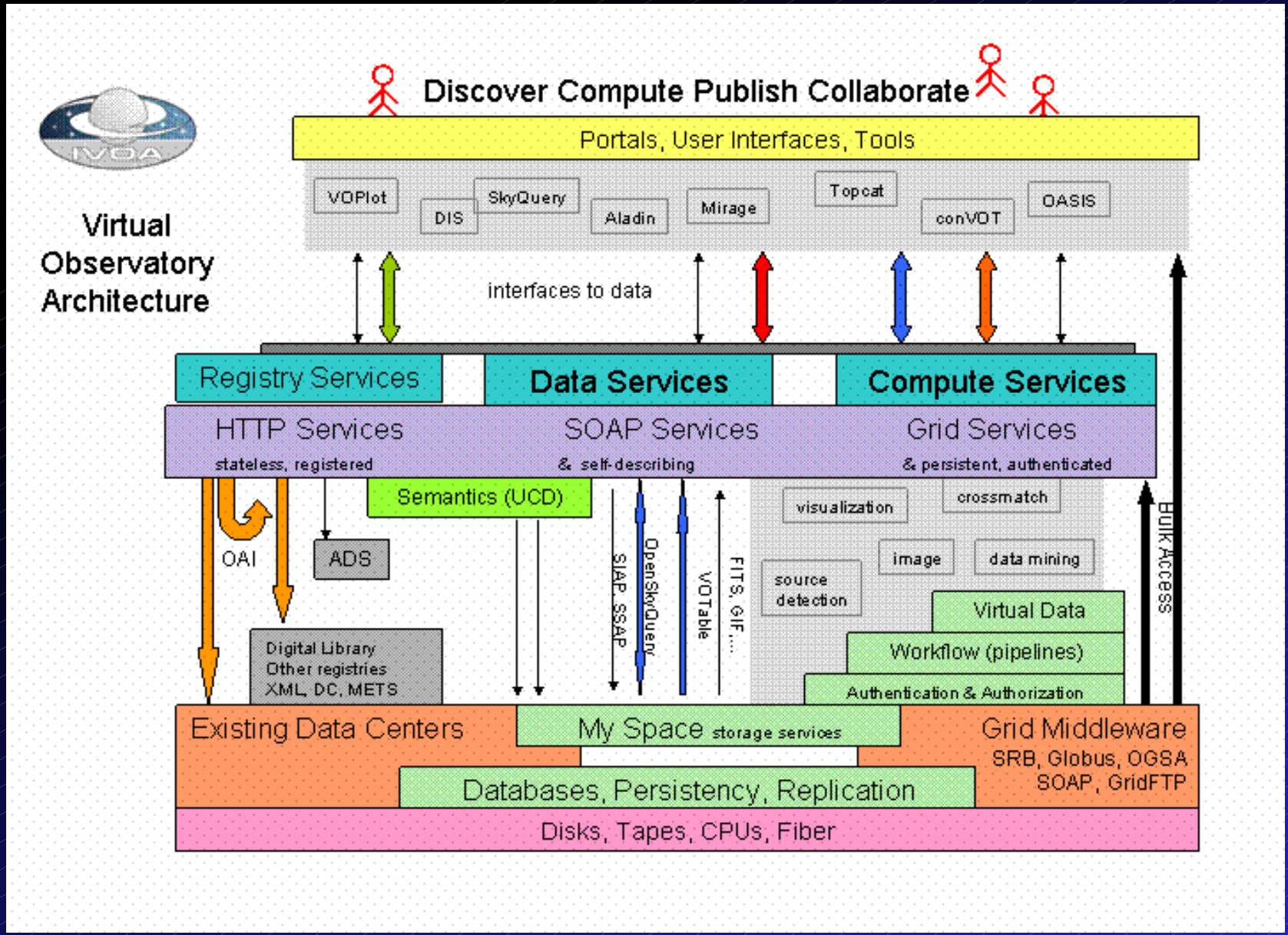
- "*The Virtual Observatory will be a system that allows astronomers to interrogate multiple data centers in a seamless and transparent way, which provides new powerful analysis and visualization tools within that system, and which gives data centers a standard framework for publishing and delivering services using their data*".
- Standardization of data and metadata, and of data exchange methods.
- Registry, listing available services and what can be done with them.

R.J.Hanisch, P.J.Quinn, in "IVOA – Guidelines for participation"

VO Paradigma

- METADATA (name of column), ontologies (name)
- Unique format (VOTable – e.g Vizier)
- Transparent search, download, conversion
- Query for data – processing done on servers
- Federation of astronomical archives (protocols)
- Unified presentation – automatic units conversion (A,MeV,MHz->nm), Wm^-2s^-1 → Jy)
- Background computing on GRIDS
- Multiwavelength approach (SED)

Architecture of VO



Technology of VO

Unified data format– VOTable, UCD (Vizier)

Transparent transport (SOAP , REST (youtube))

Web services (WS) e-commerce, B2B, J2EE, .Net

VOregistry (DNS like) Google for data+WS

protocols (CGI)

- ConeSearch (searching in circle on sky)
- SIAP (Simple Image Access Protocol)
- SSAP(Simple Spectral Access Protocol)
- SLAP(Simple Line Access Protocol)
- TAP (Table Access Protocol)
- VOEVENT (transients, robotic telescopes,Sun)

Technology of VO

ADQL (Astronomical Data Query Language)

XMATCH, REGION (2 catalogues - shifted)

Application interoperabilty – PLASTIC, SAMP

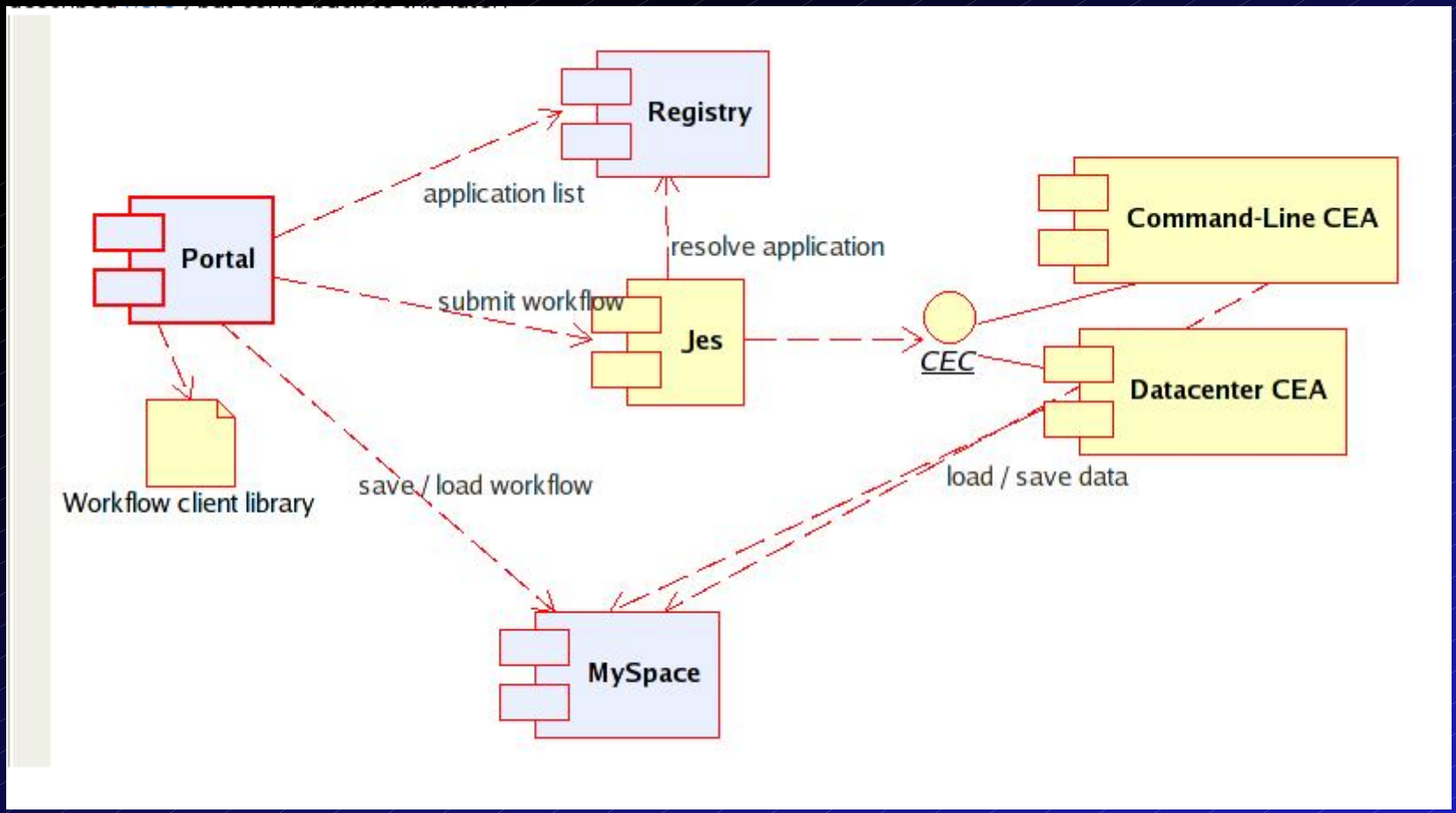
- Allows develop applications as bricks
- sending VOTABLES (catalogue-spectra-images)

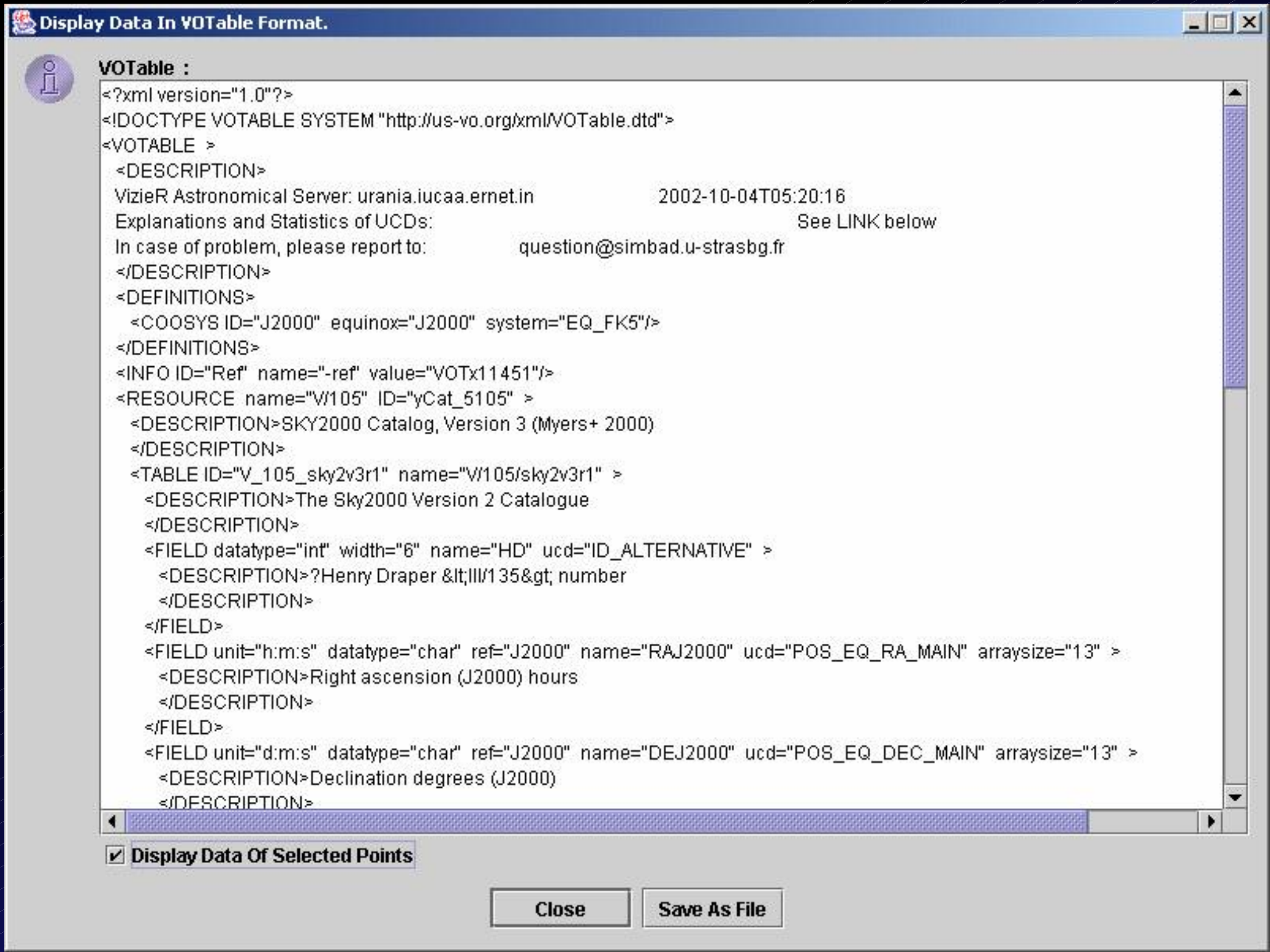
Commercial interest (GoogleSky, MS WWT)

Planetariums, Outreach (Stellarium)

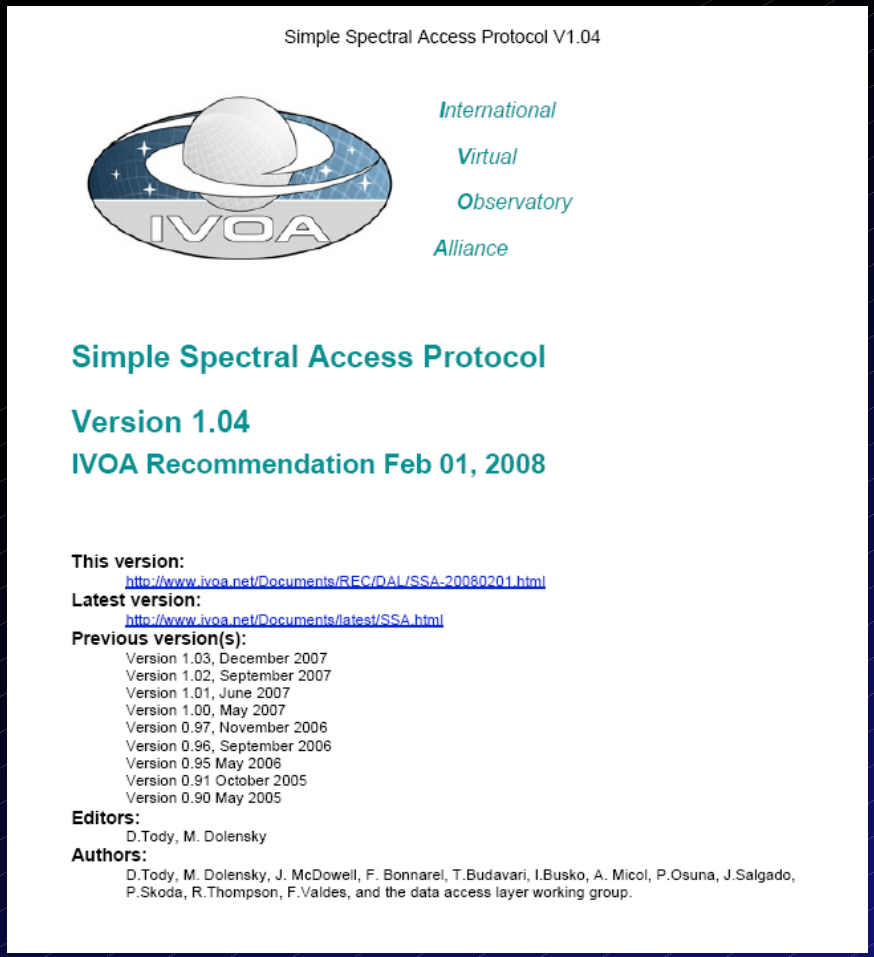
Workflows - Astrogrid

Running remote services – e.g. Sextractor, CASJobs, AstroNeural MLP....




Display Data In VOTable Format.
VOTable :
<?xml version="1.0"?>
<!DOCTYPE VOTABLE SYSTEM "http://us-vo.org/xml/VOTable.dtd">
<VOTABLE >
<DESCRIPTION>
VizieR Astronomical Server: urania.iucaa.ernet.in 2002-10-04T05:20:16
Explanations and Statistics of UCDs: See LINK below
In case of problem, please report to: question@simbad.u-strasbg.fr
</DESCRIPTION>
<DEFINITIONS>
<COOSYS ID="J2000" equinox="J2000" system="EQ_FK5"/>
</DEFINITIONS>
<INFO ID="Ref" name="-ref" value="VOTx11451"/>
<RESOURCE name="V/105" ID="yCat_5105" >
<DESCRIPTION>SKY2000 Catalog, Version 3 (Myers+ 2000)
</DESCRIPTION>
<TABLE ID="V_105_sky2v3r1" name="V/105/sky2v3r1" >
<DESCRIPTION>The Sky2000 Version 2 Catalogue
</DESCRIPTION>
<FIELD datatype="int" width="6" name="HD" ucd="ID_ALTERNATIVE" >
<DESCRIPTION>?Henry Draper <III/135> number
</DESCRIPTION>
</FIELD>
<FIELD unit="h:m:s" datatype="char" ref="J2000" name="RAJ2000" ucd="POS_EQ_RA_MAIN" arraysize="13" >
<DESCRIPTION>Right ascension (J2000) hours
</DESCRIPTION>
</FIELD>
<FIELD unit="d:m:s" datatype="char" ref="J2000" name="DEJ2000" ucd="POS_EQ_DEC_MAIN" arraysize="13" >
<DESCRIPTION>Declination degrees (J2000)
</DESCRIPTION>
Display Data Of Selected Points
Close
Save As File

Simple Spectra Access Protocol
Spectral Data Model



Simple Spectral Access Protocol V1.04

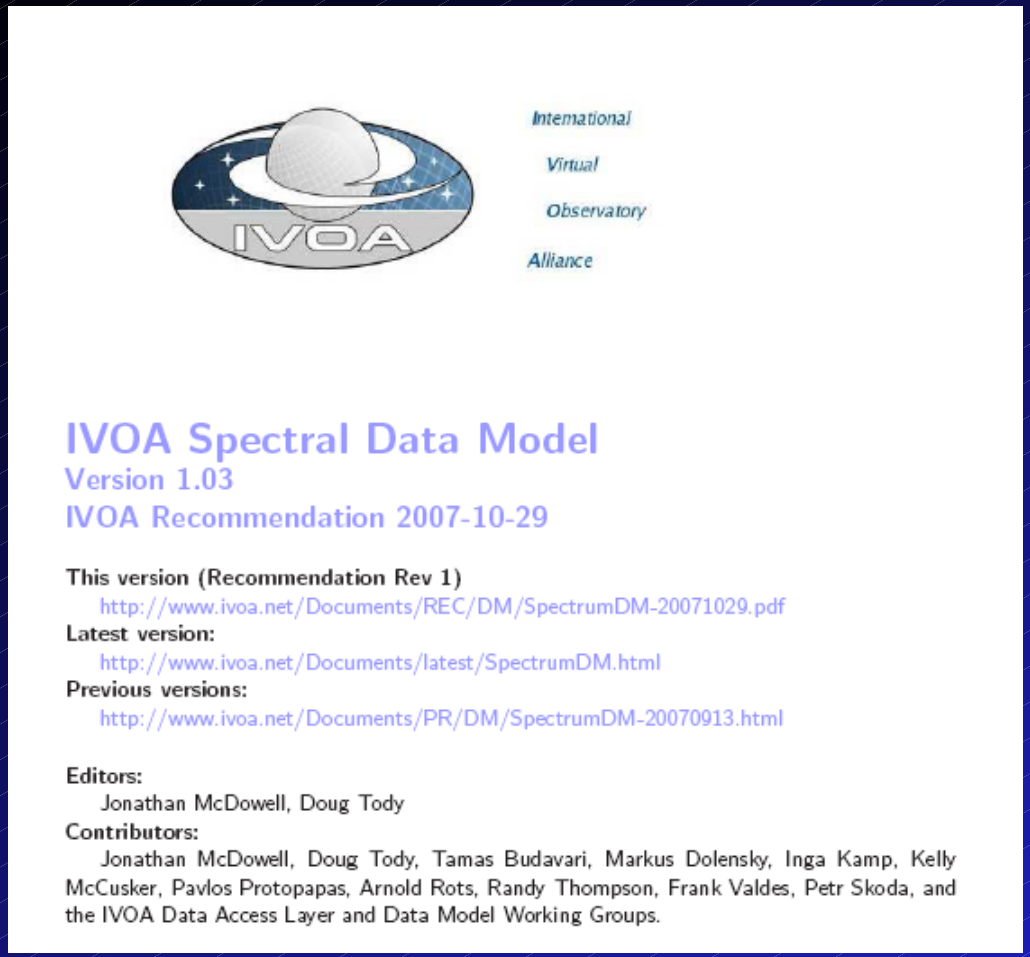
International Virtual Observatory Alliance

Simple Spectral Access Protocol

Version 1.04

IVOA Recommendation Feb 01, 2008

This version:
http://www.ivoa.net/Documents/REC/DAL/SSA-20080201.html
Latest version:
http://www.ivoa.net/Documents/latest/SSA.html
Previous version(s):
Version 1.03, December 2007
Version 1.02, September 2007
Version 1.01, June 2007
Version 1.00, May 2007
Version 0.97, November 2006
Version 0.96, September 2006
Version 0.95 May 2006
Version 0.91 October 2005
Version 0.90 May 2005
Editors:
D.Tody, M. Dolensky
Authors:
D.Tody, M. Dolensky, J. McDowell, F. Bonnarel, T.Budavari, I.Busko, A. Micol, P.Osuna, J.Salgado, P.Skoda, R.Thompson, F.Valdes, and the data access layer working group.



