

Dynamical systems in fluid mechanics

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Mathematical model

STATE VARIABLES

Mass density

$$\rho = \rho(t, \mathbf{x})$$

Absolute temperature

$$\vartheta = \vartheta(t, \mathbf{x})$$

Velocity field

$$\mathbf{u} = \mathbf{u}(t, \mathbf{x})$$

THERMODYNAMIC FUNCTIONS

Pressure

$$p = p(\rho, \vartheta)$$

Internal energy

$$e = e(\rho, \vartheta)$$

Entropy

$$s = s(\rho, \vartheta)$$

TRANSPORT

Viscous stress

$$\mathbb{S} = \mathbb{S}(\vartheta, \nabla_{\mathbf{x}} \mathbf{u})$$

Heat flux

$$\mathbf{q} = \mathbf{q}(\vartheta, \nabla_{\mathbf{x}} \vartheta)$$



Field equations



Claude Louis
Marie Henri
Navier
[1785-1836]

Equation of continuity

$$\partial_t \rho + \operatorname{div}_x(\rho \mathbf{u}) = 0$$

Momentum balance

$$\partial_t(\rho \mathbf{u}) + \operatorname{div}_x(\rho \mathbf{u} \otimes \mathbf{u}) + \nabla_x p(\rho, \vartheta) = \operatorname{div}_x \mathbb{S} + \rho \mathbf{f}$$



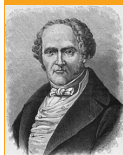
George
Gabriel
Stokes
[1819-1903]

Entropy production

$$\partial_t(\rho s(\rho, \vartheta)) + \operatorname{div}_x(\rho s(\rho, \vartheta) \mathbf{u}) + \operatorname{div}_x \left(\frac{\mathbf{q}}{\vartheta} \right) = \sigma$$

$$\sigma = (\geq) \frac{1}{\vartheta} \left(\mathbb{S} : \nabla_x \mathbf{u} - \frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} \right)$$

Constitutive relations



François Marie Charles Fourier
[1772-1837]

Fourier's law

$$\mathbf{q} = -\kappa(\vartheta)\nabla_x\vartheta$$

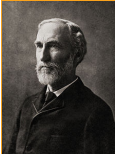


Isaac Newton
[1643-1727]

Newton's rheological law

$$\mathbb{S} = \mu(\vartheta) \left(\nabla_x \mathbf{u} + \nabla_x^t \mathbf{u} - \frac{2}{3} \operatorname{div}_x \mathbf{u} \right) + \eta(\vartheta) \operatorname{div}_x \mathbf{u} \mathbb{I}$$

Gibbs' relation



Willard Gibbs
[1839-1903]

Gibbs' relation:

$$\vartheta Ds(\varrho, \vartheta) = De(\varrho, \vartheta) + p(\varrho, \vartheta)D\left(\frac{1}{\varrho}\right)$$

Thermodynamics stability:

$$\frac{\partial p(\varrho, \vartheta)}{\partial \varrho} > 0, \quad \frac{\partial e(\varrho, \vartheta)}{\partial \vartheta} > 0$$

Boundary conditions

Impermeability

$$\mathbf{u} \cdot \mathbf{n}|_{\partial\Omega} = 0$$

No-slip

$$\mathbf{u}_{\text{tan}}|_{\partial\Omega} = 0$$

No-stick

$$[\mathbb{S}\mathbf{n}] \times \mathbf{n}|_{\partial\Omega} = 0$$

Thermal insulation

$$\mathbf{q} \cdot \mathbf{n}|_{\partial\Omega} = 0$$

A bit of history of global existence for large data



Jean Leray - Royal society (1998)

Jean Leray [1906-1998]
Global existence of weak
solutions for the
incompressible
Navier-Stokes system (3D)



**Olga Aleksandrovna
Ladyzhenskaya**
[1922-2004] Global
existence of classical
solutions for the
incompressible 2D
Navier-Stokes system



Pierre-Louis Lions [*1956] Global existence of weak
solutions for the compressible barotropic Navier-Stokes
system (2,3D)

and many, many others...



