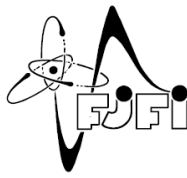




Seminář ATLAS – FZU AV ČR



28/3/2008

Uvádění pixelového detektoru experimentu ATLAS do provozu

Pavel Jež

FZU AVČR, v.v.i.

FJFI ČVUT



Pixelový detektor - status

- Hlavní rozcestník: <https://twiki.cern.ch/twiki/bin/view/Atlas/PixelWiki>
- Narozdíl od SCT a TRT není pixelový detektor stále zapojen (řada technických problémů – více info na pixel meetingu, každý čtvrtek od 16h v CERN)
 - Díky tomu se pixelový detektor neúčastní „milestone“ (MX) testů, kdy je (téměř) celý detektor zapojen
 - k dispozici data pouze z ROD simulací

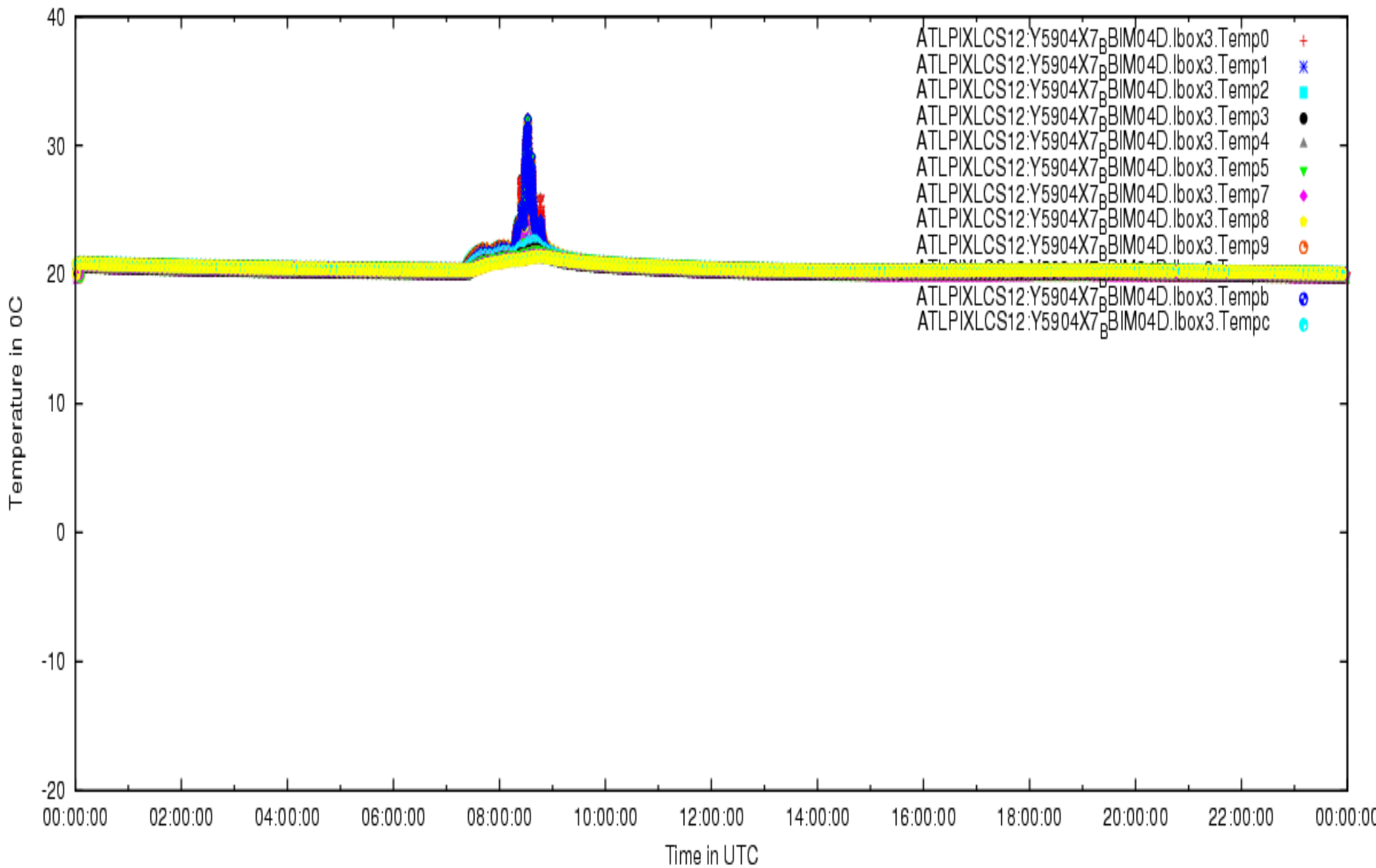
Pixelový detektor – status II

- Testuje a vyvíjí se monitorovací software
 - Není zřejmé, které veličiny je třeba sledovat
- Online monitoring – kontroluje především hardware detektoru, úzká souvislost s DCS (detector control system)
 - Histogramy (dočasně) ukládány v http://www.slac.stanford.edu/~behera/Pixel/DCS_Mon/
- Offline monitoring – slouží ke zjištění kvality dat proudících z detektoru
 - Samostatný balíček v Atheně