International Virtual Observatory Alliance

IVOA Spectral Data Model

Version 1.03

IVOA Recommendation 2007-10-29

This version (Recommendation Rev 1)
http://www.ivoa.net/Documents/REC/DM/SpectrumDM-20071029.pdf
Latest version:
http://www.ivoa.net/Documents/latest/SpectrumDM.html
Previous versions:
http://www.ivoa.net/Documents/PR/DM/SpectrumDM-20070913.html

Editors:
Jonathan McDowell, Doug Tody
Contributors:
Jonathan McDowell, Doug Tody, Tamas Budavari, Markus Dolensky, Inga Kamp, Kelly McCusker, Pavlos Protopapas, Arnold Rots, Randy Thompson, Frank Valdes, Petr Skoda, and the IVOA Data Access Layer and Data Model Working Groups.

SSAP Parameters

4.1.1 Mandatory Query Parameters

The following parameters **must** be implemented by a compliant service:

Parameter	*Sample value*	*Physical unit*	*Datatype*
POS	52,-27.8	degrees; defaults to ICRS	string
SIZE	0.05	degrees	double
BAND	2.7E-7/0.13	meters	string
TIME	1998-05-21/1999	ISO 8601 UTC	string
FORMAT	votable	-	string

4.1.2 Recommended and Optional Query Parameters

Parameter	Sample value	Unit	Req	Datatype
APERTURE	0.00028 (=1°)	degrees	OPT	double
SPECRP	2000	λ/dλ	REC	double
SPATRES	0.05	degrees	REC	double
TIMERES	31536000 (=1yr)	seconds	OPT	double
SNR	5.0	dimensionless	OPT	double
REDSHIFT	1.3/3.0	dimensionless	OPT	string
VARAMPL	0.77	dimensionless	OPT	string
TARGETNAME	mars		OPT	string
TARGETCLASS	star		OPT	string
FLUXCALIB	relative		OPT	string
WAVECALIB	absolute		OPT	string
PUBDID	ADS/col#R5983		REC	string
CREATORDID	ivo://auth/col#R1234		REC	string
COLLECTION	SDSS-DR5		REC	string
TOP	20	dimensionless	REC	int
MAXREC	5000		REC	string
MTIME	2005-01-01/2006-01-01	ISO 8601	REC	string
COMPRESS	true		REC	boolean
RUNID			REC	string

The spatial, spectral, and time resolution of the data may all be used as query constraints

IVOA


Aus-VO
EURO VO
China-VO
VO-i
Astro Grid
CVO
RVO
SVO
NVO
IVOA
VObs.it
ArVO
JVO
HVO
GAVO
France

VO-enabled tools

Aladin

VOPlot

TOPCAT

VOSpec

SpecView

SPLAT

ViSiVO (HPC simulations, cosmology)

VOSED

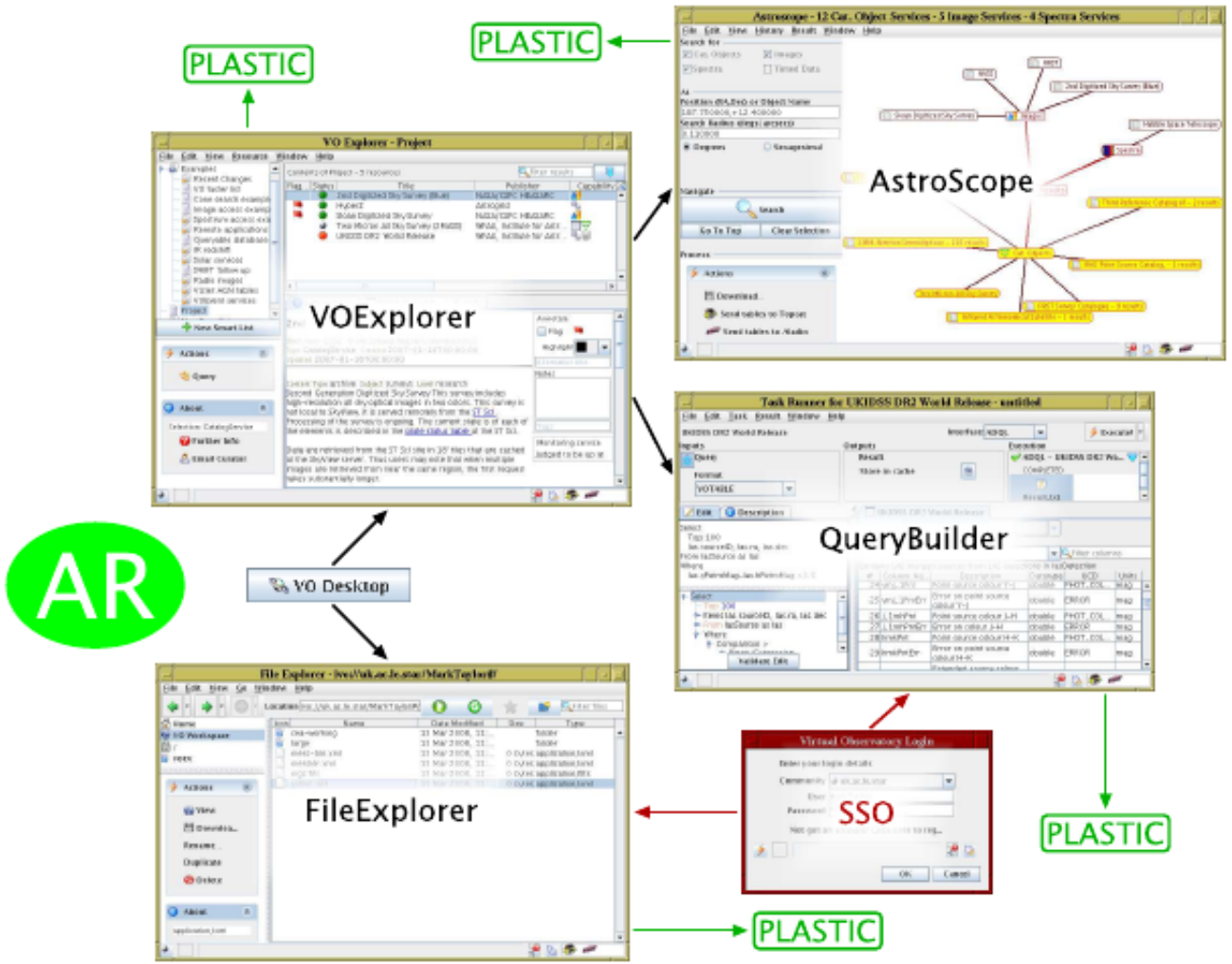
BASTI

SExtractor – WESIX (Web Enabled Source Identification with Cross Matching)

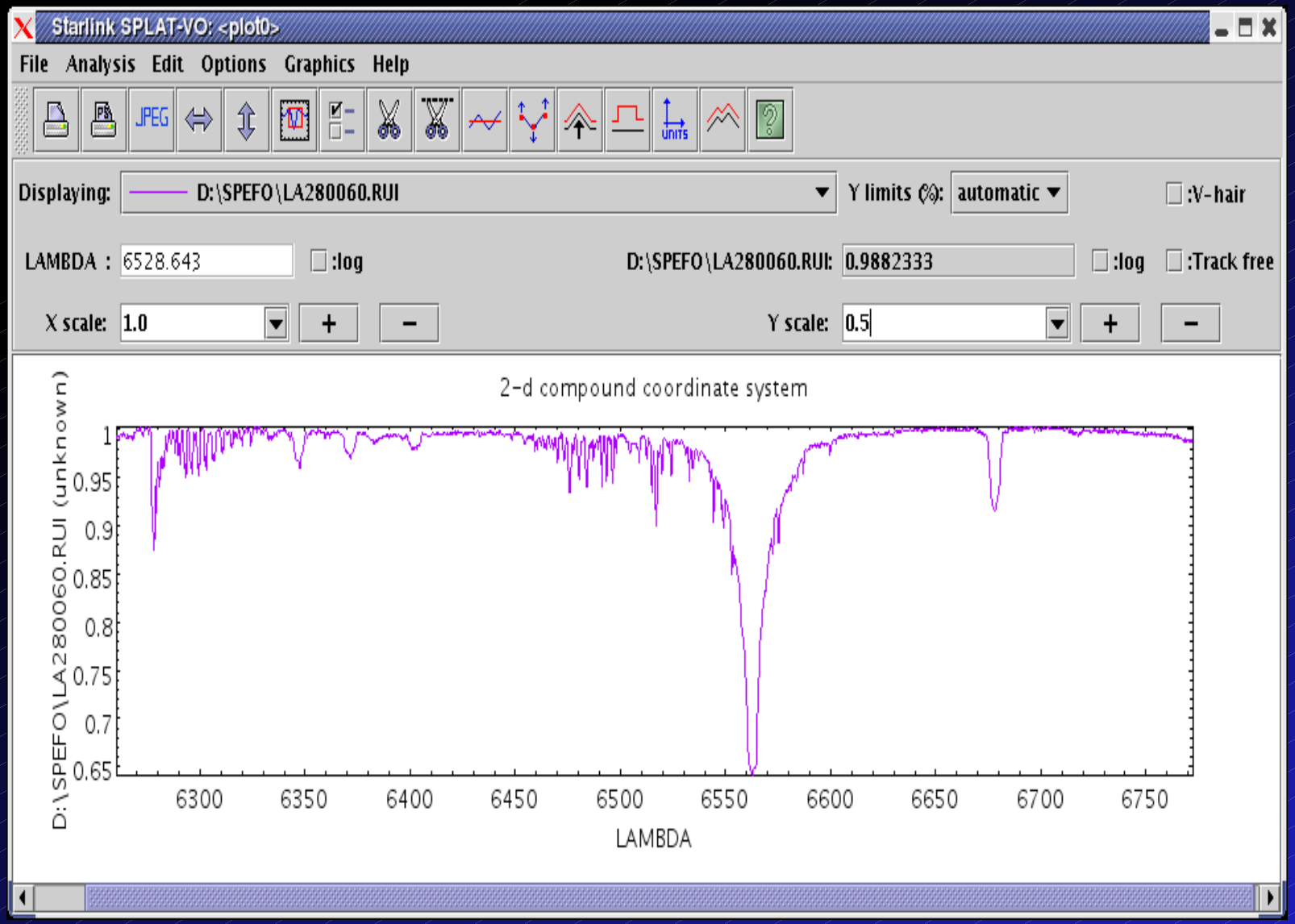
Period04 (since 18.9.08) - PLASTIC

AstroGRID VODesktop

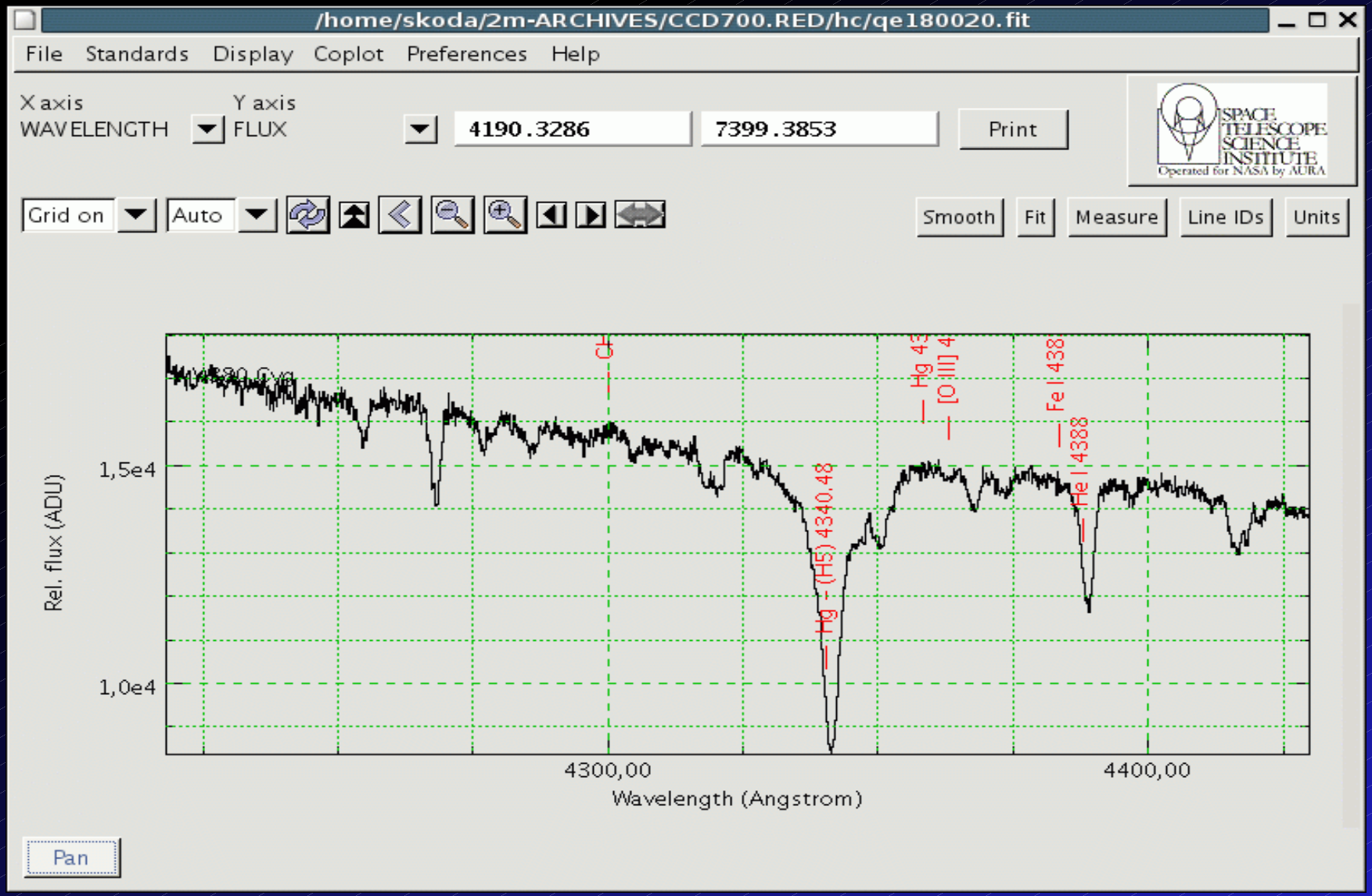
VODesktop Overview



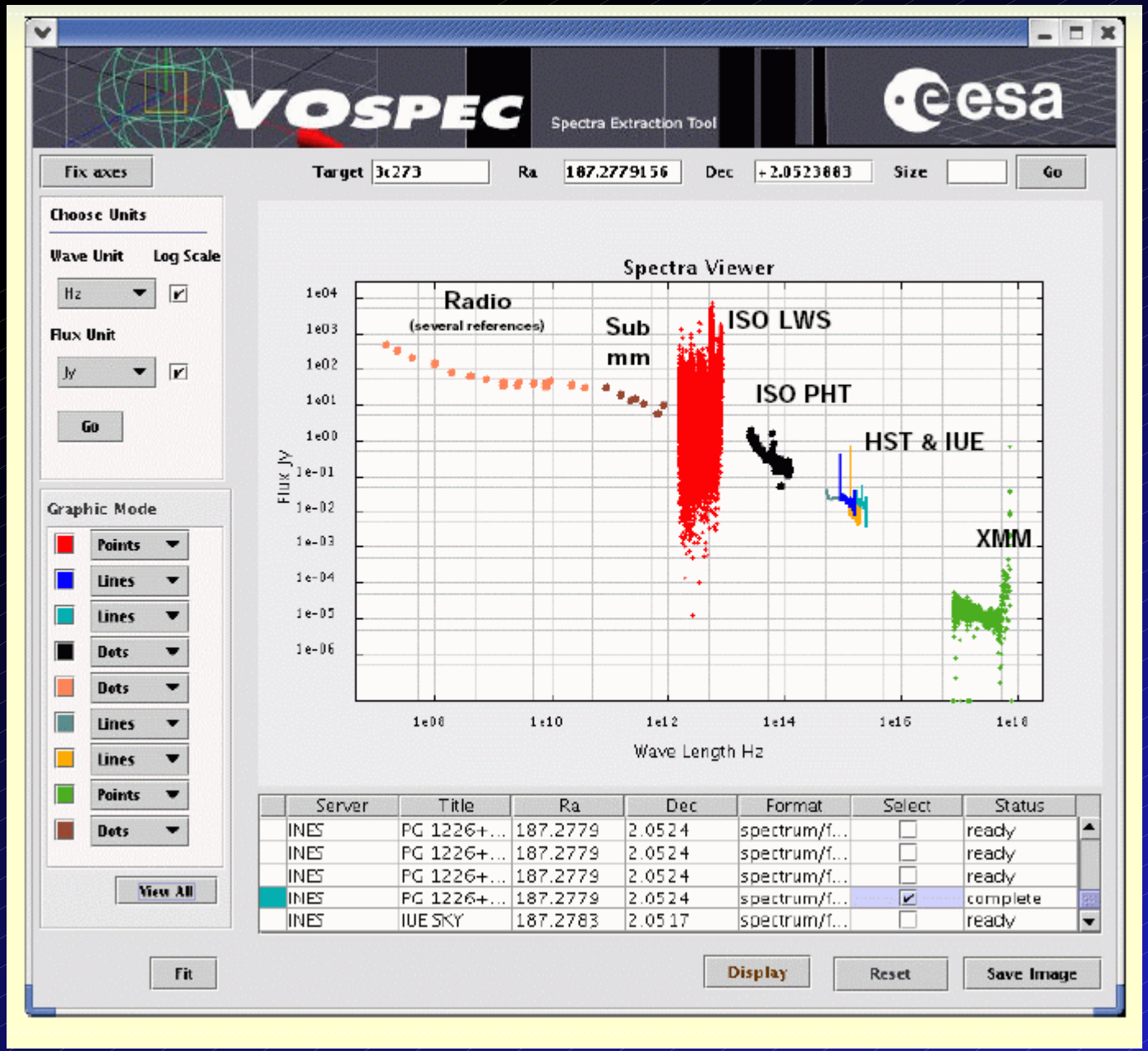
SPLAT-VO (Starlink, JAC)



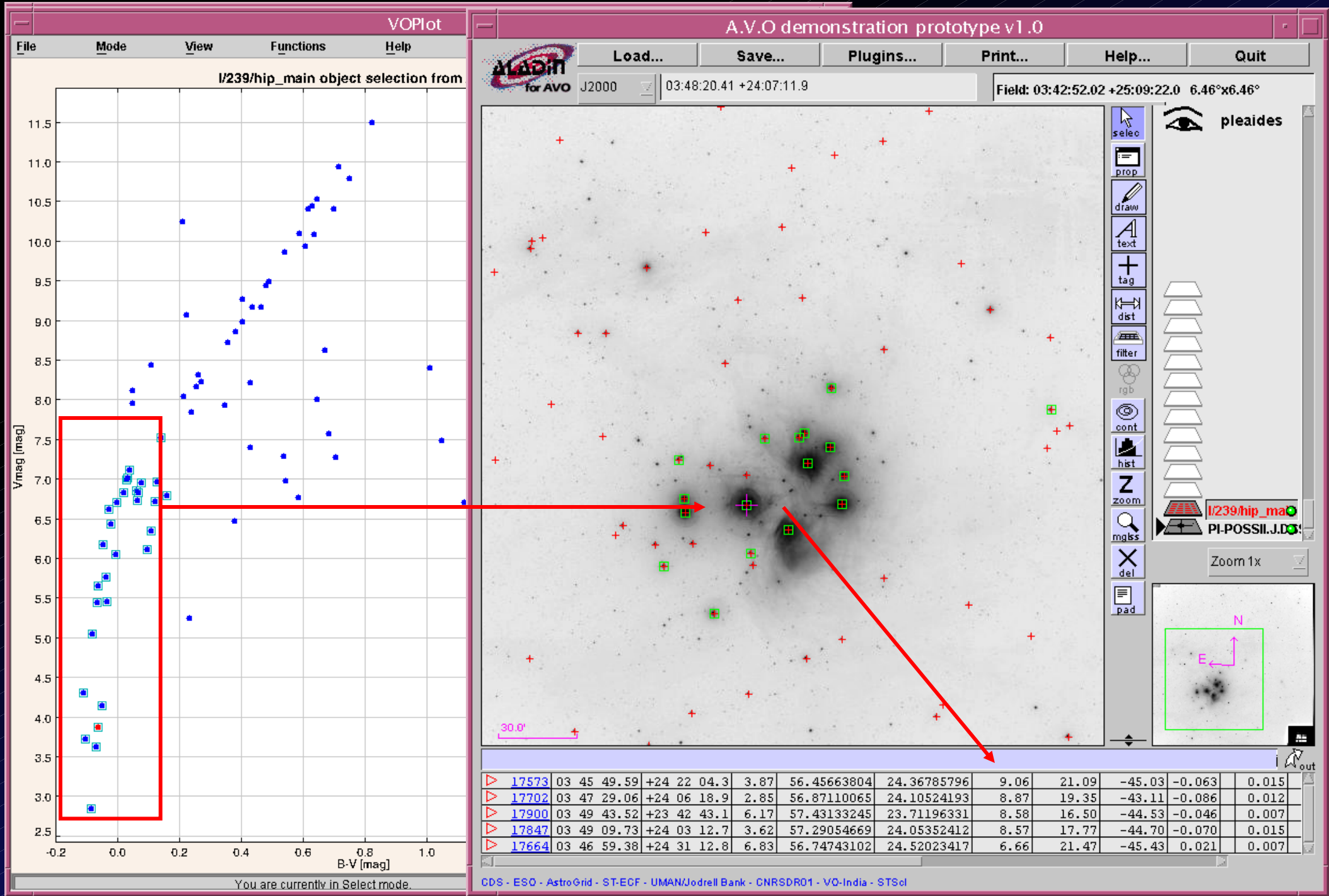
SpecView (STScI)