Weak solutions to the complete system

- Equation of continuity holds in the sense of distributions (renormalized equation also satisfied)
- Momentum balance holds in the sense of distributions
- Entropy production equation holds in the sense of distributions, entropy production rate satisfies the inequality
- The system is augmented by

Total energy balance

$$\frac{d}{dt} \int_{\Omega} \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho, \vartheta) - \varrho F \right) dx = 0$$

Technical hypotheses

Pressure

$$p(\varrho, \vartheta) = \vartheta^{5/2} P\left(\frac{\varrho}{\vartheta^{3/2}}\right) + \frac{a}{3} \vartheta^4$$

$$P(0) = 0, P'(Z) > 0, P(Z)/Z^{5/3} \rightarrow p_\infty > 0 \text{ as } Z \rightarrow \infty$$

Internal energy

$$e(\varrho, \vartheta) = \frac{3}{2} \vartheta \frac{\vartheta^{3/2}}{\varrho} P\left(\frac{\varrho}{\vartheta^{3/2}}\right) + \frac{a}{\varrho} \vartheta^4$$

Transport coefficients

$$\mu(\vartheta) \approx (1 + \vartheta^\alpha), \alpha \in [1/2, 1], \kappa(\vartheta) \approx (1 + \vartheta^3)$$

Conservative vs. dissipative system

Conservative character

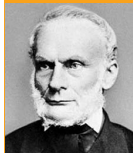
$$\text{total mass } \int_{\Omega} \varrho(t, \cdot) \, dx = M_0,$$

$$\text{total energy } \int_{\Omega} \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho, \vartheta) - \varrho F \right) (t, \cdot) \, dx = E_0$$

Dissipative character

$$\text{total entropy } \int_{\Omega} \varrho s(\varrho, \vartheta) \, dx = S(t) \nearrow S_{\infty}$$

Uniform stabilization to equilibria



DIE ENERGIE DER WELT IST CONSTANT;
DIE ENTROPIE DER WELT
STREBT EINEM MAXIMUM ZU

Rudolph Clausius, 1822-1888

Equilibrium solutions

CONSERVATIVE DRIVING FORCE

$$\mathbf{f} = \nabla_x F, \quad F = F(x)$$

TOTAL ENERGY CONSERVATION

$$\frac{d}{dt} \int_{\Omega} \left(\frac{1}{2} \rho |\mathbf{u}|^2 + \rho e(\rho, \vartheta) - \rho F \right) dx = 0$$

Static solutions

$$\nabla_x \rho(\tilde{\rho}, \bar{\vartheta}) = \tilde{\rho} \nabla_x F, \quad \bar{\vartheta} > 0 \text{ constant}$$

Total mass and energy

$$\int_{\Omega} \tilde{\rho} dx = M_0, \quad \int_{\Omega} (\tilde{\rho} e(\tilde{\rho}, \bar{\vartheta}) - \tilde{\rho} F) dx = E_0$$

Total dissipation balance

Ballistic free energy

$$H_{\Theta}(\varrho, \vartheta) = \varrho \left(e(\varrho, \vartheta) - \Theta s(\varrho, \vartheta) \right)$$

Relative entropy

$$\begin{aligned} & \mathcal{E}(\varrho, \vartheta, \mathbf{u} | \tilde{\varrho}, \bar{\vartheta}) \\ &= \int_{\Omega} \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + H_{\bar{\vartheta}}(\varrho, \vartheta) - \partial_{\varrho} H_{\bar{\vartheta}}(\tilde{\varrho}, \bar{\vartheta})(\varrho - \tilde{\varrho}) - H_{\bar{\vartheta}}(\tilde{\varrho}, \bar{\vartheta}) \right) dx \end{aligned}$$