InnerDetector/InDetMonitoring/PixelMonitoring

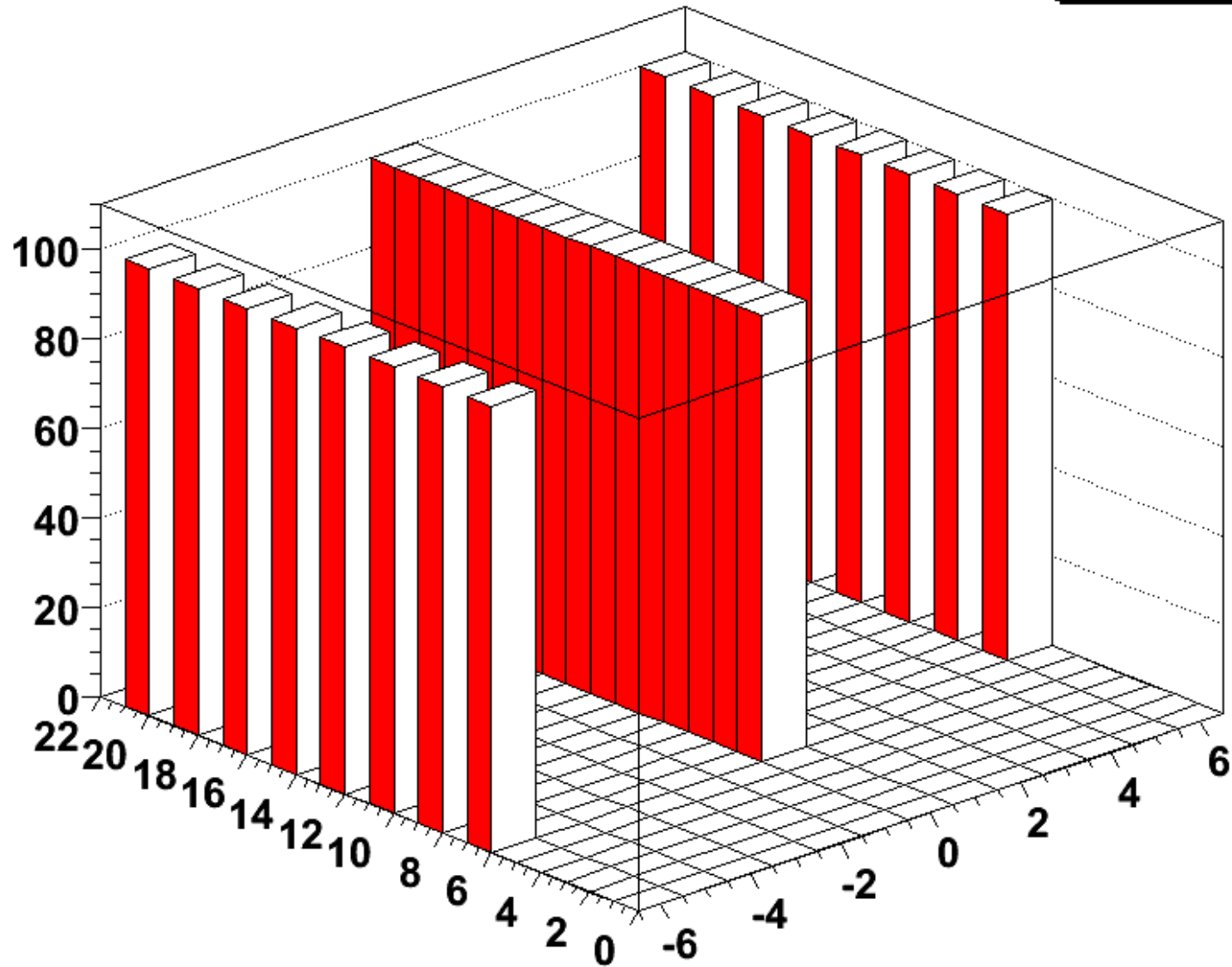
Teplota pixelového modulu

From 18-03-2008 00:01 to 18-03-2008 23:59 UTC



Barrel layer 0 occupancy

Entries 3200



Occupancy = number of hits during run

Pixel Offline Software

- Součást SW rámce **Athena**
 - soubor společných knihoven a nástrojů pro offline software na ATLAS (např. HEPVector...)
 - zároveň je to soubor veškerých algoritmů pro MC simulaci eventů i detektoru, rekonstrukci a fyzikální analýzu
 - balíčková struktura
 - všechny algoritmy v C++
- CVS:** <http://atlas-sw.cern.ch/cgi-bin/viewcvs-atlas.cgi/offline/>
- LXR:** <http://alxr.usatlas.bnl.gov/> (s vyhledáváním)

ATHENA

- SW balíčky spravovány pomocí CMT (configuration management tool)
- Uživatel (vývojář) zvolí, s jakou verzí SW chce pracovat a připraví prostředí
<https://twiki.cern.ch/twiki/bin/view/Atlas/ReleaseRecipes>
- stáhne zdrojový kód požadovaného balíčku z centrálního úložiště

```
cmt co <název balíčku>
```

ATHENA

- všechny balíčky mají stejnou adresářovou strukturu:
 - *cmt* : slouží k propojení s ostatními balíčky (soubor `requirements`)
 - *<název balíčku>* : obsahuje hlavičkové soubory
 - *src* : obsahuje algoritmy v C++
 - *run* : zde se spouští jednotlivé joby
 - *share* : obsahuje příklady řídicích skriptů (jobOptions), jimiž se volají jednotlivé algoritmy a předávají parametry

ATHENA – kompilace a spuštění

- typické kroky po stažení balíčku:

```
cd cmt
```

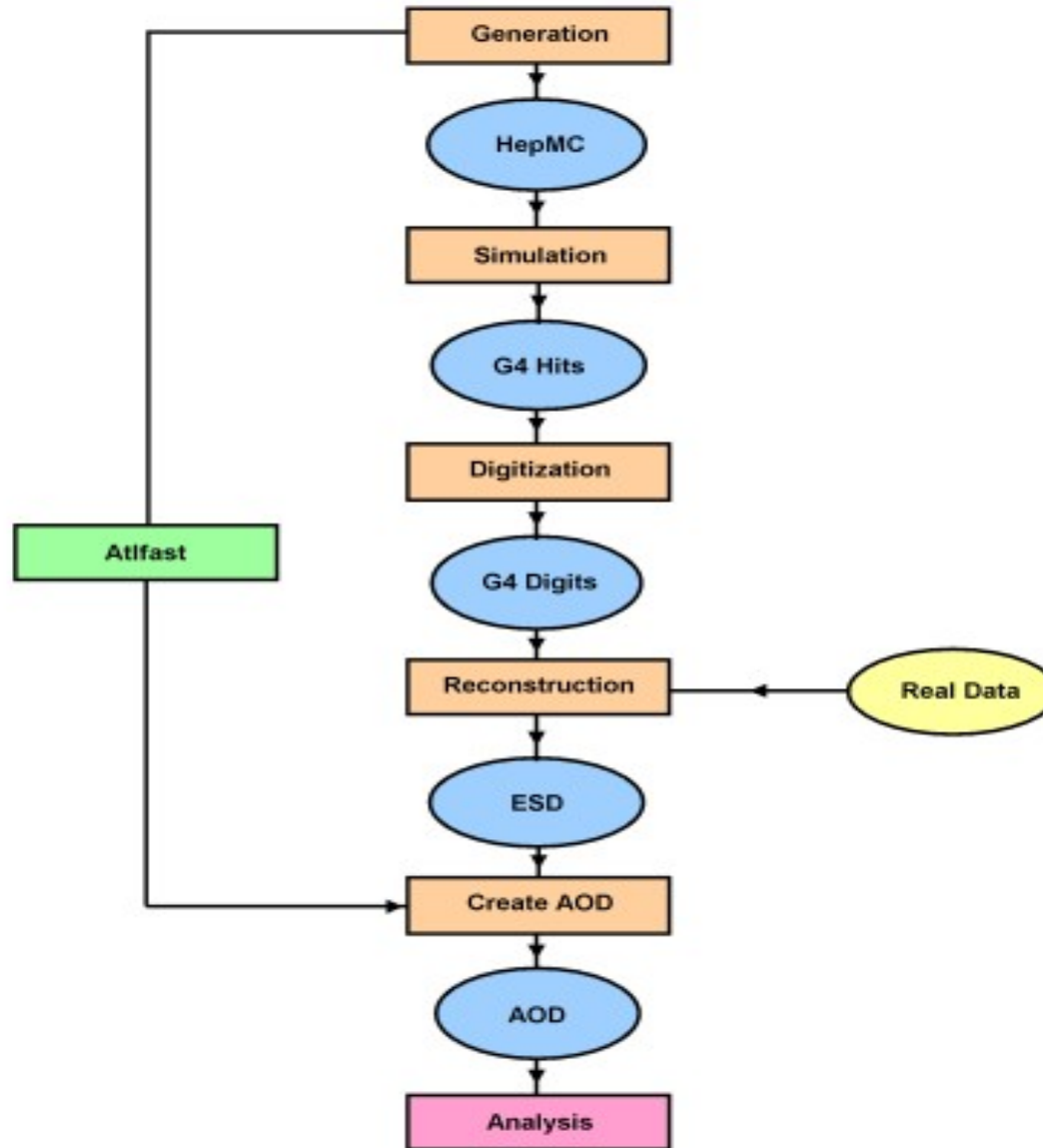
```
cmt config
```

```
source setup.sh
```

```
cd ../run
```

```
athena ../share/jobOptions.py
```

ATHENA - full chain





Cosmics in the Pit - motivation

- LHC will start commissioning with beams in 2008
- ATLAS will be completed sooner
 - Period with “nothing to do”
- Cosmic rays will be the first (and for some time the only) data for ATLAS
- We know the cosmic rays very well => they could be used for calibration and commissioning of the detector
- Need to have an idea how it will look like => Simulation of cosmic muons.