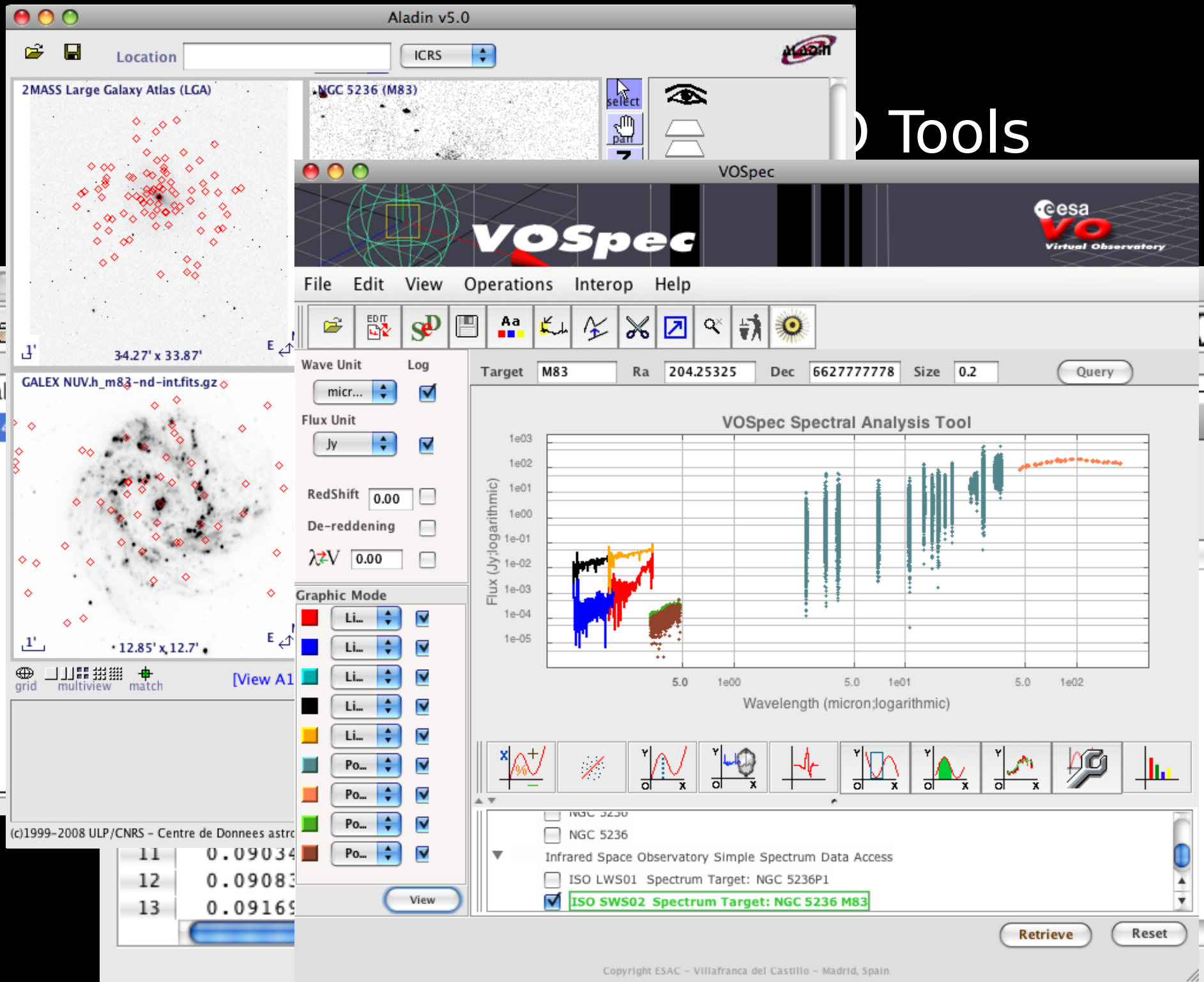


VOspec (ESAC)

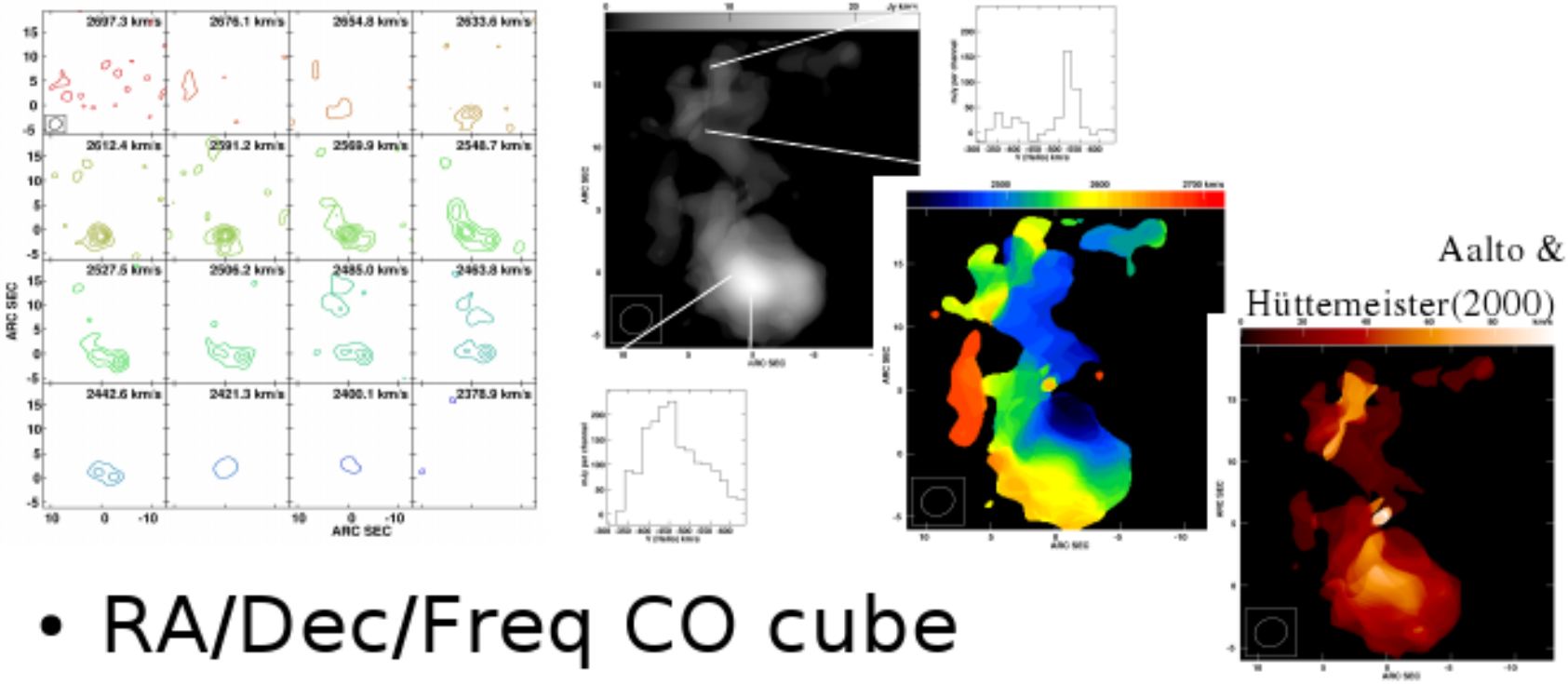

VOSPEC
Spectra Extraction Tool
esa
Fix axes
Target 3c273
Ra 187.2779156
Dec +2.0523883
Size
Go
Choose Units
Wave Unit
Log Scale
Hz
Flux Unit
Jy
Go
Graphic Mode
Points
Lines
Lines
Dots
Dots
Lines
Lines
Points
Dots
View All
Spectra Viewer
Radio
(several references)
Sub
mm
ISO LWS
ISO PHT
HST & IUE
XMM
Flux Jy
Wave Length Hz
Server
Title
Ra
Dec
Format
Select
Status
INES PG 1226+... 187.2779 2.0524 spectrum/f... ready
INES PG 1226+... 187.2779 2.0524 spectrum/f... ready
INES PG 1226+... 187.2779 2.0524 spectrum/f... ready
INES PG 1226+... 187.2779 2.0524 spectrum/f... complete
INES IUE SKY 187.2783 2.0517 spectrum/f... ready
Fit
Display
Reset
Save Image

Colour-magnitude diagram



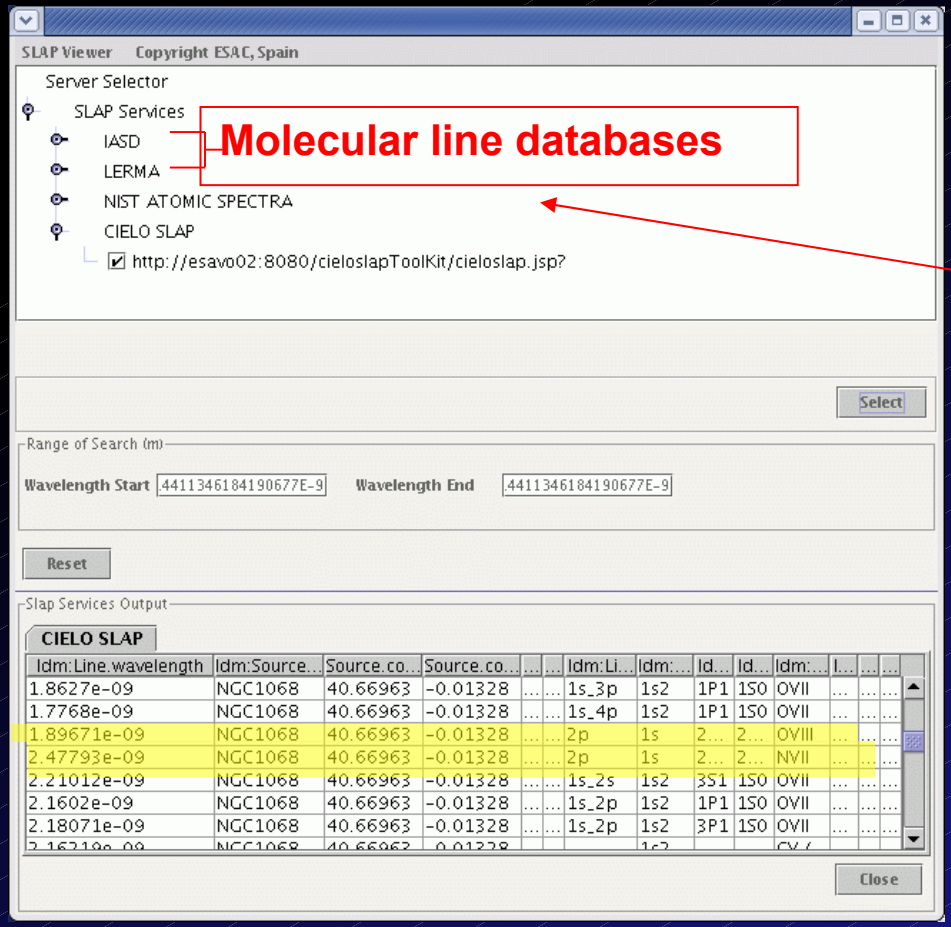

Aladin v5.0
Location
ICRS
2MASS Large Galaxy Atlas (LGA)
NGC 5236 (M83)
select
34.27' x 33.87'
GALEX NUV.h_m83-nd-int.fits.gz
12.85' x 12.7'
grid
multiview
match
[View A1
(c)1999-2008 ULP/CNRS - Centre de Donnees astro
Tools
VOSpec
VOSpec
esa
Virtual Observatory
File
Edit
View
Operations
Interop
Help
Wave Unit
Log
micr...
Flux Unit
Jy
RedShift
0.00
De-reddening
0.00
Graphic Mode
Target
M83
Ra
204.25325
Dec
6627777778
Size
0.2
Query
VOSpec Spectral Analysis Tool
Flux (Jy;logarithmic)
Wavelength (micron;logarithmic)
NGC 5236
Infrared Space Observatory Simple Spectrum Data Access
ISO LWS01 Spectrum Target: NGC 5236P1
ISO SWS02 Spectrum Target: NGC 5236 M83
View
Retrieve
Reset
Copyright ESAC - Villafranca del Castillo - Madrid, Spain
Flux

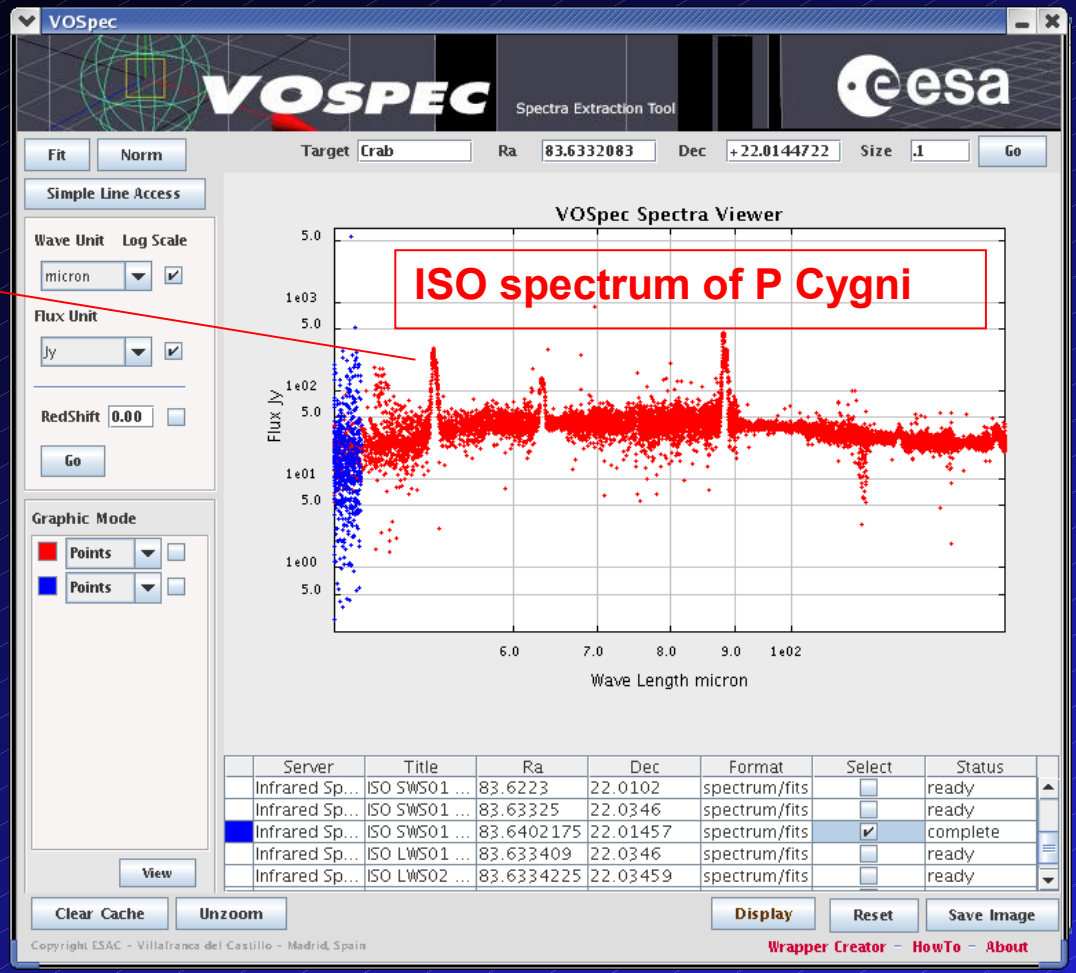
ALMA/IRAM use case



- RA/Dec/Freq CO cube
  - Convert to velocity (LSR, radio convention)
  - Cutouts, simple squashes - VO tools?
  - Smoothed spectra, moments with noise cut-off
    - Specialised server-side pipeline controlled via UWS

CIELO VO – line catalogue SLAP





(IVOA Line Data Model: Dubernet, Osuna et al., in preparation)
(Simple Line Access Protocol: Salgado et al., in preparation)

VO for Atomic and Molecular Data

VAMDC (06/2009-12/2012 FP7)

13 organizations

Virtual Atomic and Molecular Data Centre

VO principles (web services, integration, registry, SAMP, VODesktop, TOPCAT, VOSpec)

(includes VALD extractor, NIST)

extended citation system (all providers acknowledged)

Other VOs

Virtual Solar Observatory

Virtual Solar-Terrestrial Observatory

Virtual Magnetospheric Observatory

Virtual Space Physics Observatory

Virtual Meteor Observatory – not proper - XML

SKYBOT – Minor planets ephemerides (1840-2019)

Interest of climatology, meteorology

New branch of Science = e-Science

Democratization of Science

- Digital Divide (data access free, journals ?)
- International Council for Science (ICS UNO) CODATA (Committee on Data for Science and Technology)
- OECD, UNESCO
- CASPAR

  Cultural, Artistic and Scientific knowledge for Preservation, Access and Retrieval

  Digital curation centers

- ADS and VO (links to ivo://, metadata, ontologies – understanding, semantic web)
- Archive importance: 5x IUE , 3x HST results from archives than PI articles
- Effectivity – 50% of published data appears in Journals, links to data automatic ?

Objections to VO

Data quality – garbage in - garbage out

How and whom to give credit ? (button)

embedded ivo:// data in ApJ

VO for dissemination only

- technology for OPTICON, nextgen
- Virtual science – VO technology

VO only for public data ! Proprietary ?

(data jealousy)

- local archive - available data marked

The Astronomer's Data Manifesto

at 26 IAU GA Prague SPS3

(a) All significant tables, images, and spectra published in journals should appear inastronomical data centres.

(b) All data obtained with publicly-funded observatories should, after appropriate pro-prietary periods, be placed in the public domain.

(c) In any new major astronomical construction project, the data processing, storage,migration, and management requirements should be built in at an early stage of the project plan, and costed along with other parts of the project.

(d) Astronomers in all countries should have the same access to astronomical data and information.

(e) Legacy astronomical data can be valuable, and high-priority legacy data should be preserved and stored in digital form in the data centres.