Total dissipation balance

$$\frac{d}{dt} \mathcal{E}(\varrho, \vartheta, \mathbf{u} | \tilde{\varrho}, \bar{\vartheta}) + \int_{\Omega} \sigma \, dx = 0$$

$\tilde{\varrho}, \bar{\vartheta}$ – equilibrium state

Thermodynamic stability

Positive compressibility and specific heat

$$\frac{\partial p(\varrho, \vartheta)}{\partial \varrho} > 0, \quad \frac{\partial e(\varrho, \vartheta)}{\partial \vartheta} > 0$$

Coercivity of the ballistic free energy

$\varrho \mapsto H_{\Theta}(\varrho, \Theta)$ strictly convex

$\vartheta \mapsto H_{\Theta}(\varrho, \vartheta)$ decreasing for $\vartheta < \Theta$ and increasing for $\vartheta > \Theta$

Long-time behavior for conservative driving forces

$$\mathbf{f} = \nabla_x F, \quad F = F(x)$$

$$\varrho(t, \cdot) \rightarrow \tilde{\varrho} \text{ in } L^{5/3}(\Omega) \text{ as } t \rightarrow \infty$$

$$\vartheta(t, \cdot) \rightarrow \bar{\vartheta} \text{ in } L^4(\Omega) \text{ as } t \rightarrow \infty$$

$$(\varrho \mathbf{u})(t, \cdot) \rightarrow 0 \text{ in } L^1(\Omega; \mathbb{R}^3) \text{ as } t \rightarrow \infty$$

Attractors

Hypotheses

$$\int_{\Omega} \varrho(t, \cdot) \, dx > M_0, \quad t > 0$$

$$\int_{\Omega} \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho, \vartheta) - \varrho F \right) (t, \cdot) \, dx < E_0, \quad t > 0$$

$$\int_{\Omega} \varrho s(\varrho, \vartheta)(t, \cdot) \, dx > S_0, \quad t > 0$$

Conclusion

$$\|\varrho(t, \cdot) - \tilde{\varrho}\|_{L^{5/3}(\Omega)} < \varepsilon, \quad \|\vartheta(t, \cdot) - \bar{\vartheta}\|_{L^4(\Omega)} < \varepsilon \quad \text{for } t > T(\varepsilon)$$

$$\|\varrho \mathbf{u}(t, \cdot)\|_{L^1(\Omega; \mathbb{R}^3)} < \varepsilon \quad \text{for } t > T(\varepsilon)$$

Uniform decay of density oscillations

$$\partial_t \varrho_\varepsilon + \mathbf{u}_\varepsilon \cdot \nabla_x \varrho_\varepsilon = -\operatorname{div}_x \mathbf{u}_\varepsilon \varrho_\varepsilon$$

$$\varrho_\varepsilon \rightarrow \varrho, \quad \varrho_\varepsilon \log(\varrho_\varepsilon) \rightarrow \overline{\varrho \log(\varrho)} \text{ weakly in } L^1$$

$$d(t) = \int_{\Omega} \left(\overline{\varrho \log(\varrho)} - \varrho \log(\varrho) \right)(t, \cdot) \, dx$$

Density oscillations decay

$$\partial_t d(t) + \Psi(d(t)) \leq 0$$

$$\Psi(0) = 0, \quad \Psi(d) > 0 \text{ for } d > 0.$$

General time-dependent driving forces

$$\mathbf{f} = \mathbf{f}(t, \mathbf{x}), \quad |\mathbf{f}(t, \mathbf{x})| \leq \bar{F}$$

EITHER

$$E(t) \equiv \int_{\Omega} \left(\frac{1}{2} \rho |\mathbf{u}|^2 + \rho e(\rho, \vartheta) \right) (t, \cdot) \, dx \rightarrow \infty \text{ as } t \rightarrow \infty$$