Cosmics with Pixels in Cavern

Aim:

- **simulate** cosmic rays in Pixel detector
- **determine the rate** of cosmic muons
- find and test **optimal trigger**
- **reconstruct** the data to find the **efficiency** of our software

- **COMPARE WITH THE EXPERIMENT**

Done:

- ✓ **simulate** cosmic rays in Pixel detector:
 - **2 independent samples:** **Elsa+Jamie** (cosmics for whole ATLAS)
Pavel+Michal+Martin (focused on Pixels)
- ✓ **determine the rate** of cosmic muons
- **reconstruct** the data to find the **efficiency** of our software (partially done)



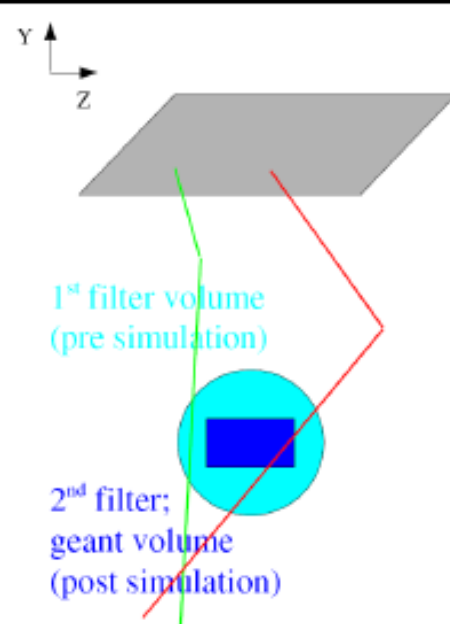
Cosmic muons simulation

Pavel Jez, Michal Marcisovsky, Martin Zeman

- Done in **Geant 4** within Athena framework
- Release **13.0.10**
- Simulation package: **G4AtlasApps-00-02-15**
- Cosmic muons generator: **CosmicGenerator-00-00-24**
 - Fixed bug with upwards heading muons
- **New package: G4PixelCosmics-00-00-01**
 - Based on G4UserAction
 - Available for testing at ~jez/public/Simulation/G4Extensions
 - Will be put into CVS and become part of nightlies



Filtering muons



- Selection criteria:

1. At the generator level:

- Only muons heading to the 10 m sphere around the origin are accepted
- Only muons with $E > 1$ GeV are accepted

➤ Rules out 99.9 % of muons

NB: Muons are generated ca. 10 m above a surface 600m x 600m over ATLAS cavern

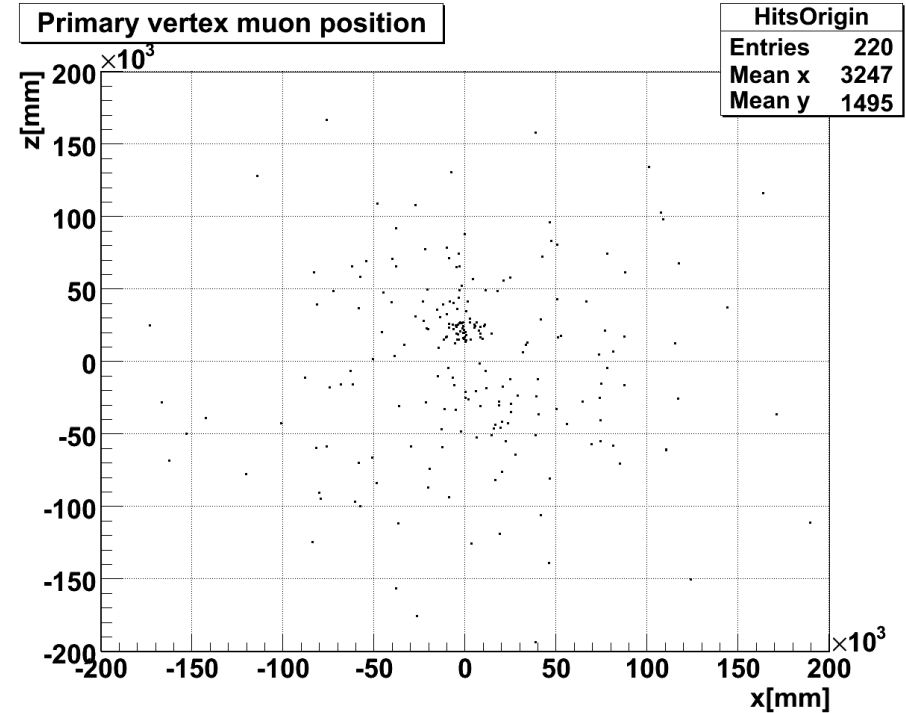
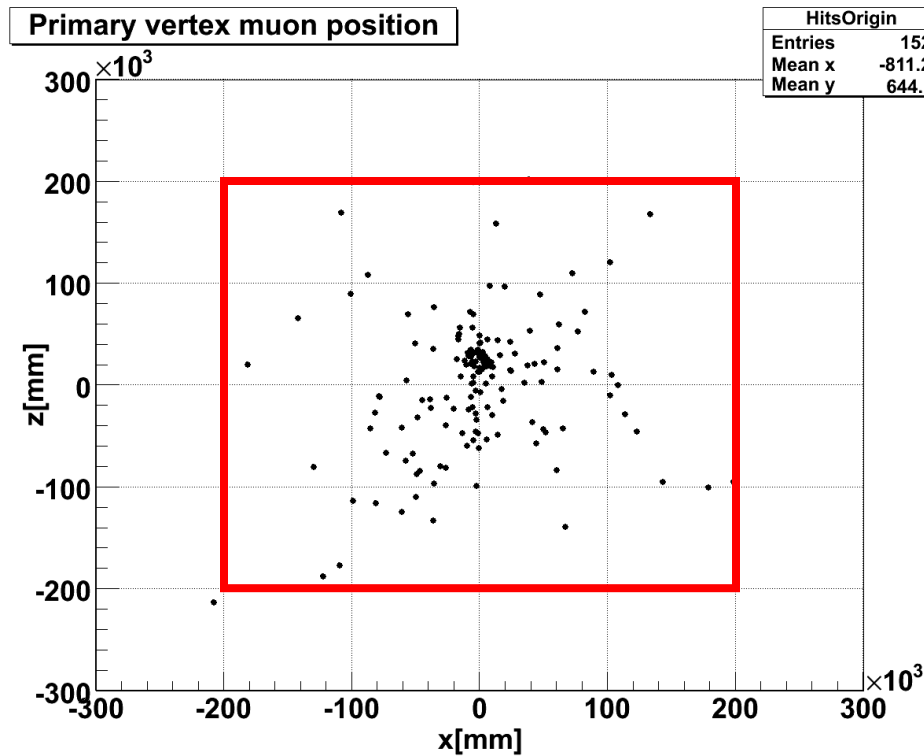
2. At the simulator/Geant 4 level:

- Done by **stepping action**
- At each step the action asks for the **name of the volume** it is in
- If the name is **Pixel:siLog** or **Pixel:siBLayLog** (i.e. active detector) and $E > 100$ MeV, the track is tagged as having gone through pixel
- Stepping action asks for **position, momentum, primary vertex and momentum, type of particle** and writes this information into a text file
- Stepping action counts **number of hits** for each track and also writes it into a text file
- **Module, ladder** and **layer** of the hit are stored as well