(f) The IAU should work with other international organisations to achieve our common goals and learn from our colleagues in other fields. ”

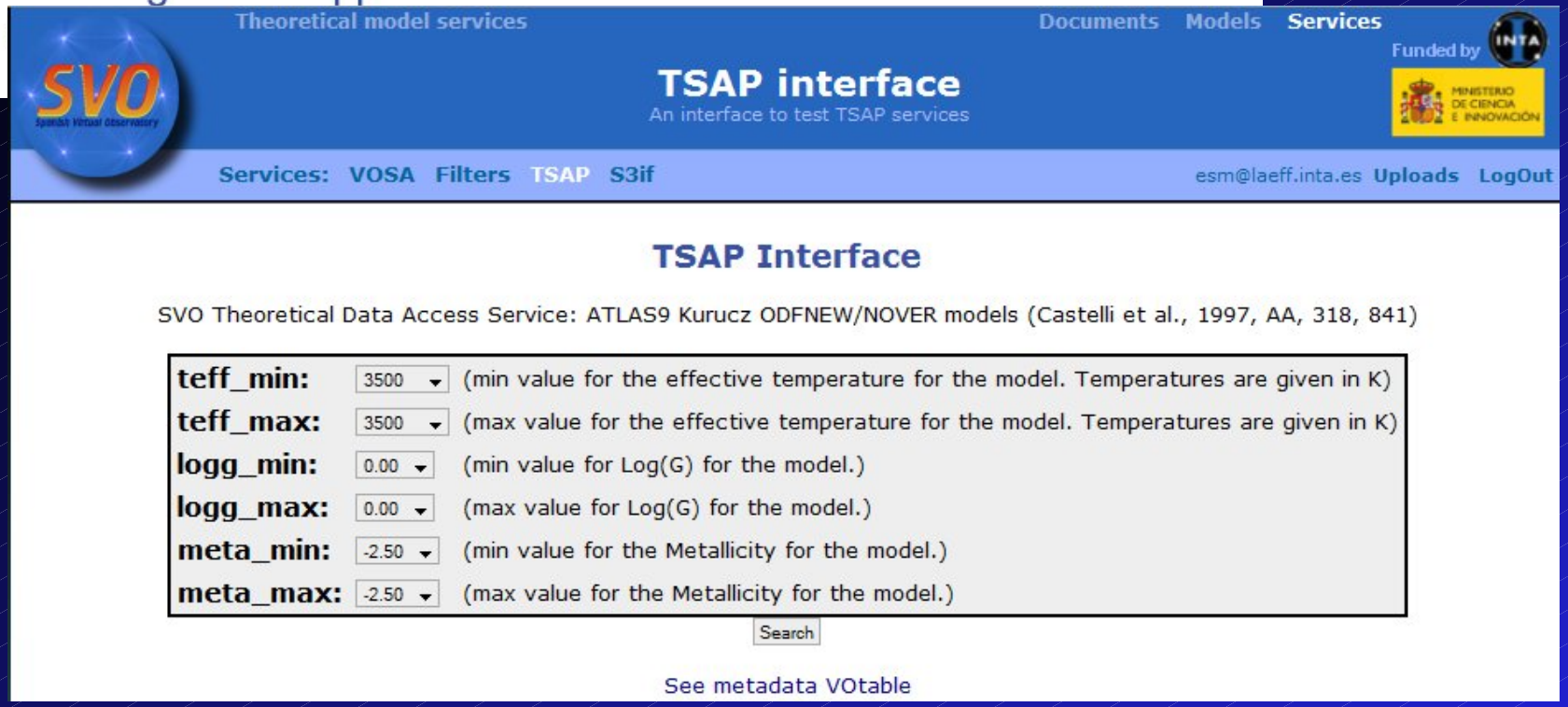
Theory VO (TVO)

- Methods of VO (parameters in DB, SQL...) for study of results of simulations , catalogues of simulated objects like SDSS...(PCA)
- Browsing of simulation space along different axes – parameters, regions...
- Evolutionary tracks, Photo Dissociation Regions
- Formation of artificial galaxies, clusters – N body models (Millenium Run 10 billions, 25TB)
- Theoretical Spectra (GAVO – Rauch, GRID)

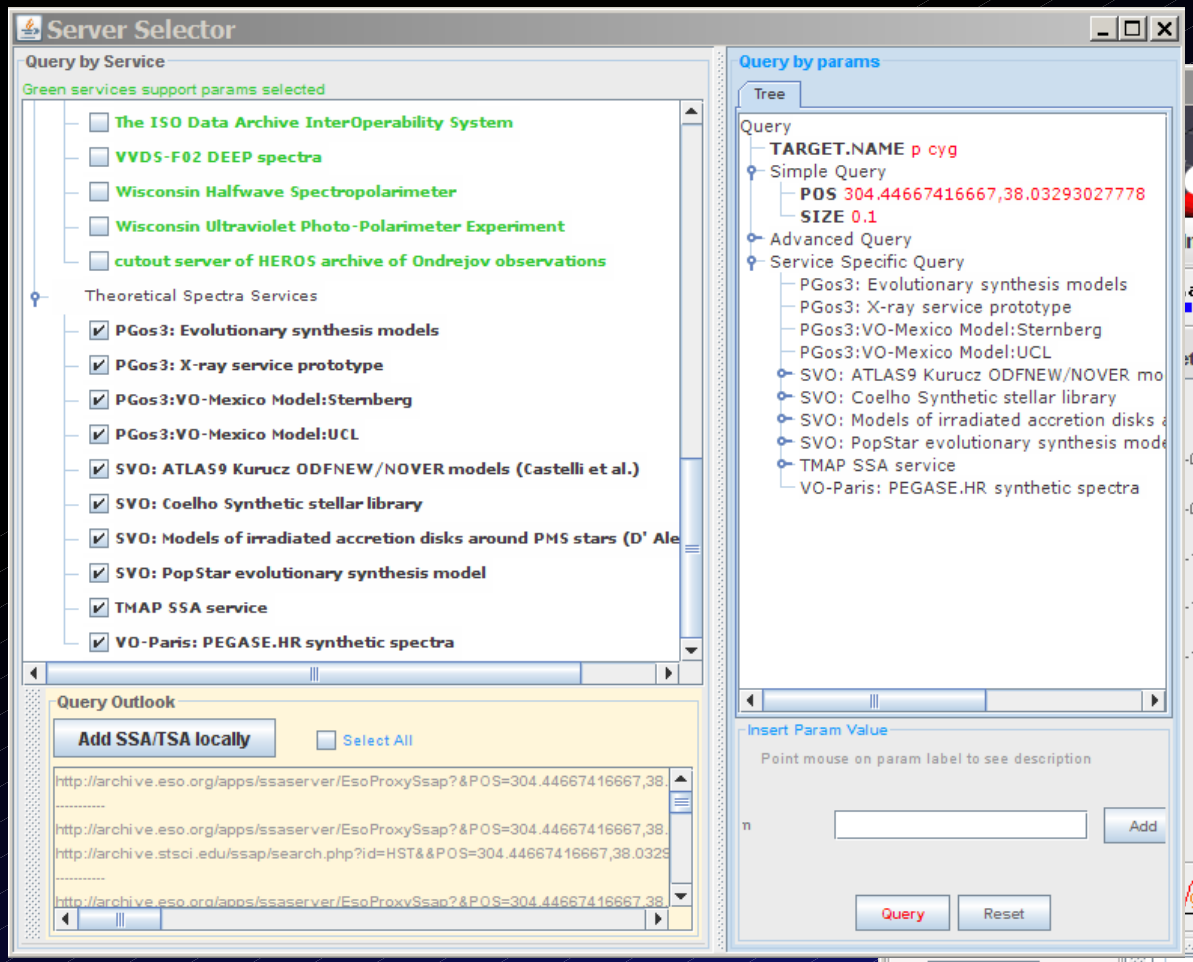
Access protocols in VO: TSAP

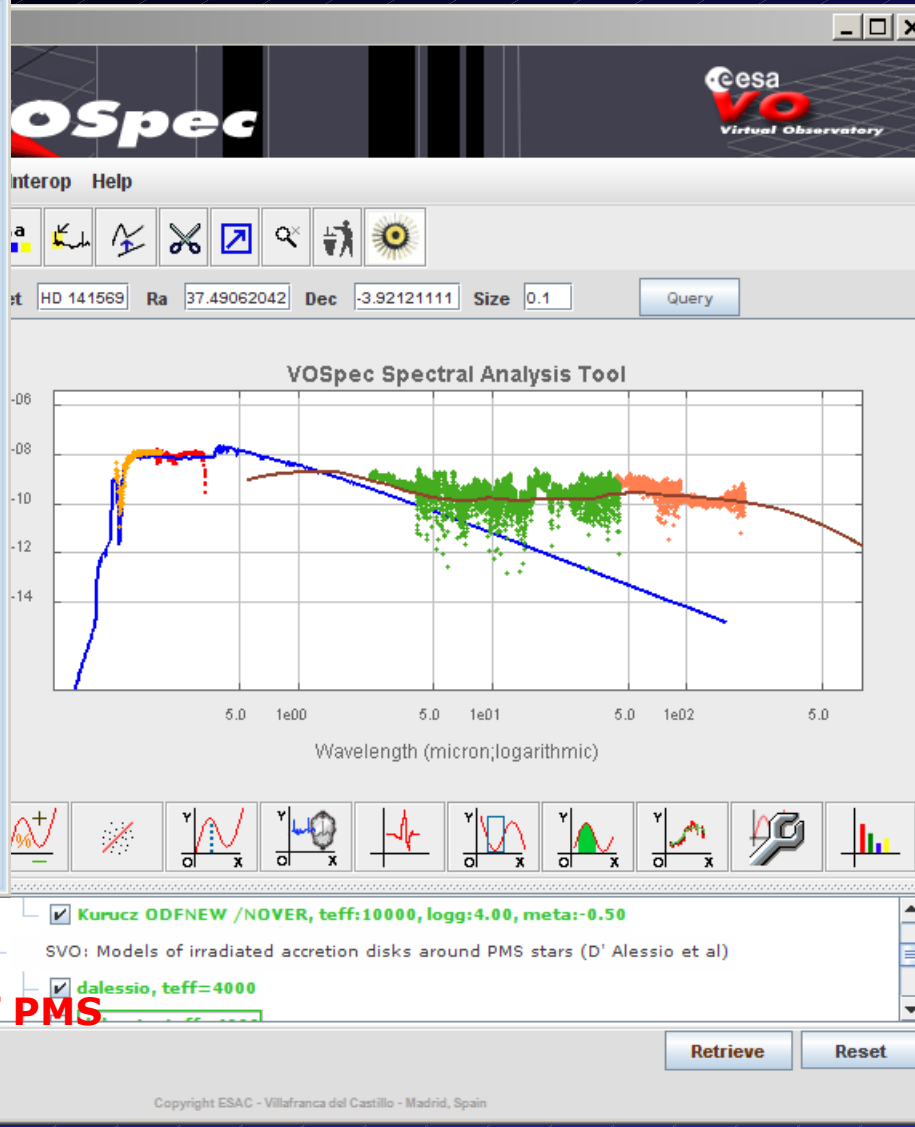
Theoretical models in the VO

- **Theoretical spectra: TSAP**
  - Included in the SSAP standard (use case for theoretical spectra)
  - A simple protocol.
  - Dialog server-application.



VOSpec – models by TSAP



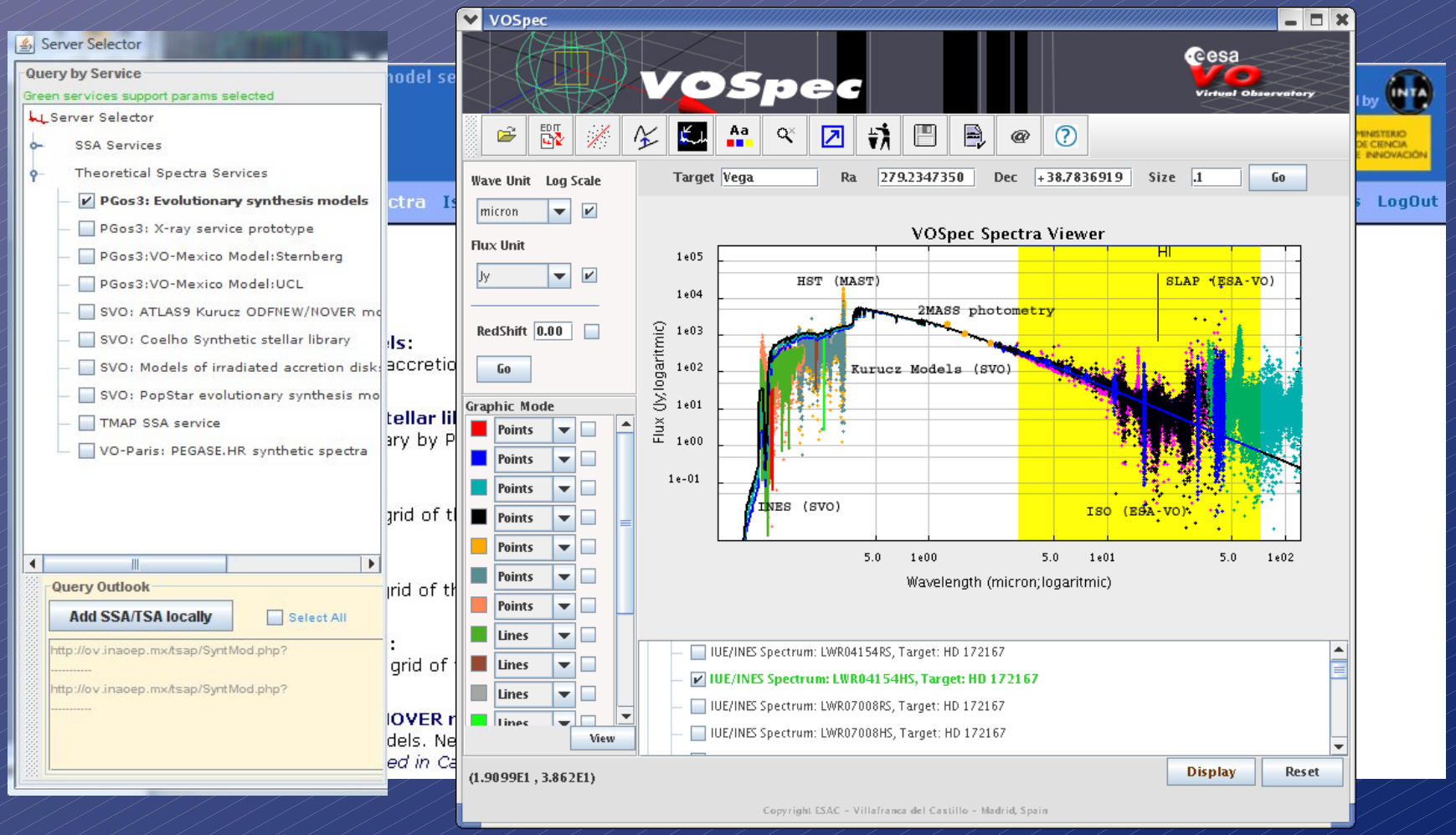


Kurucz stellar model and D'Alessio model of PMS discs fit to UV and IR spectra of HD 141569

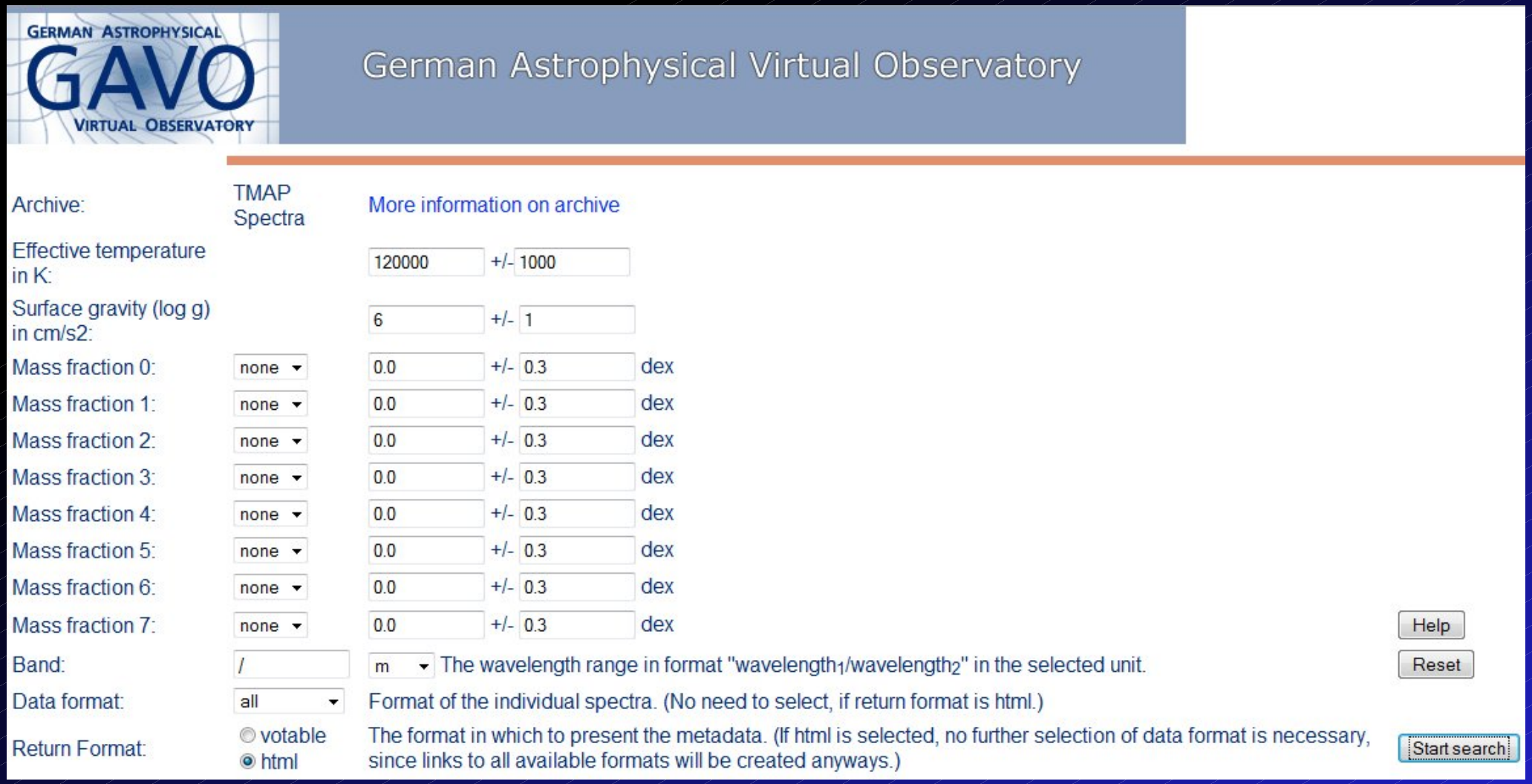
Archives, Theory, VO-Science, DataMining, E&O

Simple Spectral Access Protocol V1.04

Appendix A: Theoretical Spectral Access Use Case



- Other VO Data Centres providing theoretical spectra using TSAP



- PGos3 (Mexico), PEGASE (VO-Paris)

BaSTI database

Micro-simulations inside the VO: the BaSTI case

P. Manzato[1], M. Molinaro[1], F. Gasparo[1], F. Pasian[1], A. Pietrinferni[2], S. Cassisi [2], C. Rodrigo[3], M. Cerviño[4], E. Solano [3]
INAF - SI / Trieste Astronomical Observatory; (2) INAF – Teramo Astronomical Observatory; (3) LAEFF-INTA / Spanish VO; (4) Instituto de Astrofisica de Andalucía – CSIC / Spanish VO

S3P (Simple Self-Described Service Protocol) implementations

In collaboration with SVO (the Spanish Virtual Observatory) we presented S3P in the last IVOA Interoperability Meeting. S3P (Simple, Self-described Service) is a protocol oriented to handle theoretical data in the VO framework. It is based in the ability of the data server to describe itself in a simple standardized way.

This is a step by step protocol:

1 step: the service described it self (input and output parameters);

http://myservice.com/s3.php?format=metadata

2 step: http query and response in VOTable format;

http://myservice.com/s3.php?param1=value1&param2=value2...

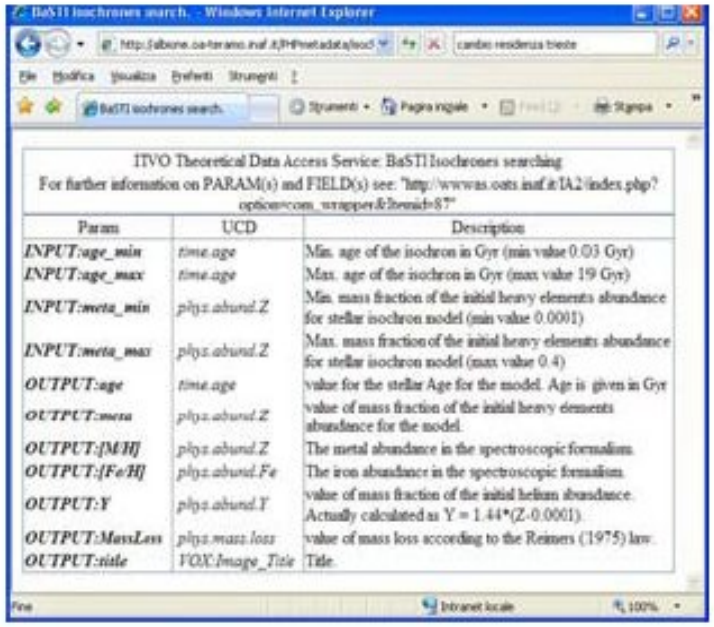
3 step: retrieve the simulated files of interest via http GET;

http://myservice.com/s3.php?id=12

We developed two prototype implementations of S3P for BaSTI: one for isochrones and one for tracks:

http://albione.oa-teramo.inaf.it/PHPmetadata/BaSTIisochron.php?format=metadata

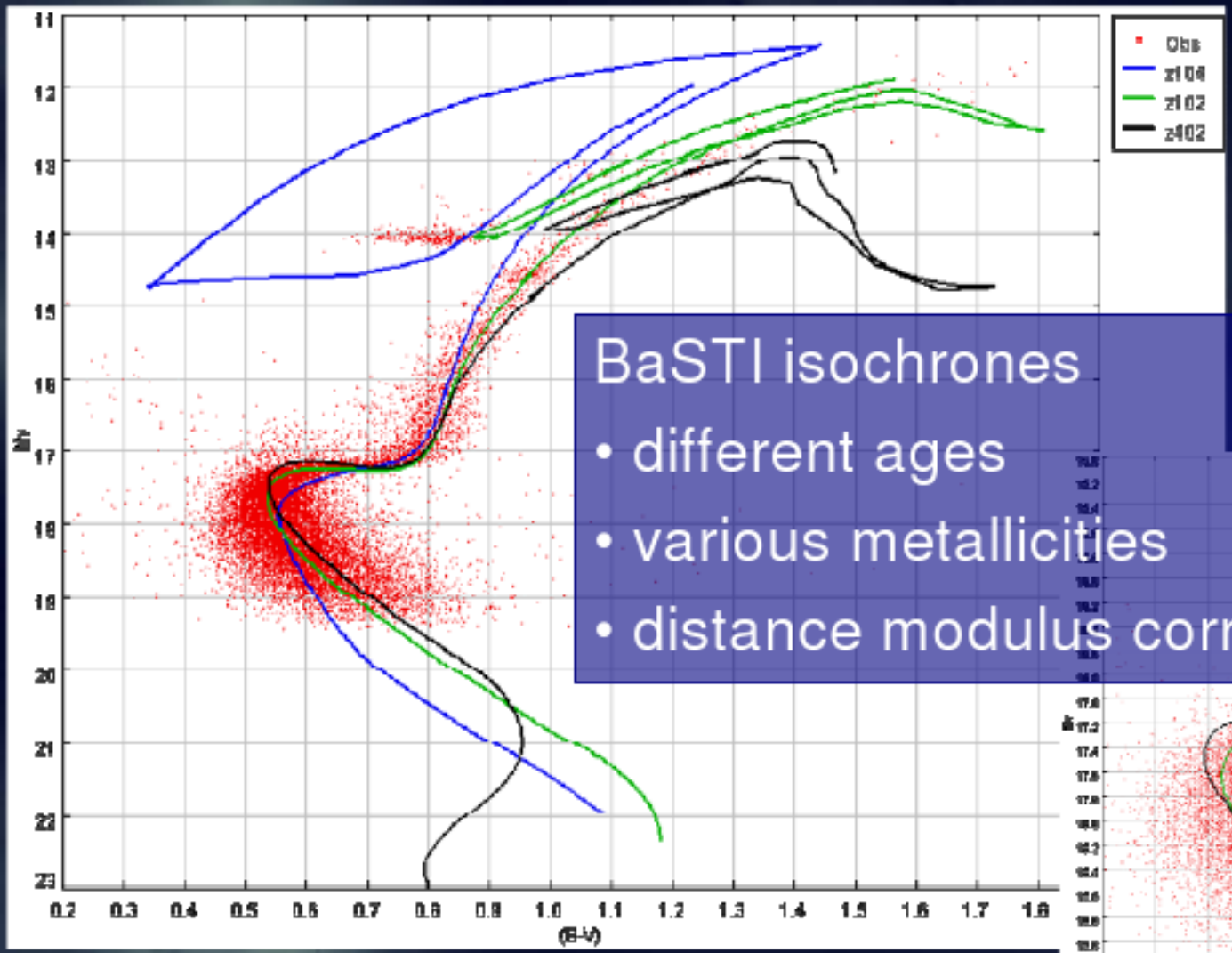
http://albione.oa-teramo.inaf.it/PHPmetadata/BaSTItrack.php?format=metadata