OR

$$|E(t)| \leq E \text{ for a.a. } t > 0$$

In the case $E(t) \leq E$, each sequence of times $\tau_n \rightarrow \infty$ contains a subsequence such that

$$\mathbf{f}(\tau_n + \cdot, \cdot) \rightarrow \nabla_x F \text{ weakly-} (*) \text{ in } L^\infty((0, 1) \times \Omega),$$

where $F = F(x)$ may depend on $\{\tau_n\}$

STEP 1:

Assume that $E(\tau_n) < E$ for certain $\tau_n \rightarrow \infty \Rightarrow$ total entropy remains bounded \Rightarrow integral of entropy production bounded

STEP 2:

For $\tau_n \rightarrow \infty$ we have $\nabla_x p(\rho, \vartheta) \approx \rho \mathbf{f}$, $\vartheta \approx \bar{\vartheta}$, meaning, $\mathbf{f} \approx \nabla_x F$

STEP 3:

The energy cannot “oscillate” since bounded entropy *static solutions* have bounded total energy

Corollaries



$$\mathbf{f} = \mathbf{f}(x) \neq \nabla_x F$$

$$\Rightarrow$$

$$E(t) \rightarrow \infty$$



$\mathbf{f} = \mathbf{f}(t, x)$ (almost) periodic in time, $\mathbf{f} \neq \nabla_x F$, $F = F(x)$

$$\Rightarrow$$

$$E(t) \rightarrow \infty$$

Rapidly oscillating driving forces

Hypotheses:

$$\mathbf{f} = \omega(t^\beta) \mathbf{w}(x), \mathbf{w} \in W^{1,\infty}(\Omega; \mathbb{R}^3), \beta > 2$$

$$\omega \in L^\infty(\mathbb{R}), \sup_{\tau > 0} \left| \int_0^\tau \omega(t) dt \right| < \infty$$

Conclusion:

$$(\rho \mathbf{u})(t, \cdot) \rightarrow 0 \text{ in } L^1(\Omega; \mathbb{R}^3) \text{ as } t \rightarrow \infty$$

$$\varrho(t, \cdot) \rightarrow \bar{\varrho} \text{ in } L^{5/3}(\Omega) \text{ as } t \rightarrow \infty$$

$$\vartheta(t, \cdot) \rightarrow \bar{\vartheta} \text{ in } L^4(\Omega) \text{ as } t \rightarrow \infty$$

Rapidly oscillating growing driving forces

Hypotheses:

$$\mathbf{f} = t^\delta \omega(t^\beta) \mathbf{w}(x), \mathbf{w} \in W^{1,\infty}(\Omega; \mathbb{R}^3)$$

$$\delta > 0, \beta - 2\delta > 2 \text{ or } \delta \leq 0, \beta - \delta > 2$$

$$\omega \in L^\infty(\mathbb{R}), \sup_{\tau > 0} \left| \int_0^\tau \omega(t) dt \right| < \infty$$

Conclusion:

$$(\rho \mathbf{u})(t, \cdot) \rightarrow 0 \text{ in } L^1(\Omega; \mathbb{R}^3) \text{ as } t \rightarrow \infty$$

$$\varrho(t, \cdot) \rightarrow \bar{\varrho} \text{ in } L^{5/3}(\Omega) \text{ as } t \rightarrow \infty$$

$$\vartheta(t, \cdot) \rightarrow \bar{\vartheta} \text{ in } L^4(\Omega) \text{ as } t \rightarrow \infty$$

Time-periodic solutions and boundary dissipation

Dissipative boundary conditions

$$\mathbf{u}|_{\partial\Omega} = 0, \quad \mathbf{q} \cdot \mathbf{n} = d(x)(\vartheta - \tilde{\vartheta})$$

Time periodic forcing

$$\mathbf{f}(t + \omega, \cdot) = \mathbf{f}(t, \cdot)$$

Time periodic solutions

$$\varrho(t + \omega, \cdot) = \varrho(t, \cdot), \quad \vartheta(t + \omega, \cdot) = \vartheta(t, \cdot), \quad \mathbf{u}(t + \omega, \cdot) = \mathbf{u}(t, \cdot)$$