Cosmics in the Pit - figures

Distribution of the primary vertices



“large” sample

“standard” sample

⇒ 400 m × 400 m is completely satisfactory

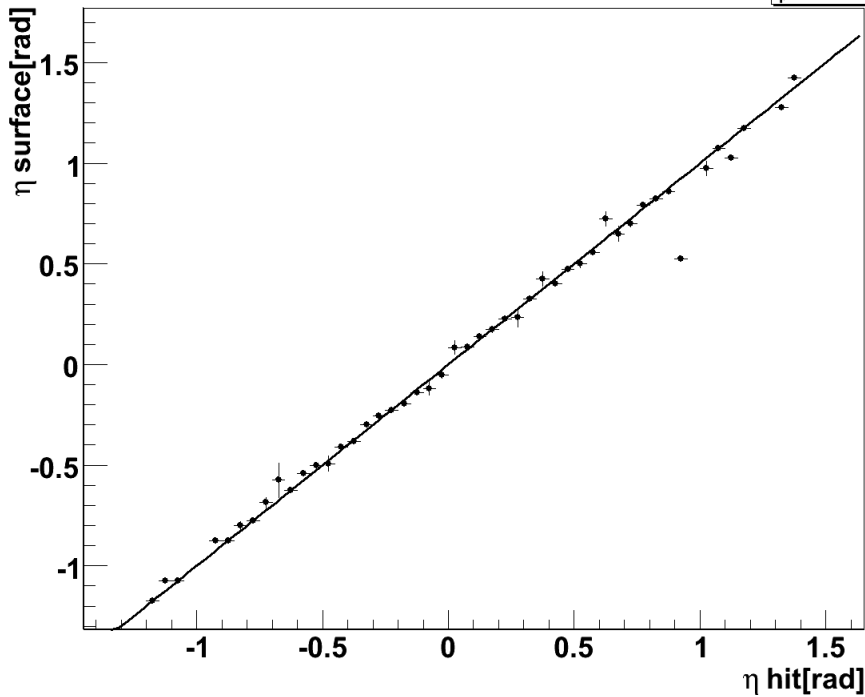


Cosmics in the Pit - figures

Fta and Phi correlation

η on the surface vs. η at the hit

ETAvsETA_pfx	
Entries	220
p0	2.032e-015
p1	1

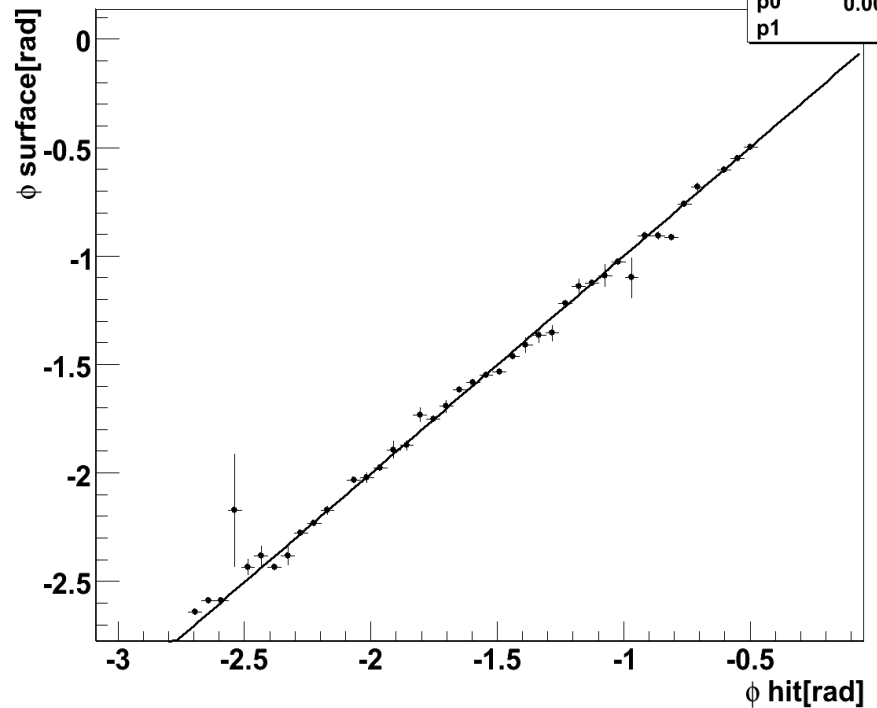


Fit: $y = x + 2e-15$

⇒ practically identity ⇒ minimal scattering

ϕ on the surface vs. ϕ at the hit

phivsphi_pfx	
Entries	220
p0	0.001132
p1	1.002



Fit: $y = 1.002x + 0.001$



Cosmics in the Pit - status

- 4 types of simulated samples:
 - “standard” = surface $400\text{ m} \times 400\text{ m}$, optimizing sphere – 10 m radius
 - Good approximation to real situation
 - Rate in Pixels – **0.85 Hz**
 - “spec1” = surface $60\text{ m} \times 60\text{ m}$, optimizing sphere – 5 m radius
 - Focused on the low energy muon through shafts
 - Rate in Pixels – **0.38 Hz**
 - “spec2” = surface $400\text{ m} \times 400\text{ m}$, optimizing sphere – 40 m radius
 - Focused on the whole cavern
 - “large” = surface $600\text{ m} \times 600\text{ m}$, optimizing sphere – 10 m radius
 - Study of the large surface effect



Cosmic muons rates standard sample

Generated	Accepted	≥ 1 pixel hit	Time	Rate
6 907 433 412	11 575 728	220	259.6 s	0.85 Hz

For reference: Result by Beate, Max and Tobi for the cosmic muon rate in Pixel detector is 0.89 ± 0.36 Hz.

Barrel hits (71 %)	1	2	3	4	5	6	7	≥ 3
Number	3	22	20	10	8	20	21	79
Rate [Hz]	0.03	0.21	0.19	0.10	0.08	0.19	0.20	0.3
Endcap Hits (29%)	1	2	3	4	5	6	7	≥ 3
Number	32	22	5	2	0	1	0	8
Rate [Hz]	0.14	0.12	0.02	0	0.02	0	0	0.02