ITVO Theoretical Data Access Service: BaSTI Isochrones searching
For further information on PARAM(s) and FIELD(s) see: "http://www.oats.inaf.it/IA2/index.php?option=com_wrapper&Itemid=87"

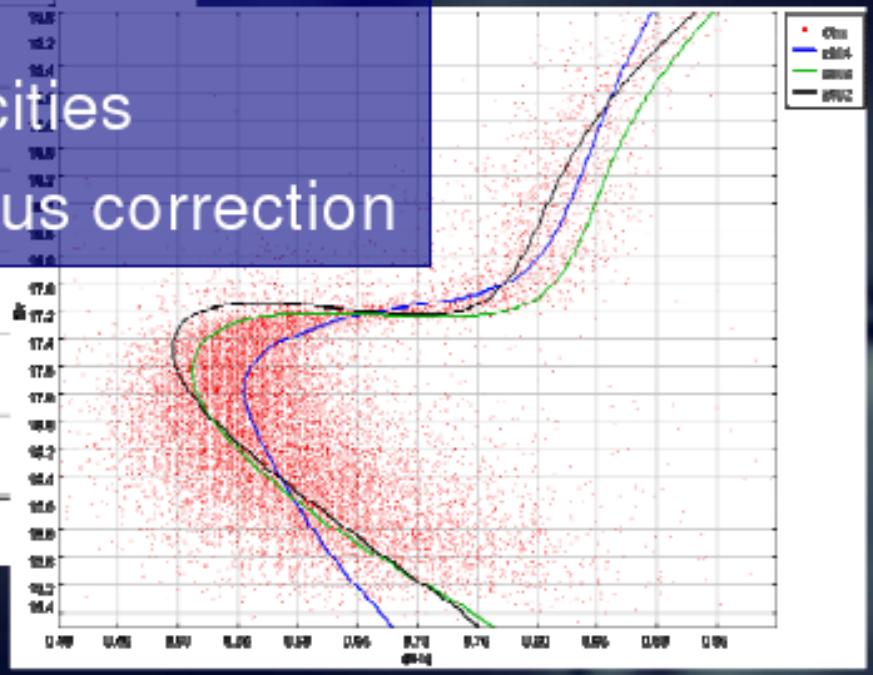
Param	UCD	Description
INPUT:age_min	*time.age*	Min. age of the isochron in Gyr (min value 0.03 Gyr)
INPUT:age_max	*time.age*	Max. age of the isochron in Gyr (max value 19 Gyr)
INPUT:meta_min	*phys.abund.Z*	Min. mass fraction of the initial heavy elements abundance for stellar isochron model (min value 0.0001)
INPUT:meta_max	*phys.abund.Z*	Max. mass fraction of the initial heavy elements abundance for stellar isochron model (max value 0.4)
OUTPUT:age	*time.age*	value for the stellar Age for the model. Age is given in Gyr
OUTPUT:meta	*phys.abund.Z*	value of mass fraction of the initial heavy elements abundance for the model.
OUTPUT:[M/H]	*phys.abund.Z*	The metal abundance in the spectroscopic formalism.
OUTPUT:[Fe/H]	*phys.abund.Fe*	The iron abundance in the spectroscopic formalism.
OUTPUT:Y	*phys.abund.Y*	value of mass fraction of the initial helium abundance. Actually calculated as Y = 1.44*(Z-0.0001).
OUTPUT:MassLoss	*phys.mass.loss*	value of mass loss according to the Reimers (1975) law.
OUTPUT:title	*VOX:Image_Title*	Title.

BaSTI Isochrones



step 1
metallicity

BaSTI isochrones
- different ages
- various metallicities
- distance modulus correction



z = 0.01 (α-enh) ; 0.008 (scaled solar)

Archives, Theory, VO-Science, DataMining, E&O

Theoretical model services

SVO

VOSA: VO Sed Analizer

VO SED Analyzer

Services: VOSA Filters TSAP S3if

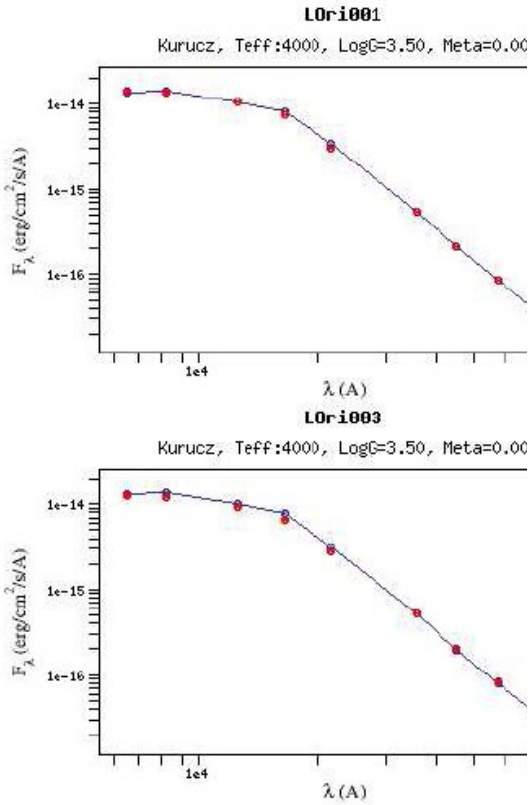
Astronomy & Astrophysics manuscript no. Synth˙VO˙PR1˙ref˙format
August 2, 2008

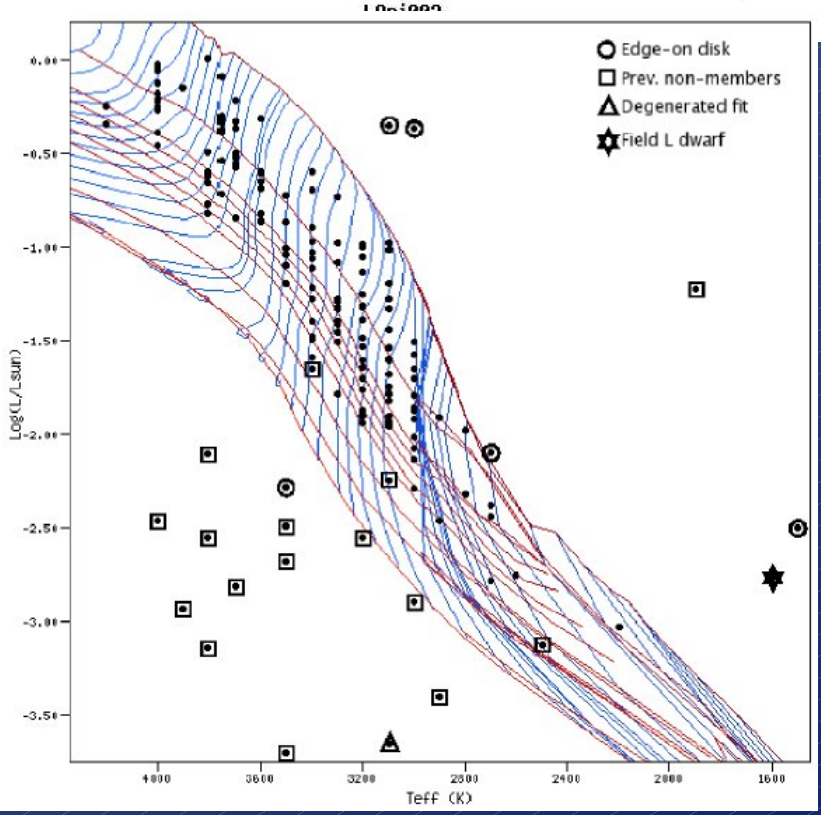
VOSA: Virtual Observatory SED Analyzer.

An application to the Collinder 69 open cluster

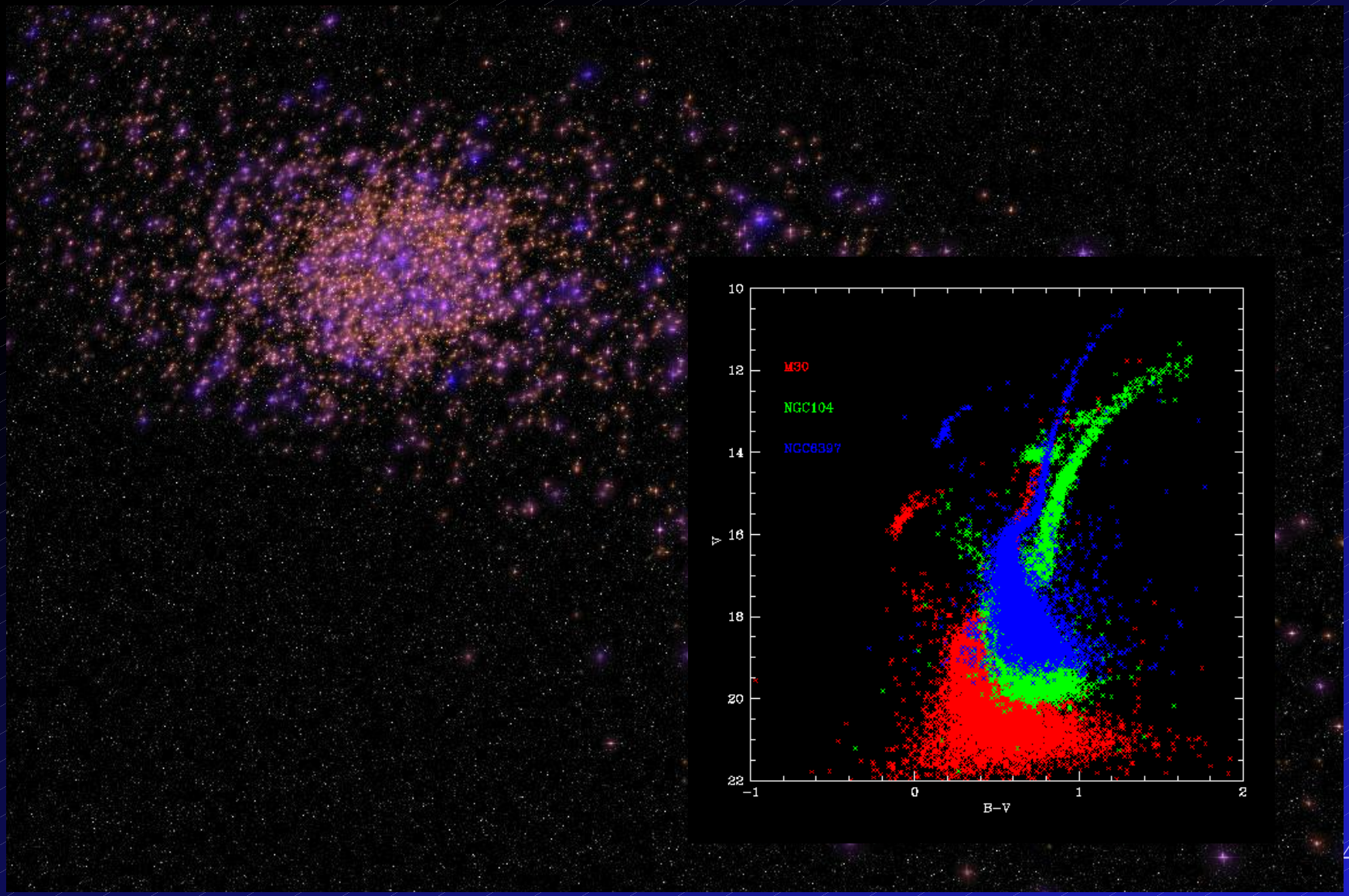
A. Bayo[1,2], C. Rodrigo[1,2], D. Barrado y Navascués[1,2], E. Solano[1,2], R. Gutiérrez[1,2], M.

Object	Model	T_{eff}	LogG	Metallicity	χ^2
LOri001	Kurucz	4000	3.50	0.00	1.71e
LOri002	NextGen	3800	4.5	0	1.44e
LOri003	Kurucz	4000	3.50	0.00	5.05e
LOri004	Kurucz	3750	4.00	0.00	2.82e
LOri005	NextGen	4000	4.0	0	9.20e
LOri006	Kurucz	4000	3.50	0.00	3.36e
LOri007	Kurucz	4000	4.50	0.00	2.49e
LOri008	Kurucz	4000	3.50	0.00	4.43e
LOri009	NextGen	4000	3.5	0	6.22e
LOri010	NextGen	4200	4.0	0	1.07e
LOri011	NextGen	3900	4.5	0	1.20e
LOri012	NextGen	4000	4.5	0	6.56e
LOri013	NextGen	3700	4.5	0	3.50e
LOri014	Kurucz	4000	4.50	0.00	4.70e
LOri015	Kurucz	4000	3.50	0.00	8.73e
LOri016	Kurucz	3750	4.50	0.00	5.63e
LOri017	NextGen	4200	4.0	0	7.56e
LOri018	Kurucz	3750	3.50	0.00	4.31e
LOri019	Kurucz	3750	3.50	0.00	1.90e
LOri020	NextGen	3800	4.5	0	2.90e
LOri021	Kurucz	4000	3.50	0.00	3.08e
LOri022	Kurucz	3750	4.00	0.00	1.78e
LOri023	NextGen	4000	4.5	0	2.35e
LOri024	Kurucz	3750	3.50	0.00	2.22e
LOri025	NextGen	3700	4.5	0	1.37e
LOri026	NextGen	3700	4.5	0	4.81e
LOri027	NextGen	4000	4.5	0	2.35e
LOri028	Kurucz	3750	4.00	0.00	1.26e
LOri029	NextGen	3100	4.5	0	7.28e
LOri030	NextGen	3700	4.5	0	2.15e
LOri031	NextGen	3800	4.5	0	2.77e
LOri032	NextGen	3700	4.5	0	1.70e
LOri033	NextGen	3700	4.5	0	7.12e
LOri034	NextGen	3000	4.0	0	1.77e
LOri035	NextGen	3700	4.5	0	1.61e

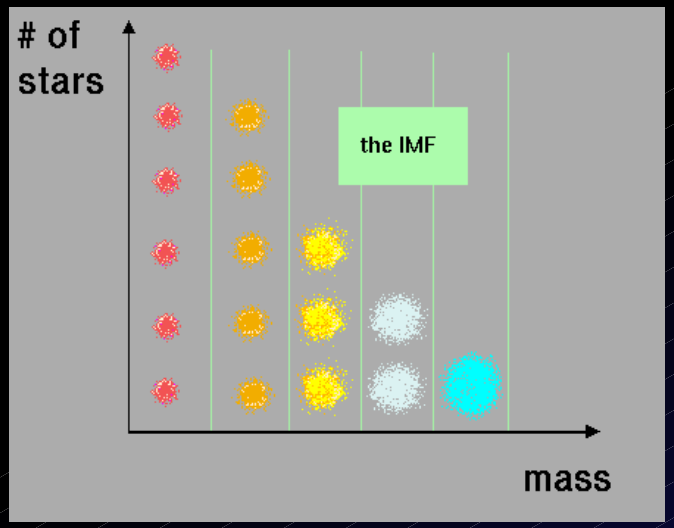


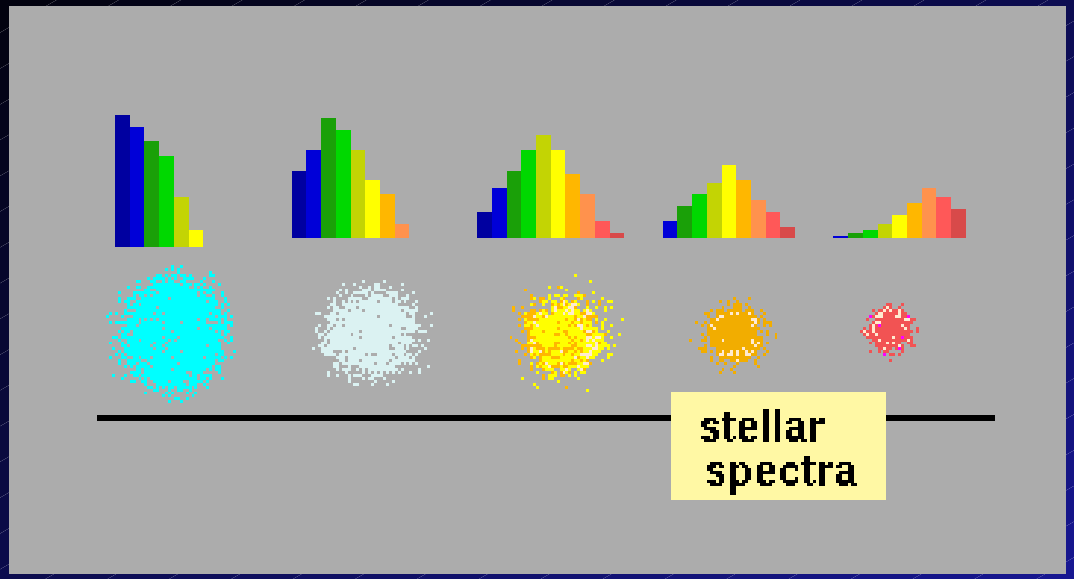


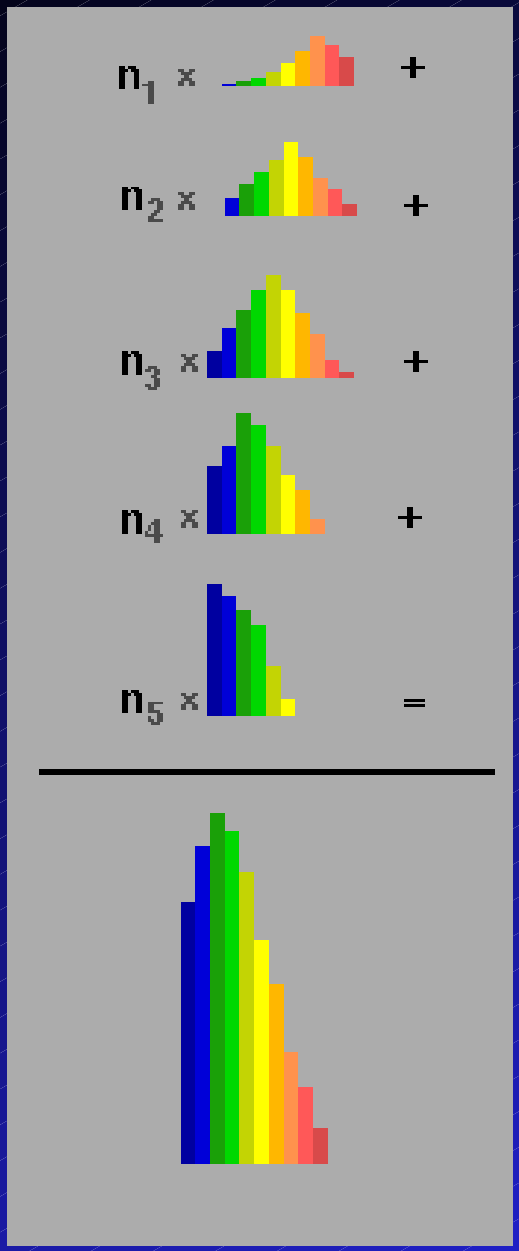
N Body Simulations of Globular Cluster Evolution