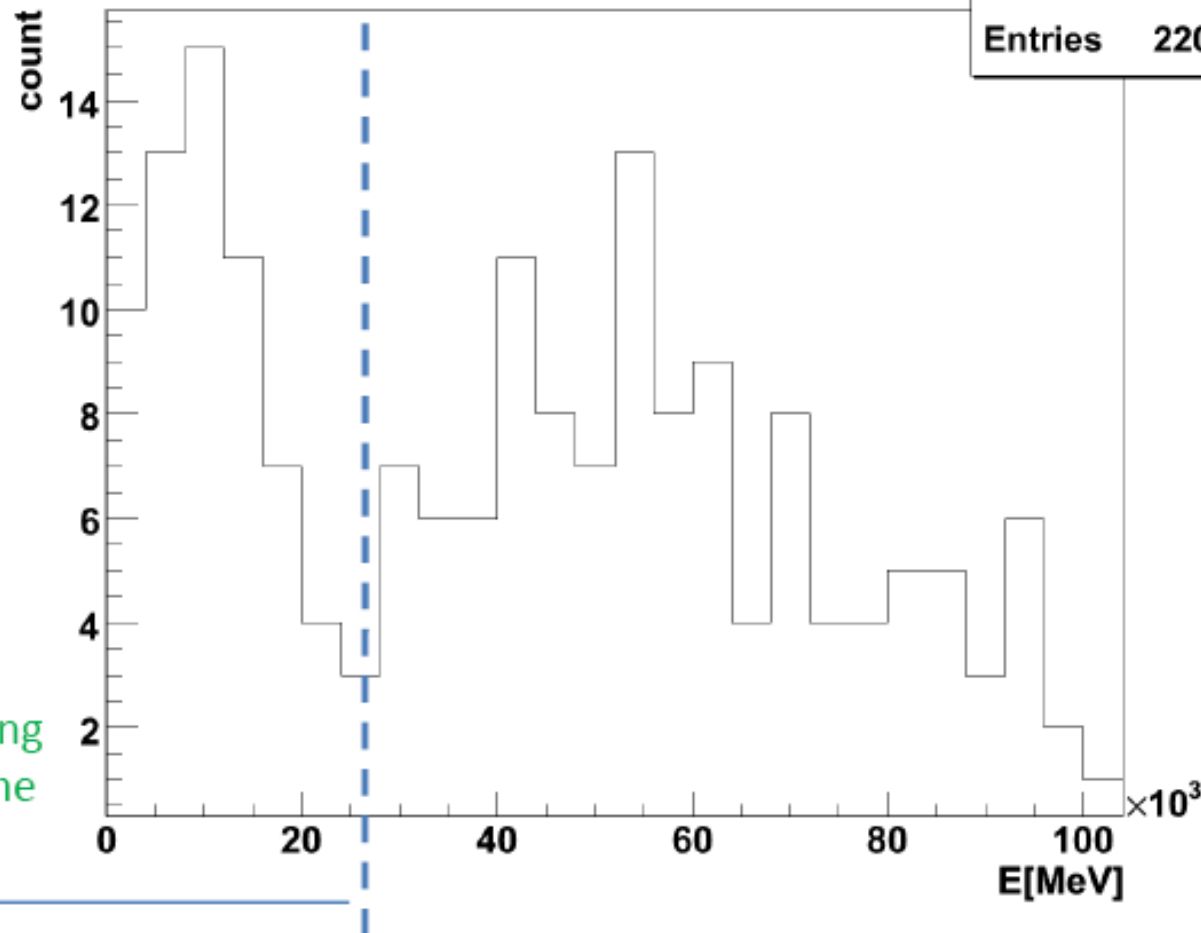


Cosmic muons distribution Energy on the surface

Energy of muons at creation

primEnergy

Entries 220



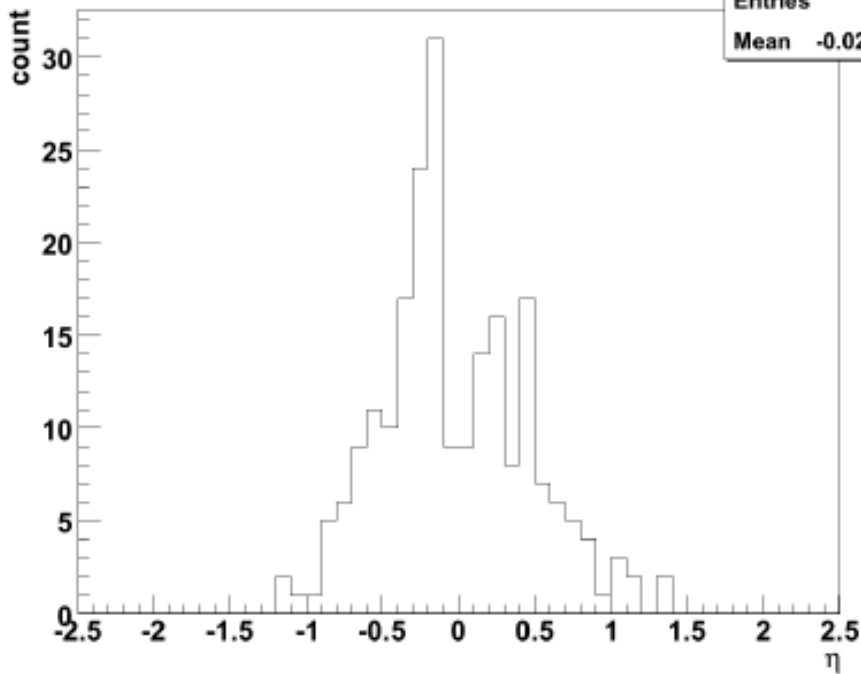
Low energy muons, spending most of the time in shafts