Stellar populations are modeled with synthesis models



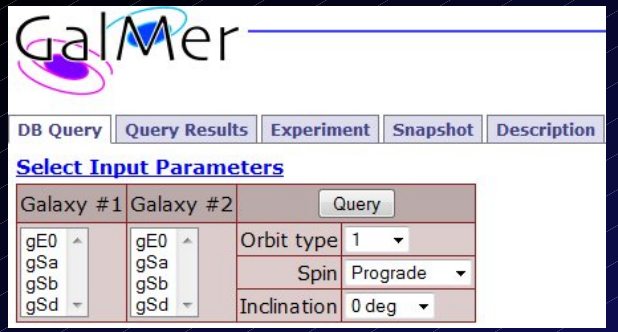


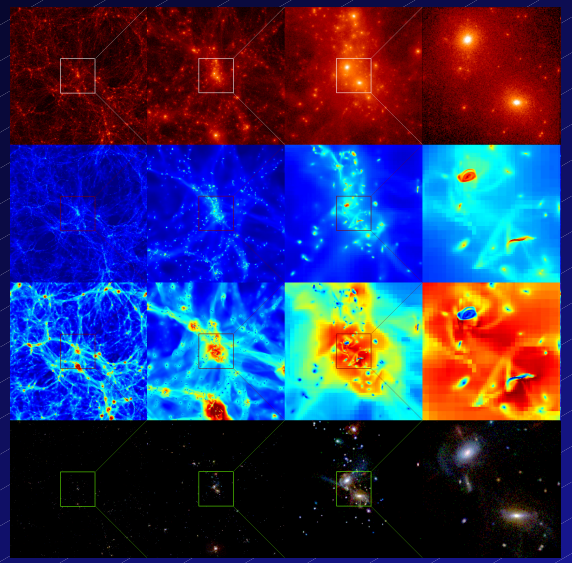


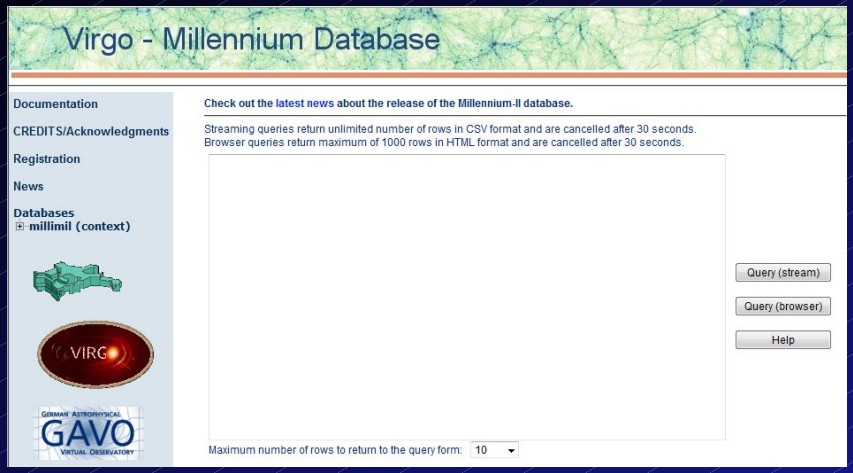
(Available as *Theoretical Simple Access Protocol* server :http://ov.inaoep.mx)

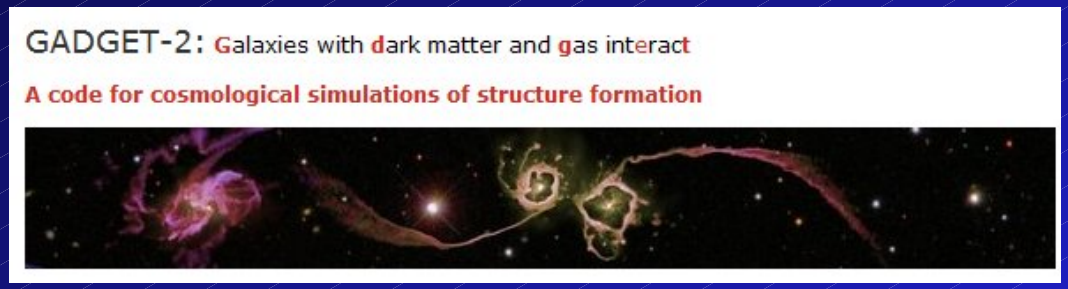
Using SimDB/SimDAP

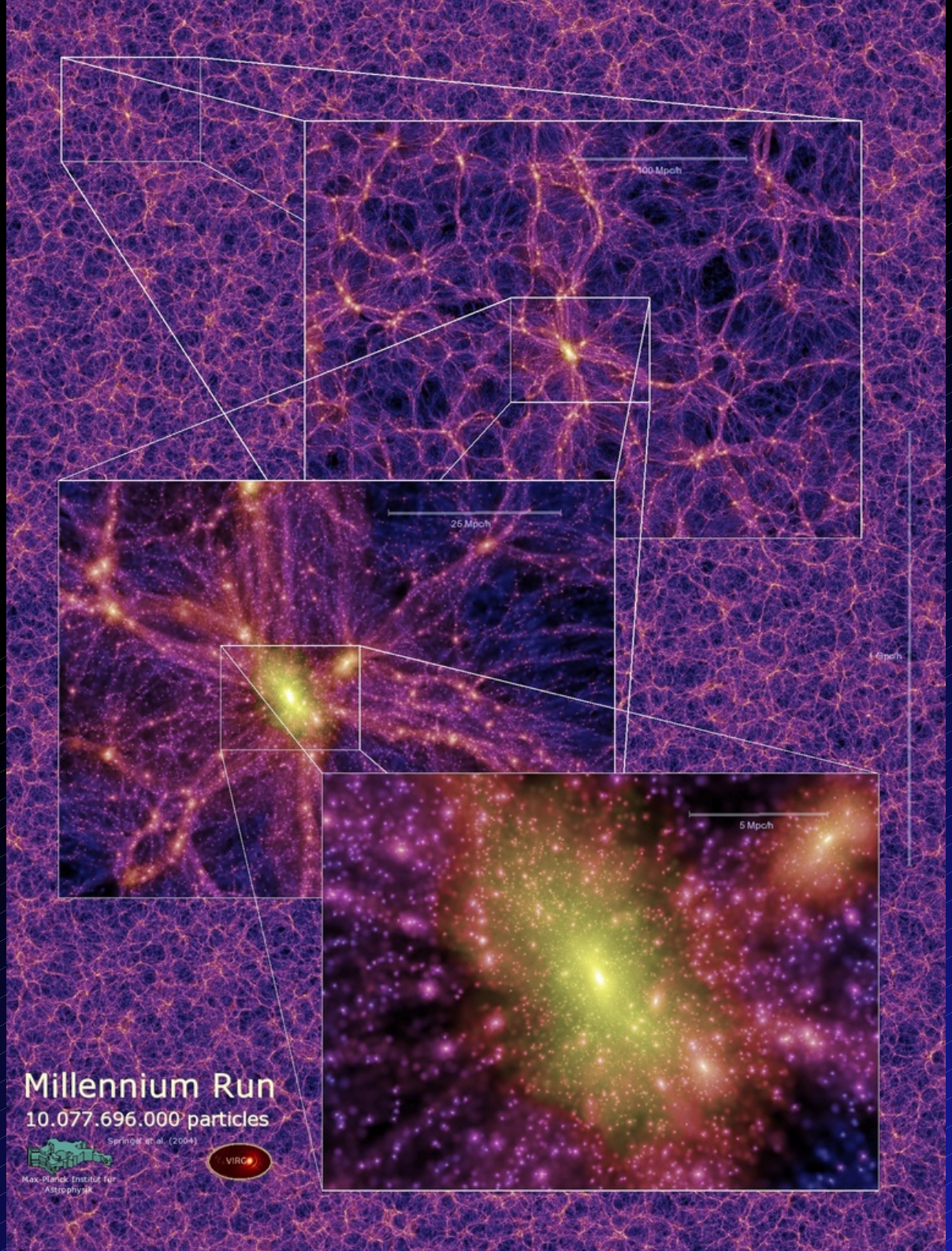
- Cosmological simulations
  - Prototypes for GalMER, Horizon
- PDR simulations
  - test implementation of Meudon PDR code
- Isochrones/evolutionary tracks
  - BaTSI
- Visualization tools
  - VisIVO







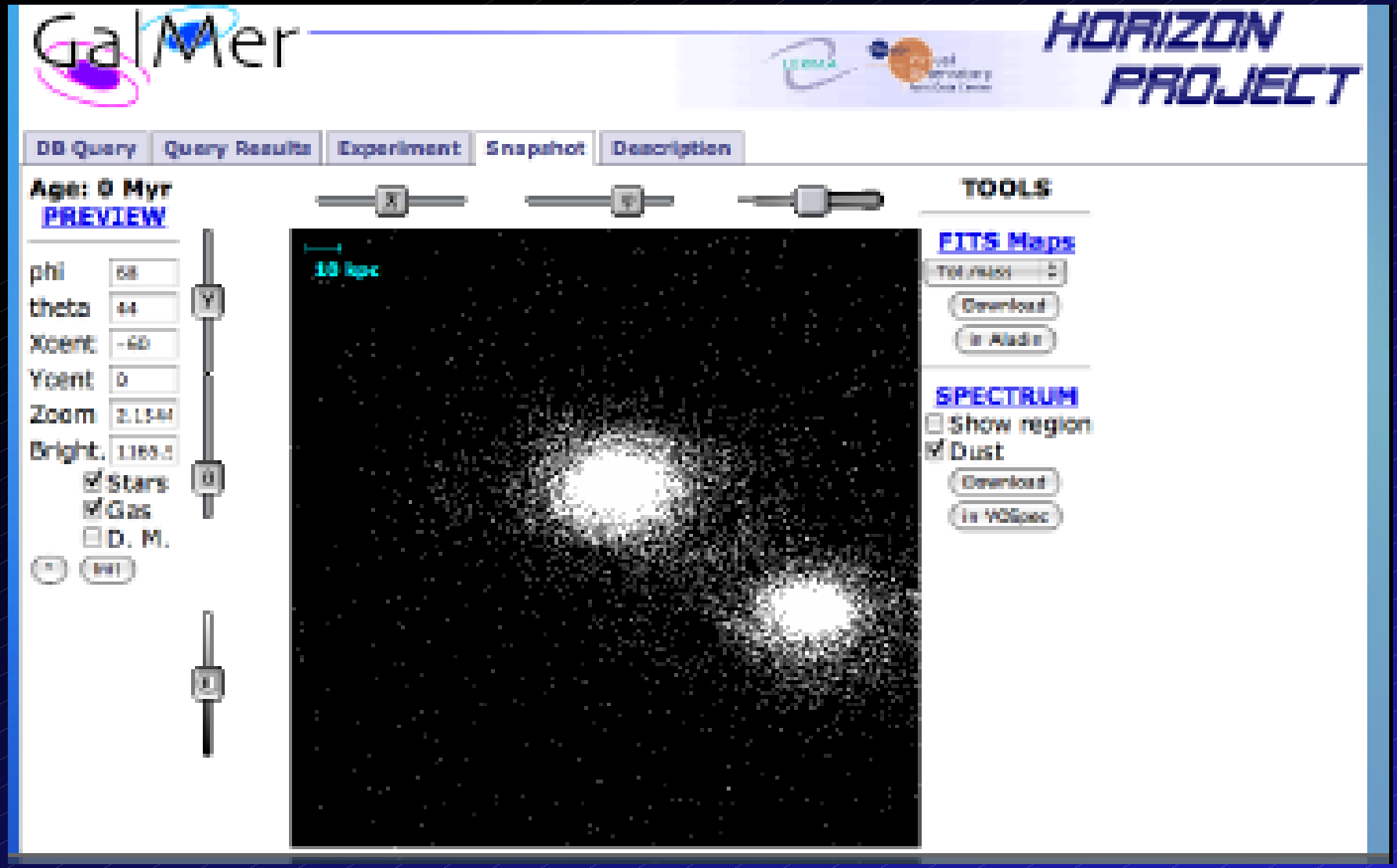




Millenium Run

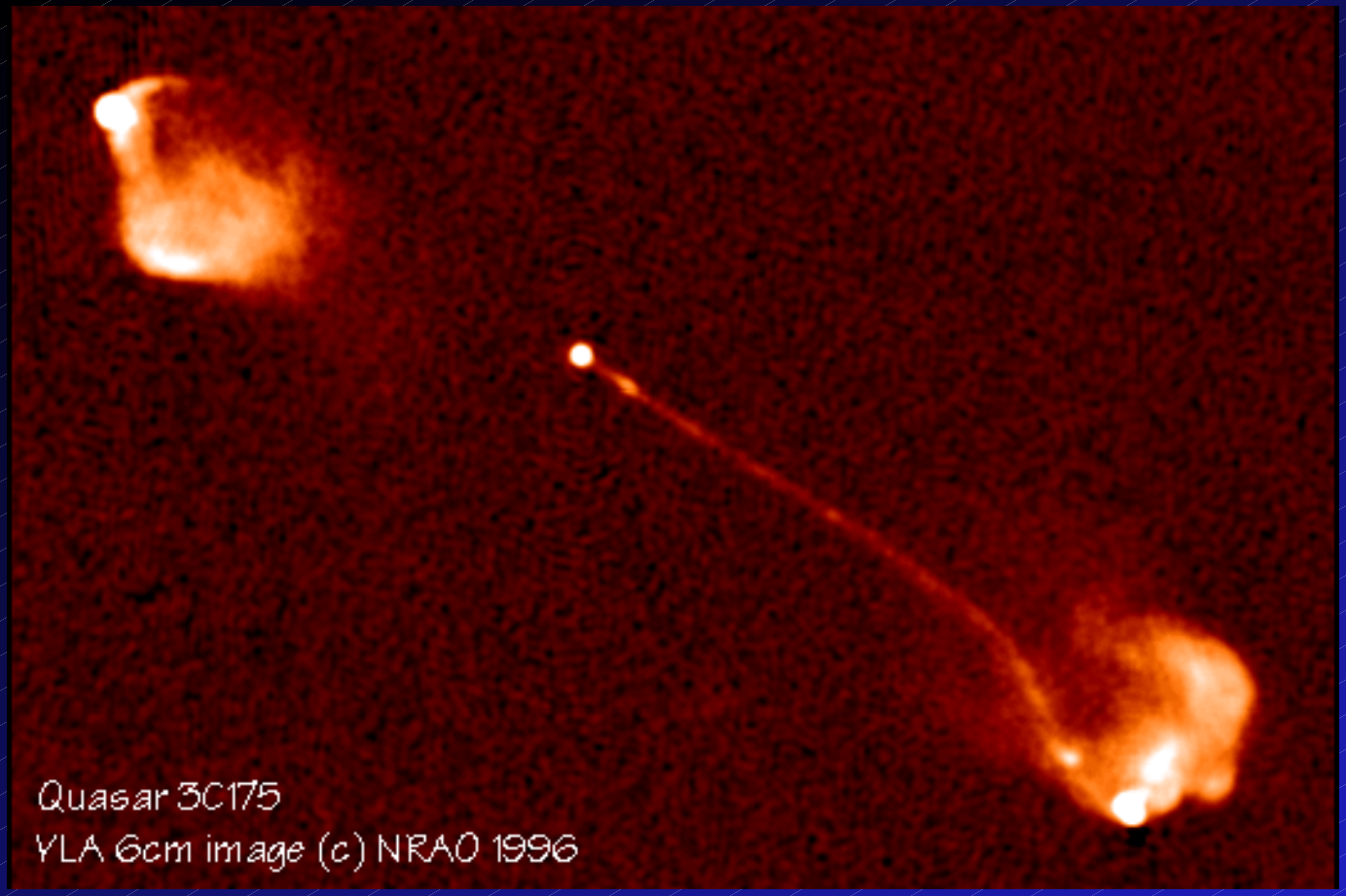
- 10^10 particles
- Several Gpc to
- 10 kpc
- Cube 2 billion ly
- One month MPSSC
- 25 TB
- Evolution of 20 mil galaxies
- Evolution merger tree