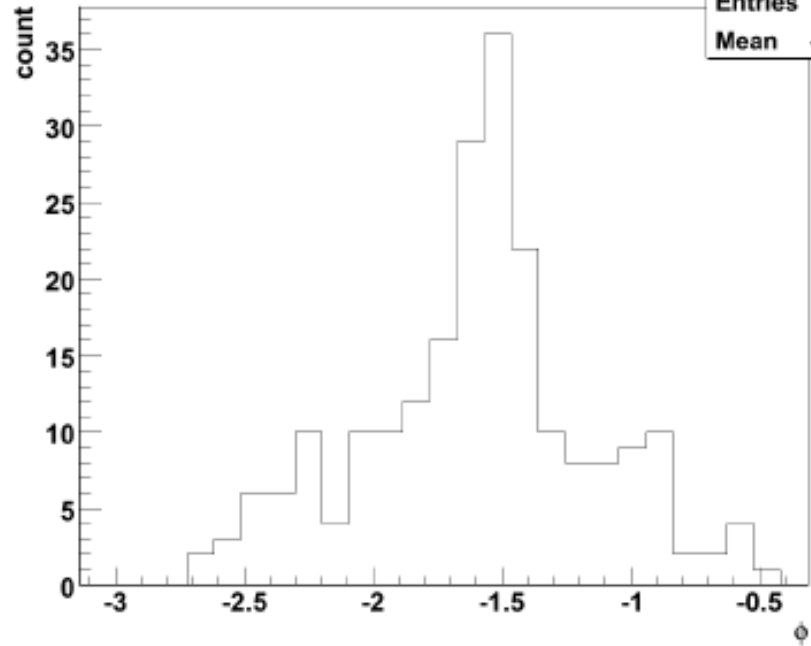


Cosmic muons distribution angular distribution

Energy of muons entering the tracker ETA

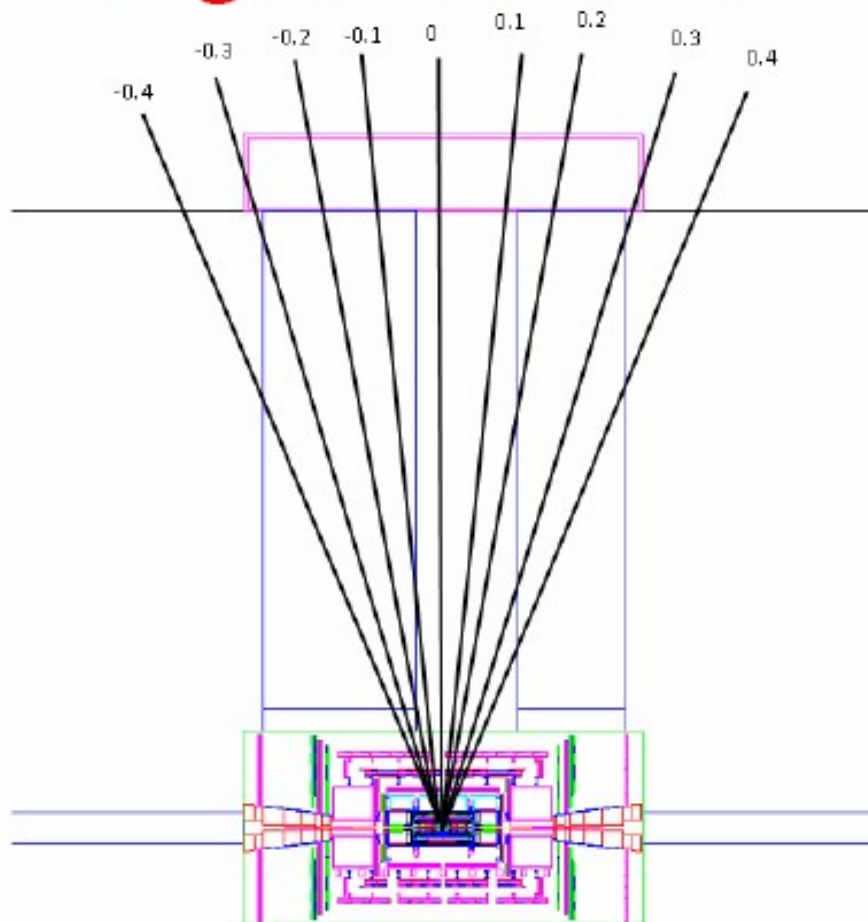


Energy of muons entering the tracker PHI





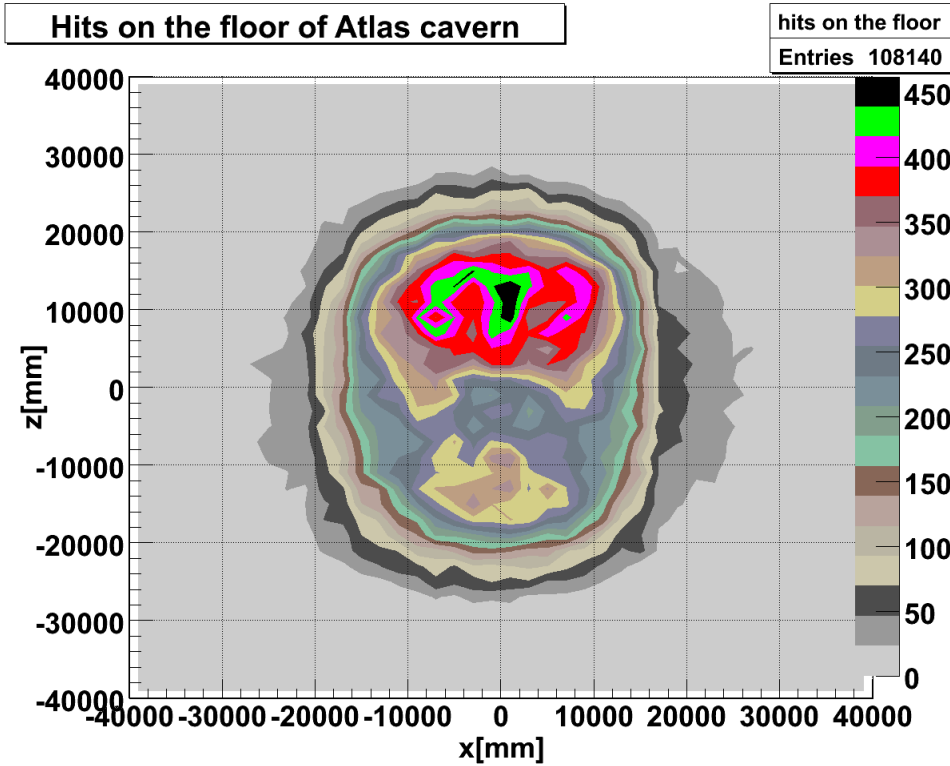
Cosmic muons distribution angular distribution