Galaxy Merger Service - Client

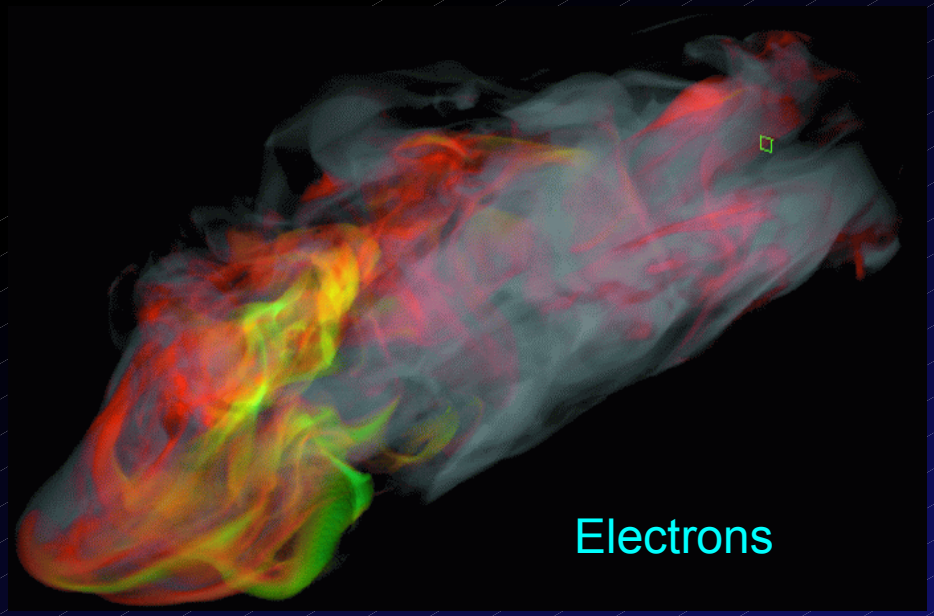


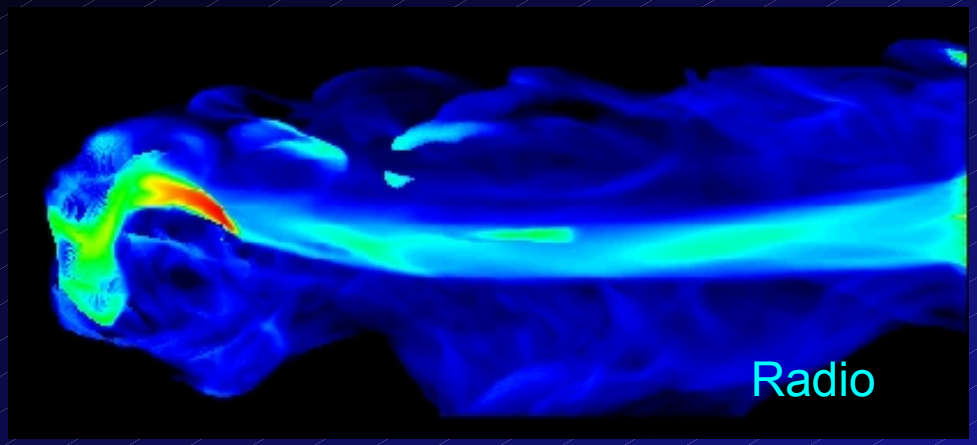
Collimated Outflows from AGN

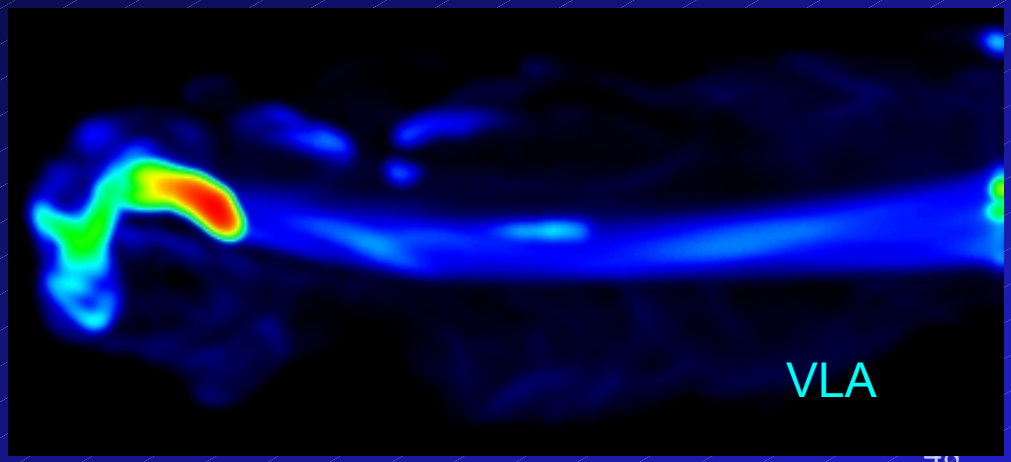
- 3C 175



MHD Simulations of Collimated Outflows from AGN – Virtual Telescope Observations

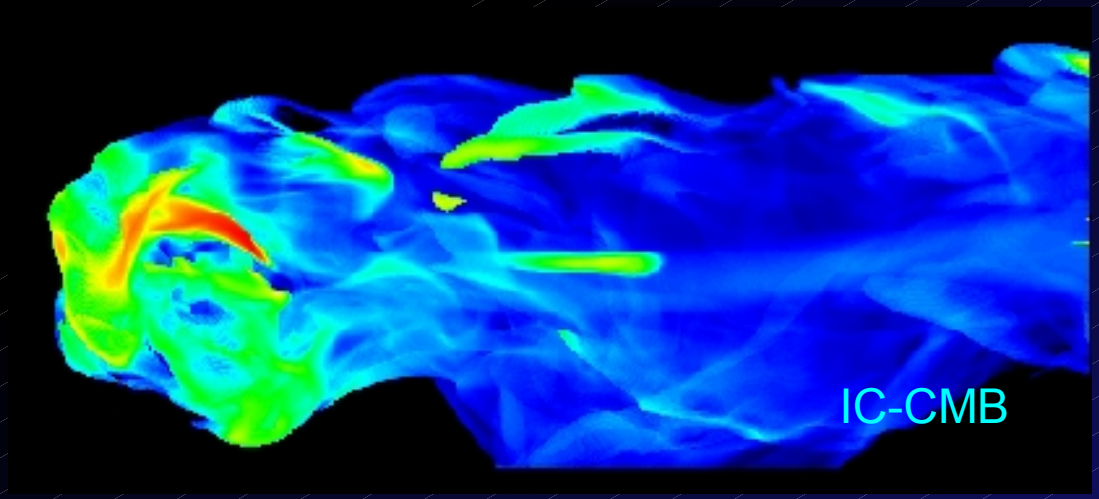


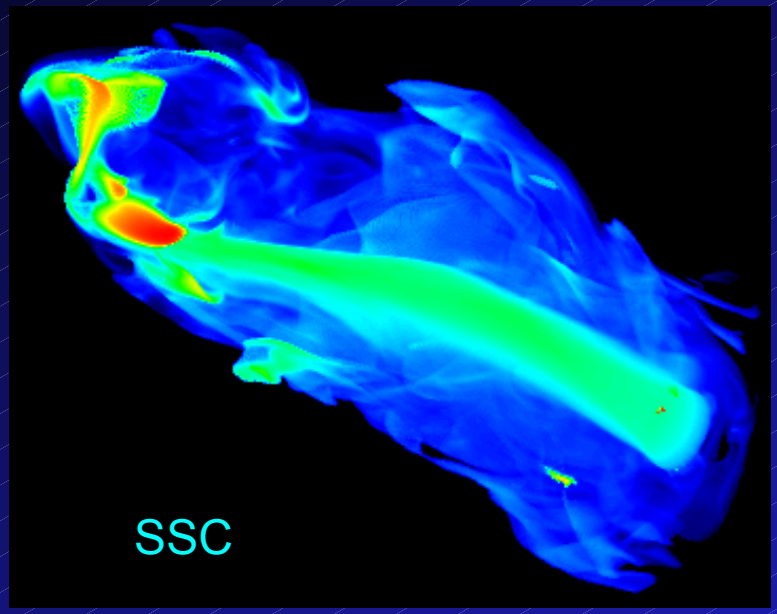


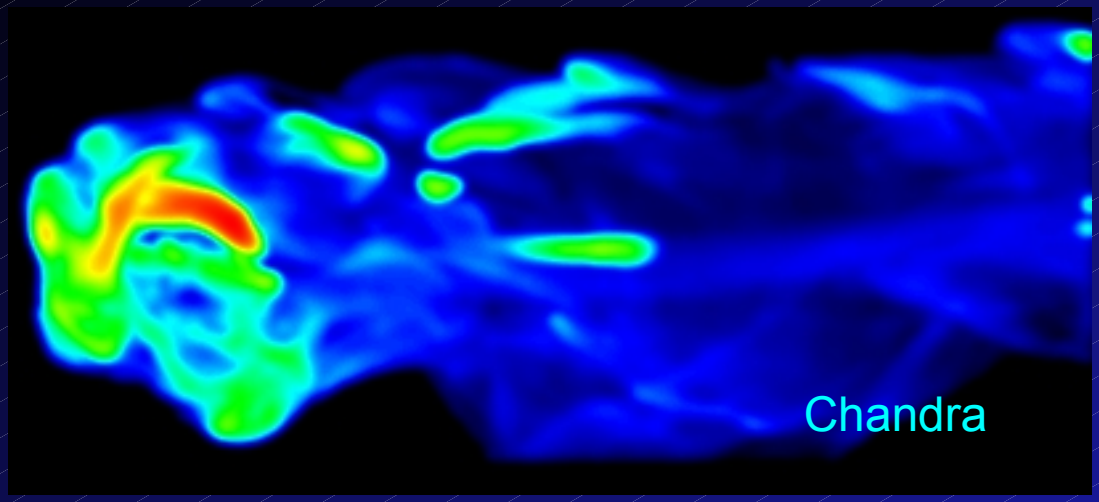


Compare with Radio Archives

MHD Simulations of Collimated Outflows from AGN – Virtual Telescope Observations



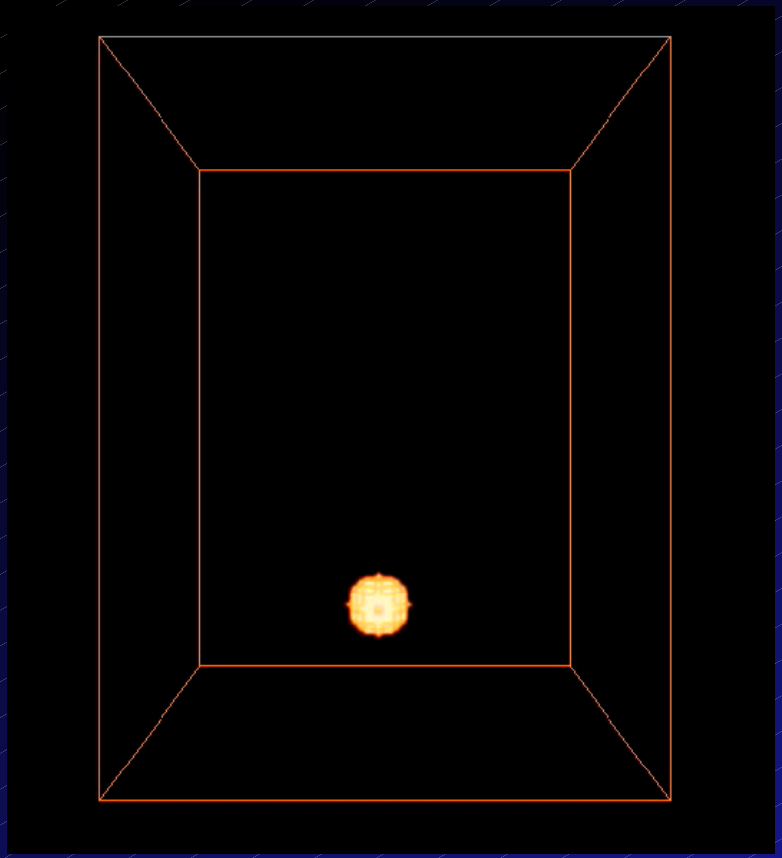


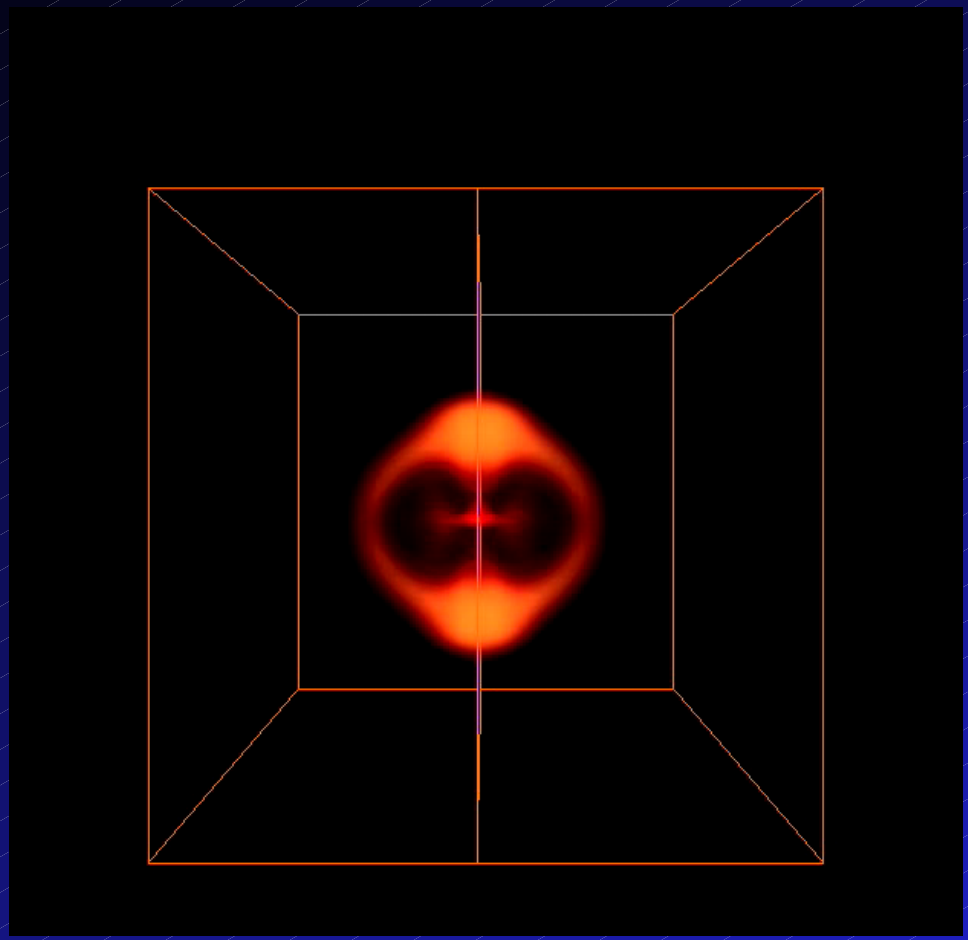


Compare with Chandra Archives

Three Dimensional MHD Calculations

- β = 3000

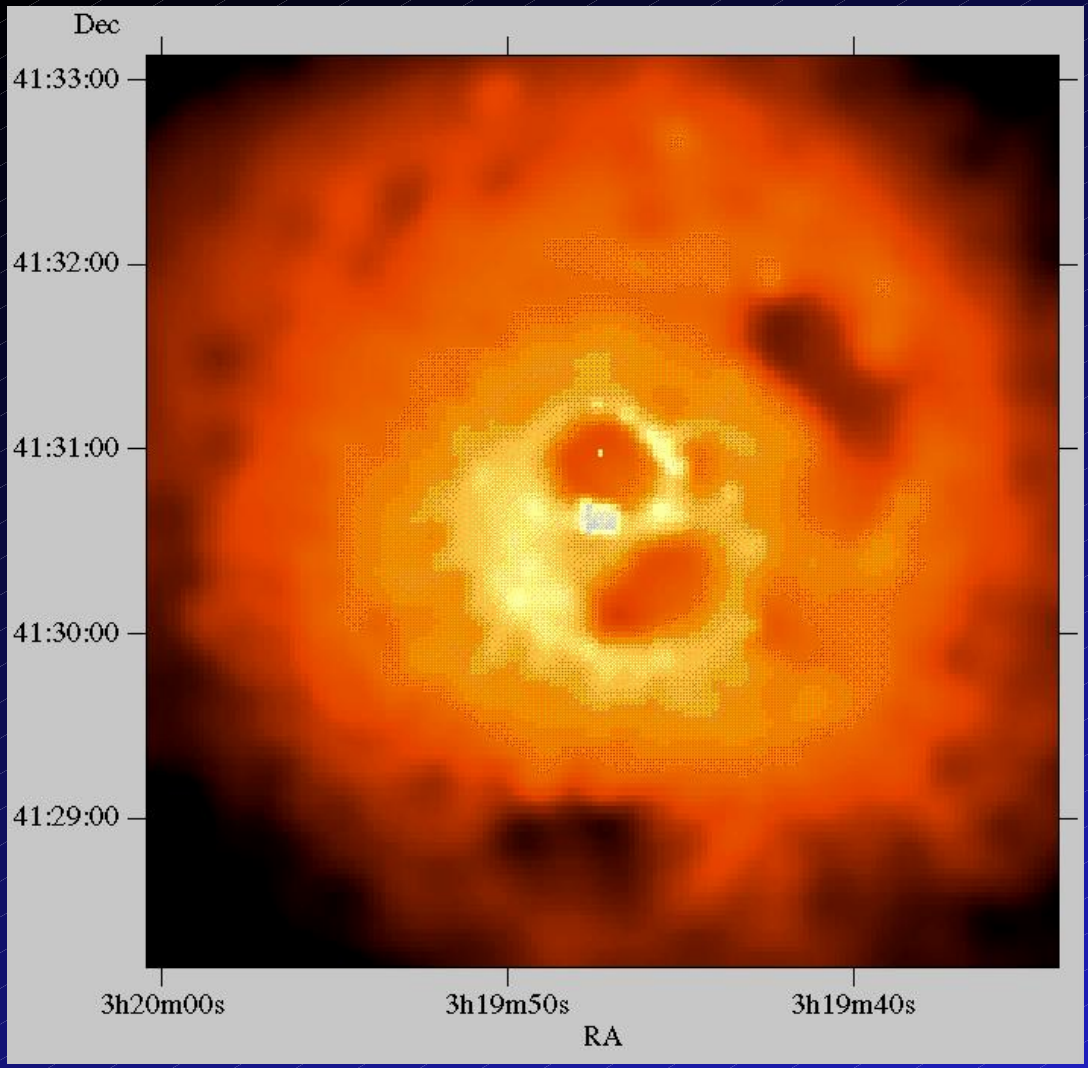




Relic Radio Bubbles in Galaxy Clusters

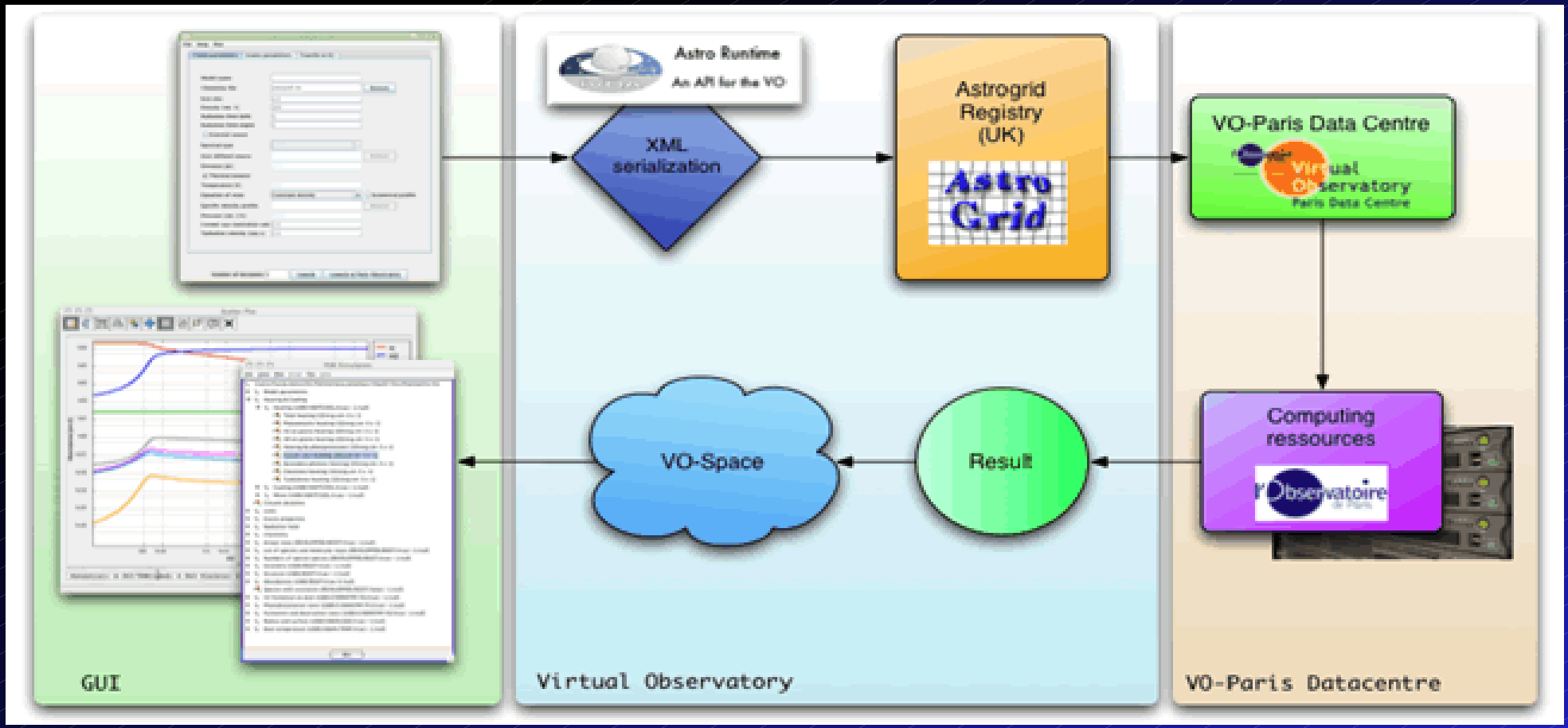
- N1275

Compare with Chandra Archives



Fabian et al. 2000

PDR VO-infrasctructure



PDR database and clients

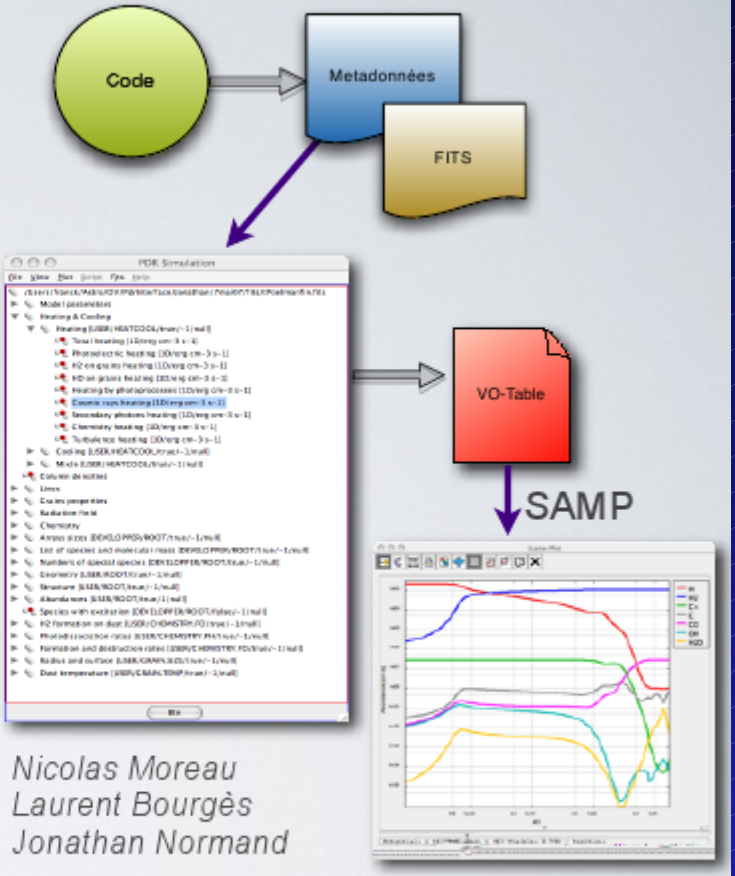
PDR Database

Output Files

Code produces
- raw data : FITS File
- XML / VO-TABLE : meta-data
  (name, description, units, UCD, ...)

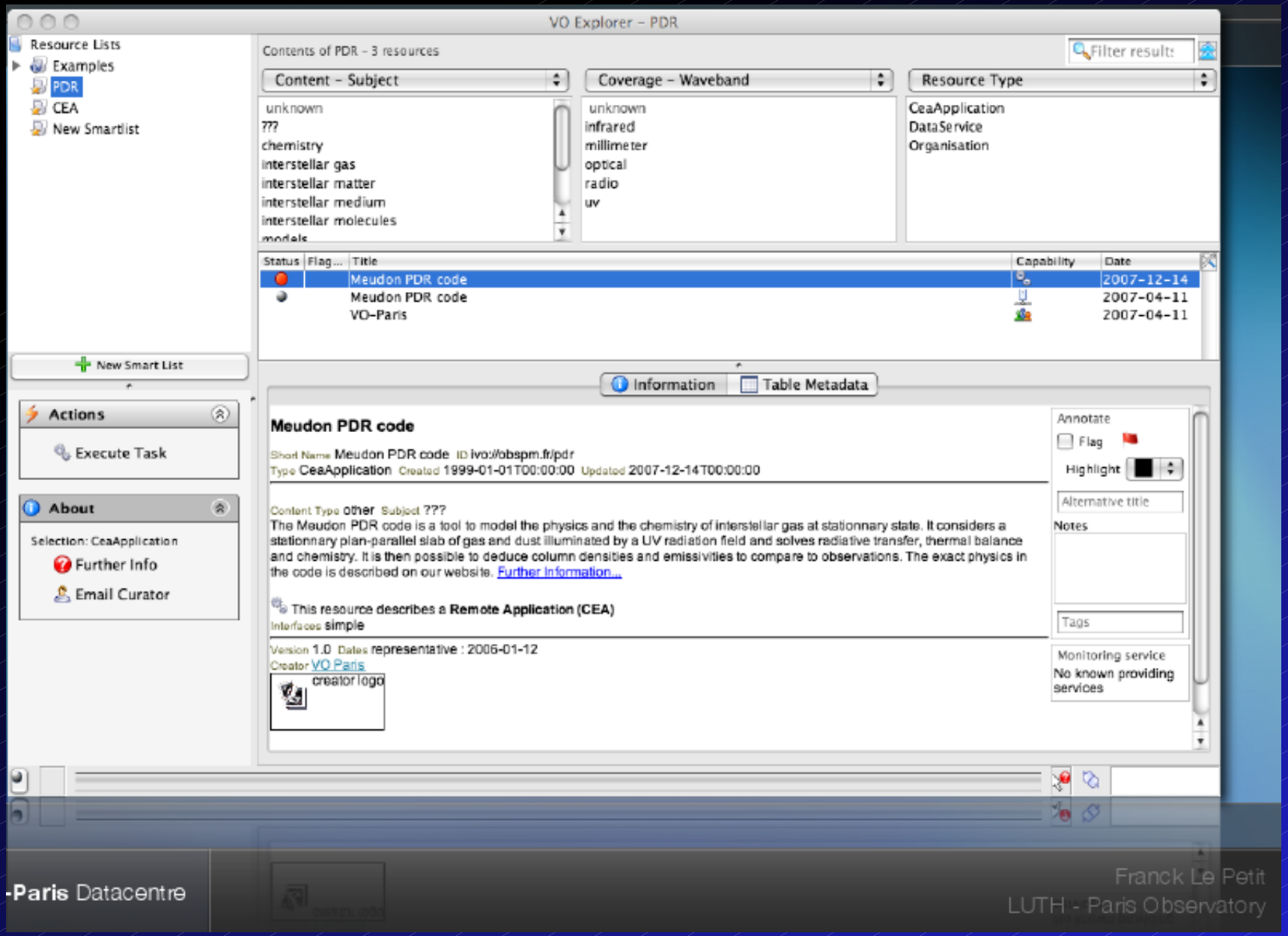
Provide all quantities computed by the code
- observables
- theoretical quantities

PDR Analyser
- browse the computed quantities
- extraction (ASCII, VO-Table)
- SAMP
- Download data from VO-Space
- Scriptable



Nicolas Moreau
Laurent Bourgès
Jonathan Normand

PDR code via VODesktop



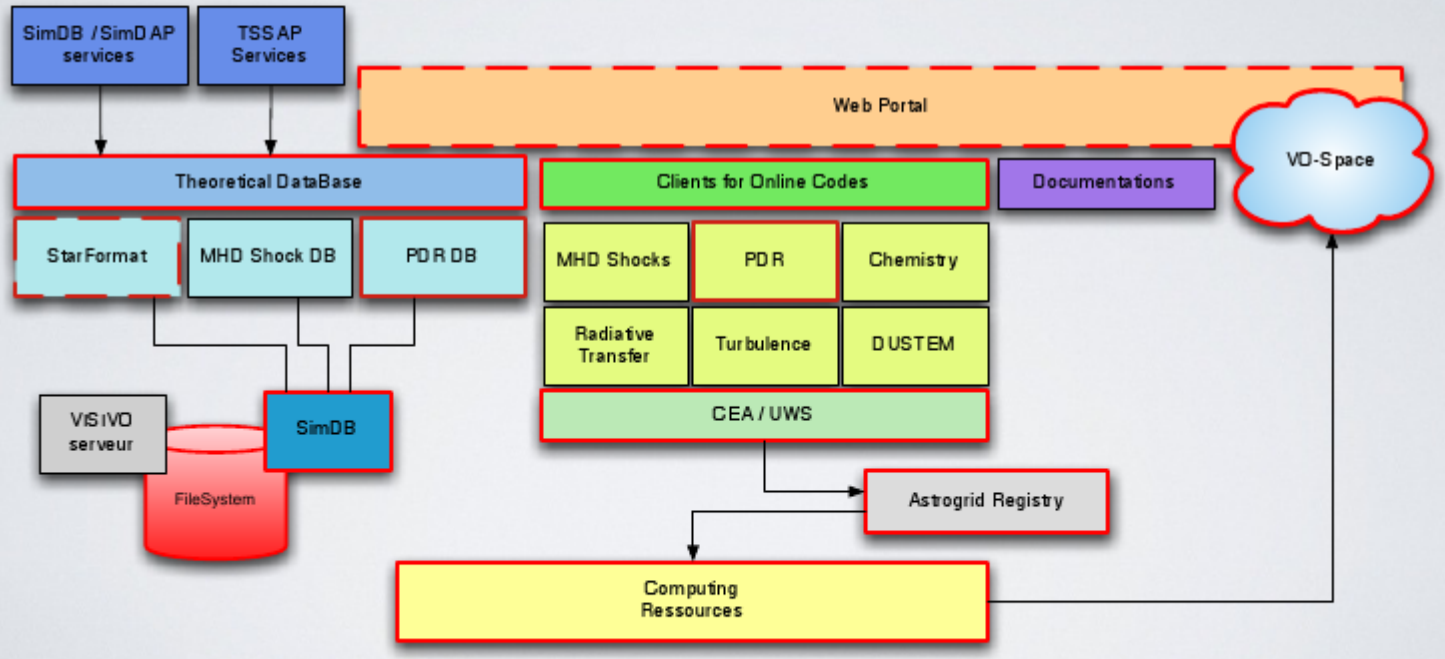
ISM platform

Interstellar Medium Platform

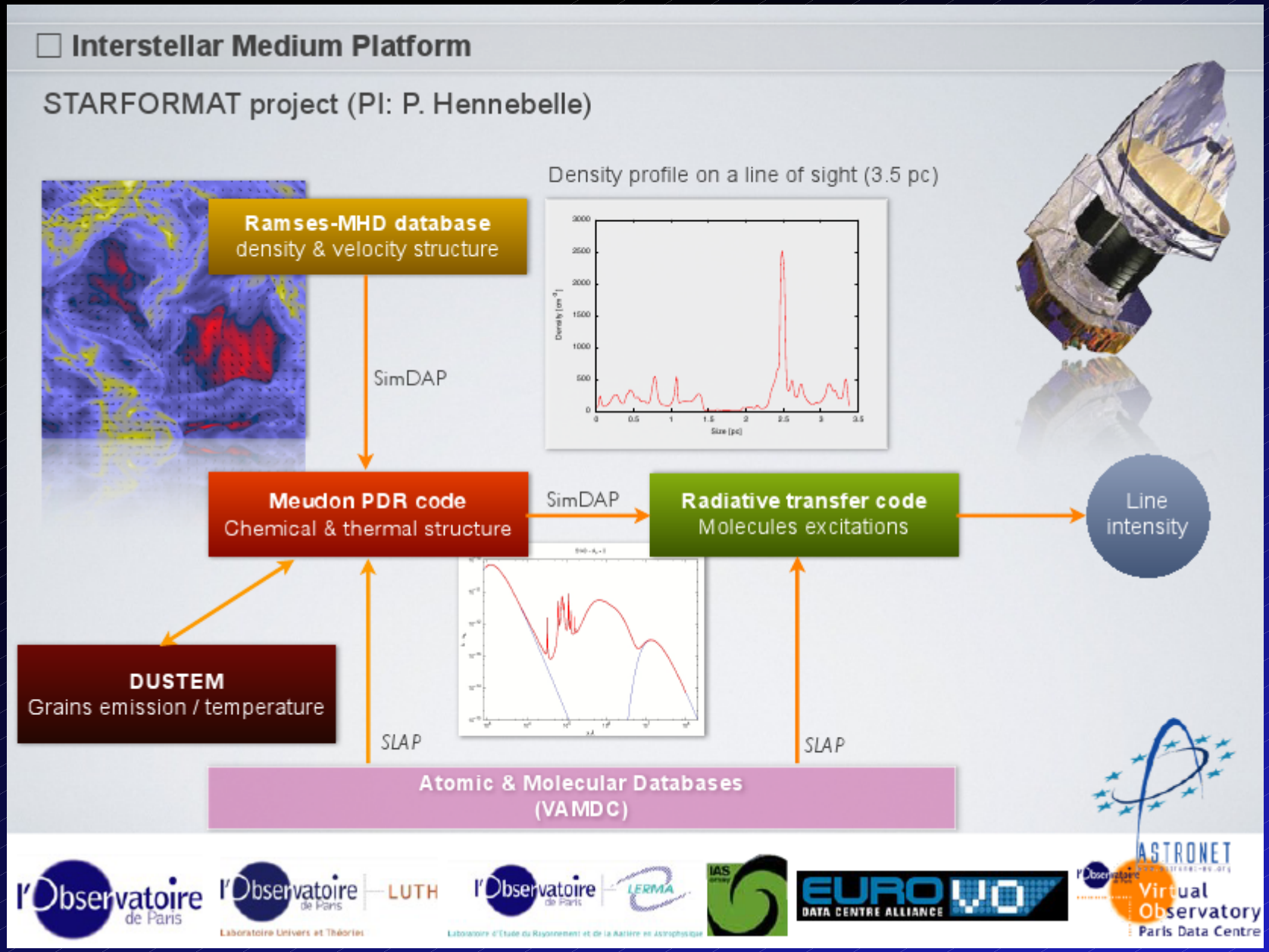
Bring together expertise in modeling / simulation of the ISM

Provide theoretical services about ISM

Codes - Databases - Tools & services



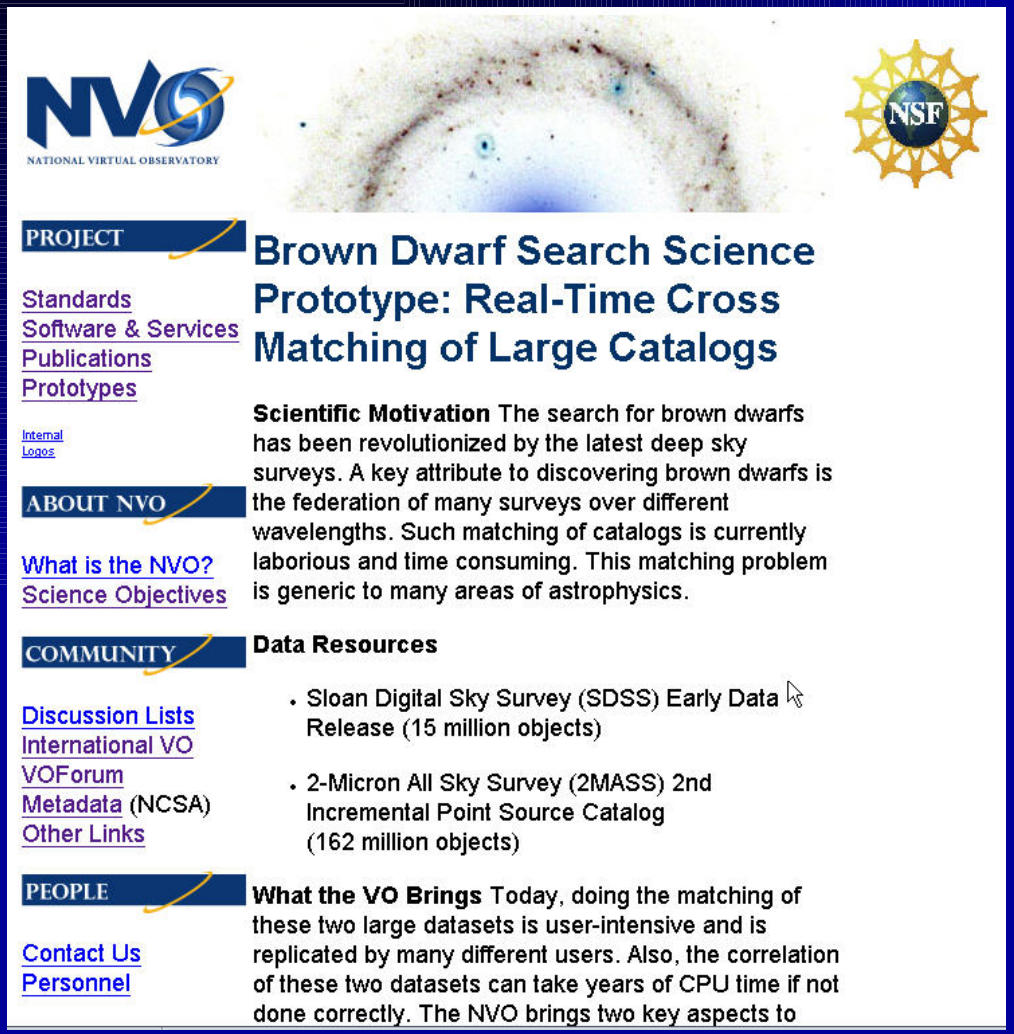
Complex join of TVO bricks



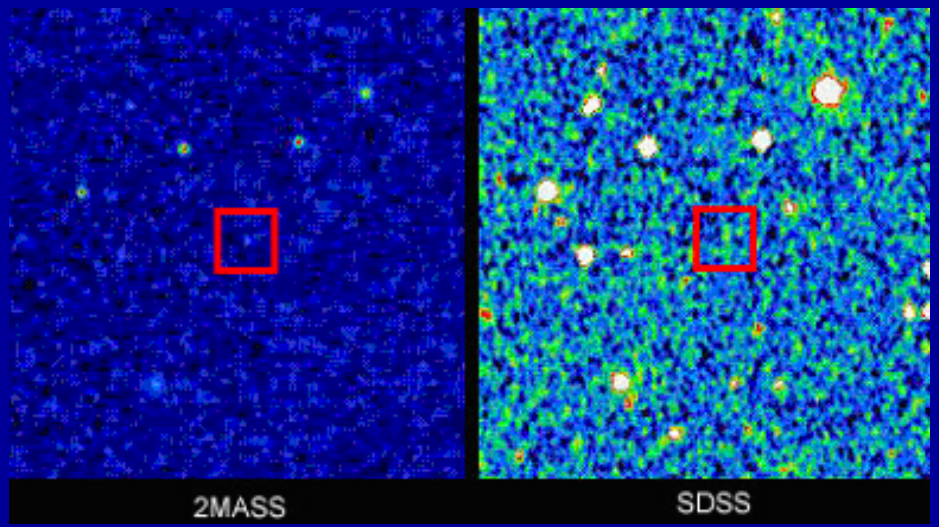
Examples of VO Science

- 31 (9) new obscured type 2 QSO (Padovani 2004)
- Brown dwarfs (about 20 candidates)
- Brightest (WD?) Albus-1 (Cabalero et al. 2008)
- Widest CPM binaries
- AGB to PNe - 100 new (200) with VO
- SED (Spectrum Energy Distribution)
- Bolometric magnitude
- VOEvent – robotic telescopes (GRB, transits,)
- Outreach , Education (MS WWT, GoogleSky)

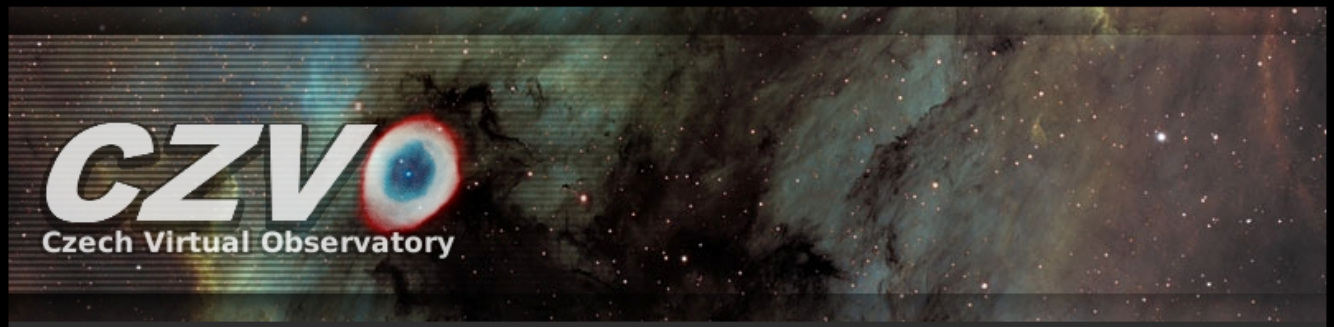
BDs discovered using VO



- **Filtering criteria:** z & J-only detections with *z- J > 2.75*
- *SDSS: 15M obj.*
- *2MASS: 160M obj.*
- *300000 objects in common.*



✓ *However, systematic searches using a VO methodology have not been performed so far.*



Navigation

- About CZVO
- Observatories
- Projects
- Spectra Archive
- Data Resources
- Links
- Publications

Home

Links

There is some links:

Virtual Observatory United Kingdom

Astrogrid

Australian Virtual Observatory/

Chinese Virtual Observatory

Canadian Virtual Observatory

European Virtual Observatory

German Astrophysical Virtual Observatory

Hungarian Virtual Observatory

Japanese Virtual Observatory

Korean Virtual Observatory

National Virtual Observatory, United States

Observatoire Virtuel France

Russian Virtual Observatory

Spanish Virtual Observatory

Italian Virtual Observatory

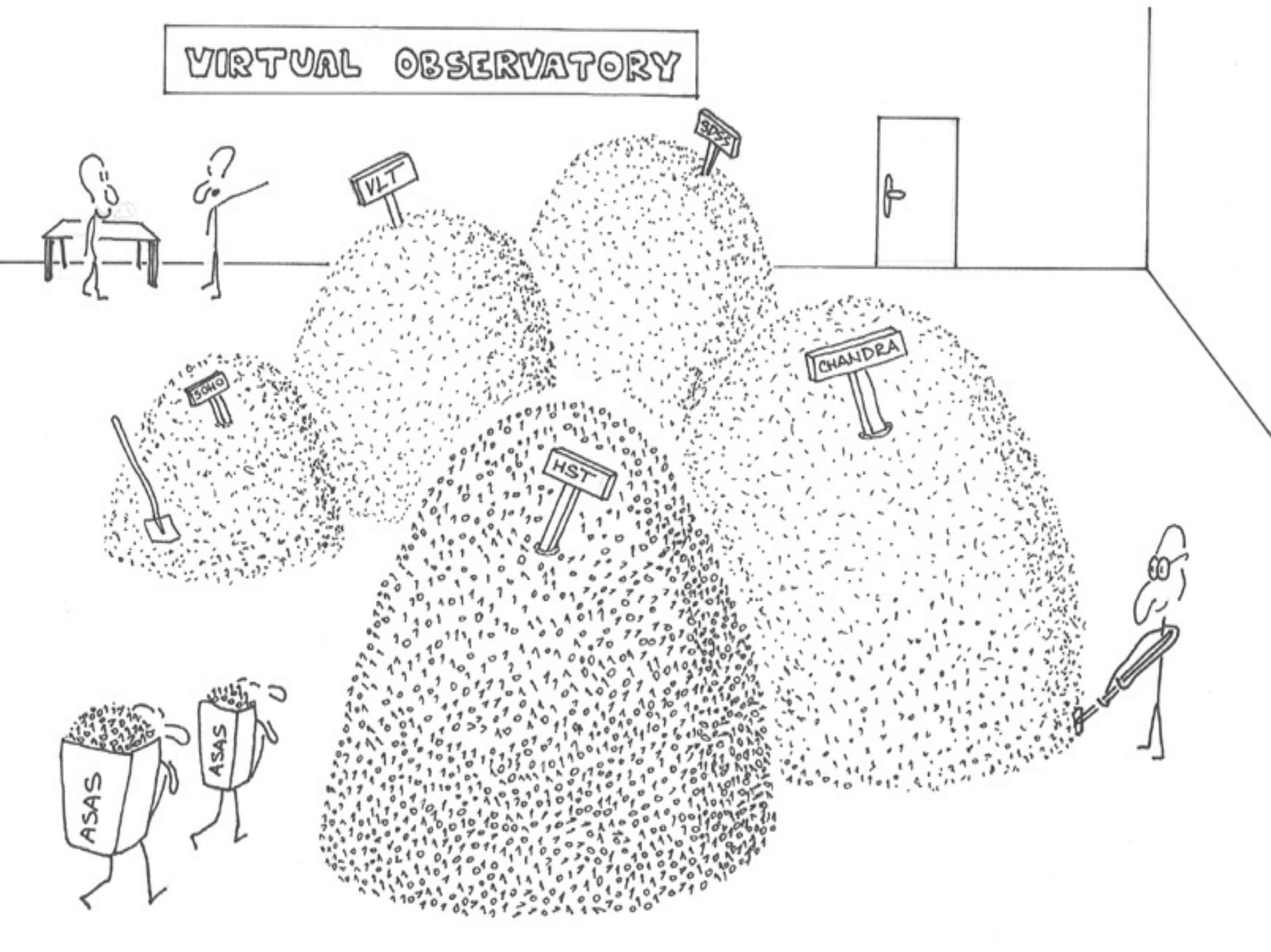
Virtual Observatory India

CZVO

- VO-KOREL (web services)
  - parallel run of many jobs – more users
  - using VO Universal Worker Server (CEA)
  - job control, queuing, jobs results polling
  - will be integrated in VODesktop
- 1D spectra cutout server (HEROS)
  - SSA access to 1D spectra + cutout of regions (lines)
  - need normalization, rebinning , convolution (resolution) on server
- Data mining – AstroNeural + Clustering

Killer spectral applications for VO

- Use VO to find all stars with emission in given line (EW<0) – find the time when it was in em.
- Use VO to get 1000 spectra of the given object cut out regions arround given lines, plot the lines, make a gray dynamic spectrum folded in time
- The same – search period, fold by period
- Get the unknown line ID of piece of spectra from SLAP overploted over SSA data
- Create Light and RV curve for given period
- Fit the grid of models (Teff, log g) to the observed spectrum – for many stars


VIRTUAL OBSERVATORY
VLT
SDSS
SOHO
HST
CHANDRA
ASAS
ASAS