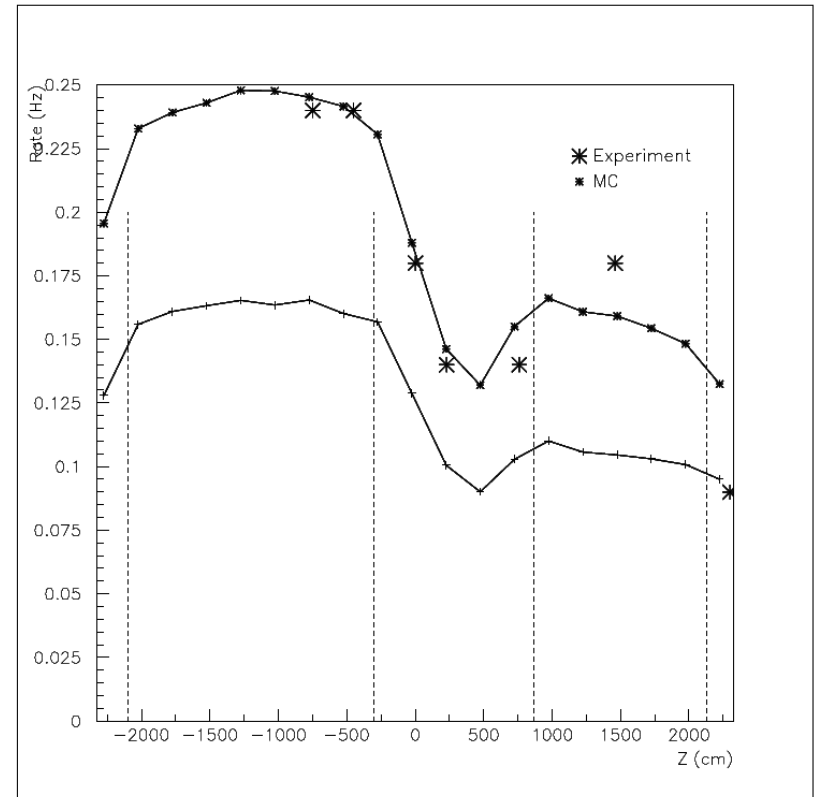


Cosmics in the Pit - figures

Cavern floor hits



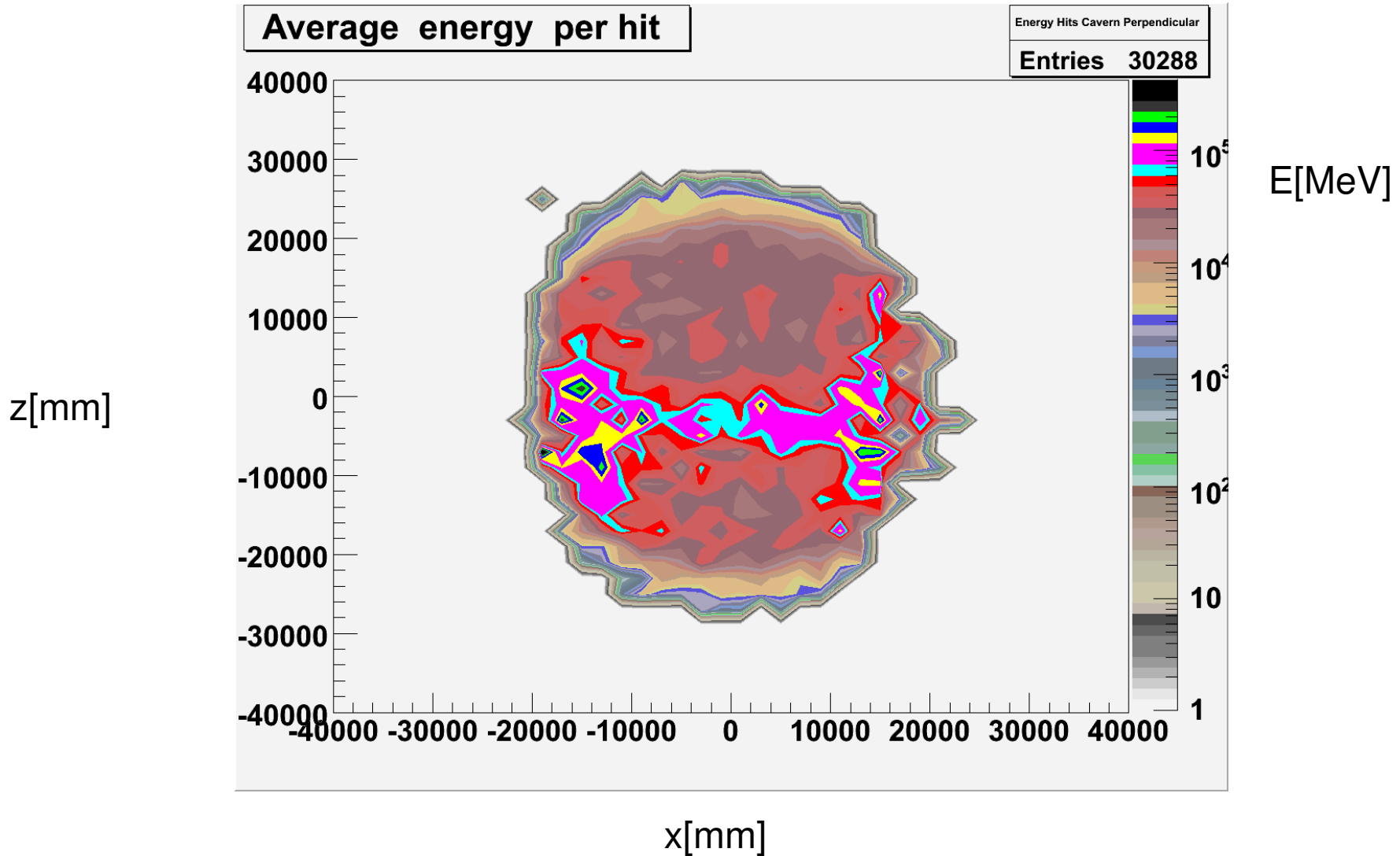
simulation



Measurement + old simulation



Cosmics in the Pit - figures



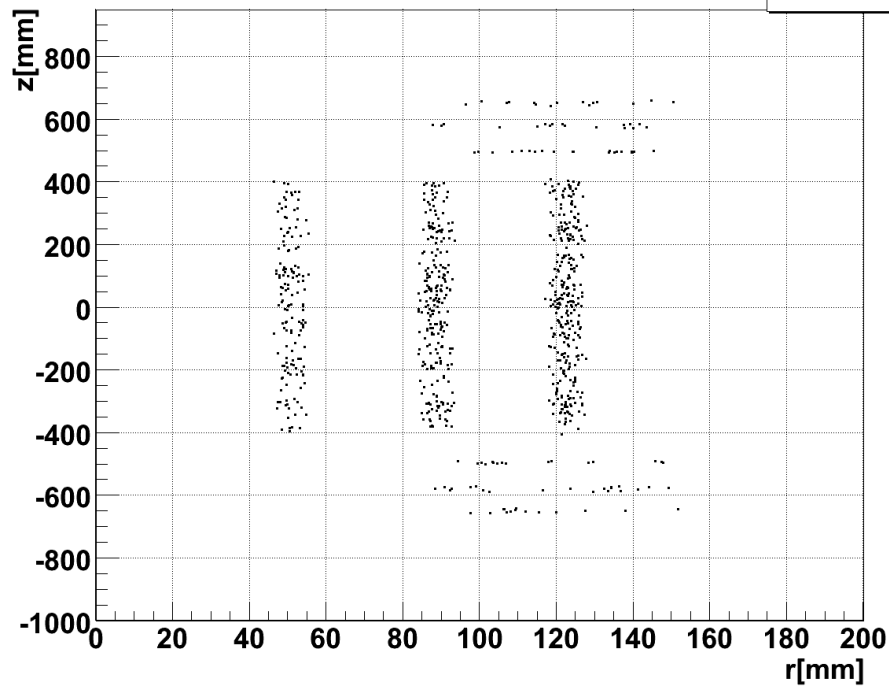


Cosmics in the Pit - figures

Hits in the Pixel detector

Hits R-Z dependence

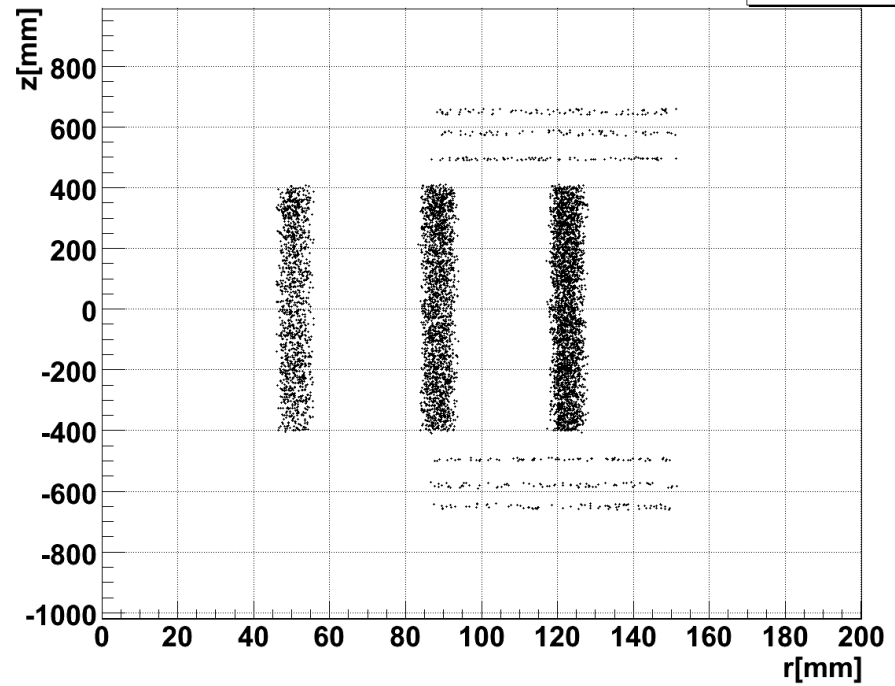
Entries 787



Standard sample

Hits RZ dependence

Entries 7016

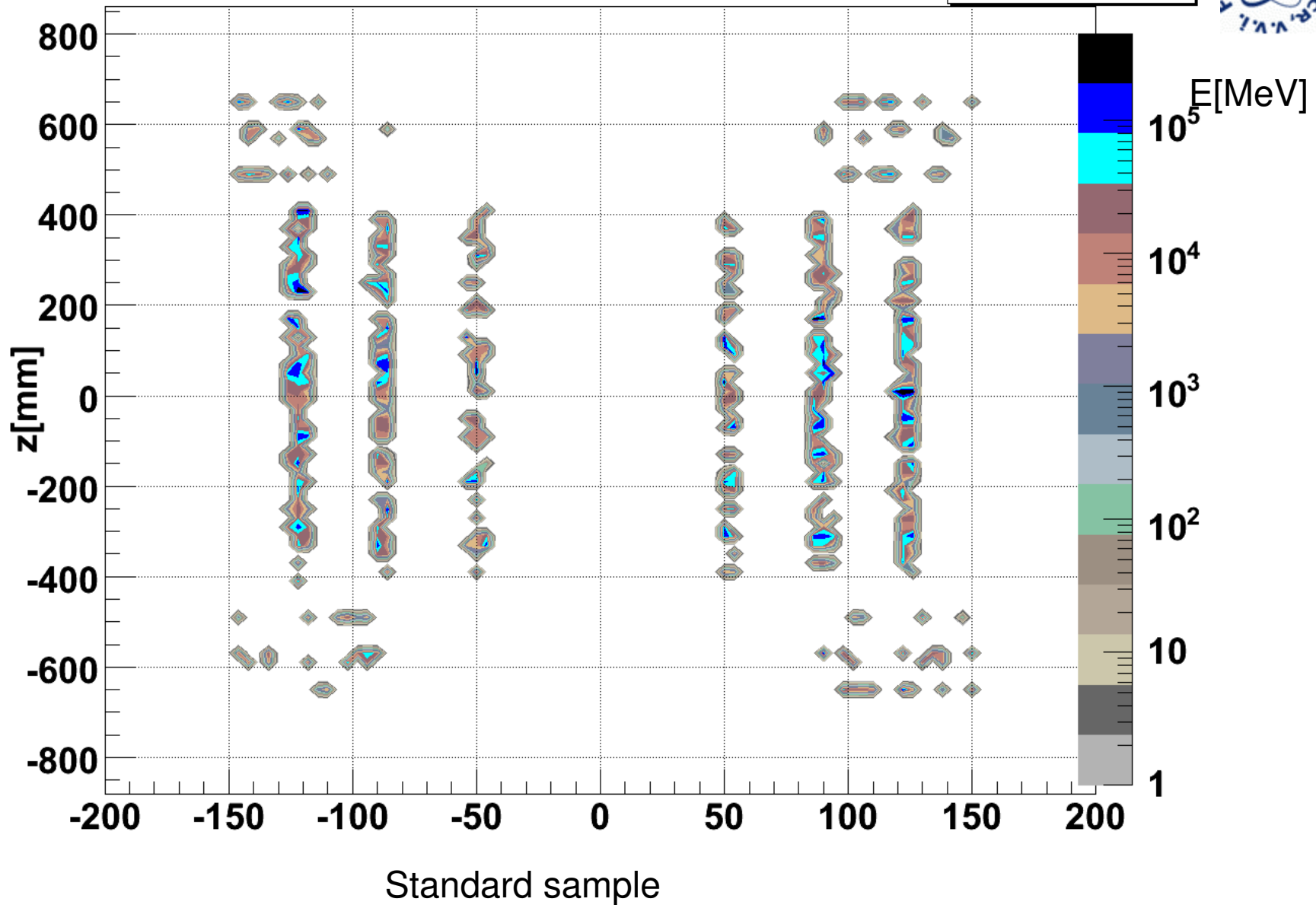


Spec1 sample



Average Energy above and below the beamline

AverageEnergy	
Entries	787

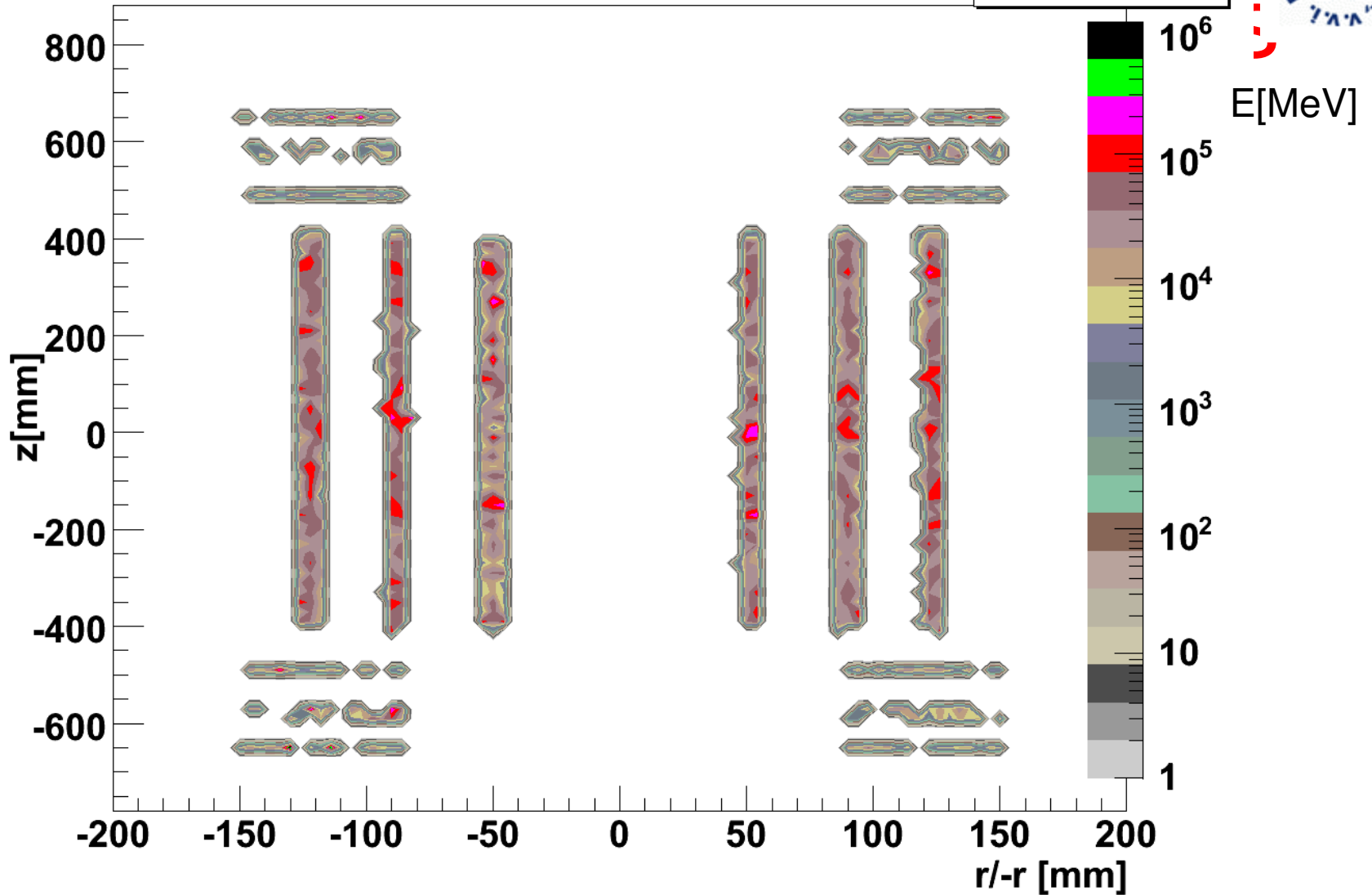




Average energy above and below beamline

AverageEnergy

Entries 7016



Spec 1 sample

Simulace kosmiky - závěr

- cross-check oficiálních výsledků
- potvrzení toho, že simulace funguje kvalitativně správně
- možnost porovnání s experimentálními daty pro SCT a TRT (ne Pixel. Stále není zřejmé, zda vůbec nějaká data budou)
- vynikající cvičení na použití Atheny a Geant4