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## An anelastic approximation arising in astrophysics

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Preprint No. 18-2016
PRAHA 2016

# An anelastic approximation arising in astrophysics 

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#### Abstract

We identify the asymptotic limit of the compressible non-isentropic NavierStokes system in the regime of low Mach, low Froude and high Reynolds number. The system is driven by a long range gravitational potential. We show convergence to an anelastic system for ill-prepared initial data. The proof is based on frequency localized Strichartz estimates for the acoustic equation based on the recent work of Metcalfe and Tataru.


Keywords: Anelastic approximation; low Mach number limit, high Reynolds number limit; stratified flow

## 1 Introduction

We consider a low Mach number approximation of the dynamics of certain type of supernovas

$$
\begin{equation*}
\operatorname{div}_{x}\left(\varrho_{0} \mathbf{V}\right)=0 \tag{1.1}
\end{equation*}
$$

[^0]\[

$$
\begin{align*}
\partial_{t} \mathbf{V}+\mathbf{V} \cdot \nabla_{x} \mathbf{V}+\nabla_{x} \Pi & =-\frac{\varrho_{0}}{\mathcal{R}} \nabla_{x} F  \tag{1.2}\\
\partial_{t}(\mathcal{R})+\operatorname{div}_{x}(\mathcal{R} \mathbf{V}) & =0 \tag{1.3}
\end{align*}
$$
\]

with the unknown density $\mathcal{R}$, velocity $\mathbf{V}$, and the pressure $\Pi$, see Almgren et al. [1], [2], [3]. The background density profile $\varrho_{0}$ is determined as the unique solution of the static equation

$$
\begin{equation*}
\nabla_{x} \varrho_{0}^{\gamma}=\varrho_{0} \nabla_{x} F \text { in } R^{3}, \varrho_{0} \rightarrow \bar{\varrho}>0 \text { as }|x| \rightarrow \infty \tag{1.4}
\end{equation*}
$$

A similar system, written in terms of the potential temperature $\mathcal{T}=\varrho_{0} / \mathcal{R}$, has been identified as a low Mach number approximation of strongly stratified fluid flows in meteorology, see [10]. The principal difference between [10] and the present problem is the geometry of the physical space; an infinite strip is relevant in meteorology, while the whole Euclidean space $R^{3}$ is considered in astrophysics mimicking a large neighborhood of a gaseous star. Accordingly, the potential $F$ shares the asymptotic properties with the standard gravitational potential:

$$
\begin{align*}
F \in C^{\infty}\left(R^{3}\right), F(x) & >0 \text { for all } x \in R^{3},  \tag{1.5}\\
\underline{F} \frac{1}{|x|} & \leq F(x) \leq \bar{F} \frac{1}{|x|} \text { for all }|x|>R,  \tag{1.6}\\
|x|^{2}\left|\nabla_{x} F(x)\right|+|x|^{3}\left|\nabla_{x}^{2} F(x)\right| & \leq c \text { for all } x \in R^{3} . \tag{1.7}
\end{align*}
$$

Our goal is to identify system (1.1-1.4) as the asymptotic limit of the primitive system

$$
\begin{align*}
\partial_{t} \varrho+\operatorname{div}_{x}(\varrho \mathbf{u}) & =0  \tag{1.8}\\
\partial_{t}(\varrho \mathbf{u})+\operatorname{div}_{x}(\varrho \mathbf{u} \otimes \mathbf{u})+\frac{1}{\varepsilon^{2}} \nabla_{x}(\varrho \Theta)^{\gamma} & =\varepsilon^{\alpha} \operatorname{div}_{x} \mathbb{S}\left(\nabla_{x} \mathbf{u}\right)+\frac{1}{\varepsilon^{2}} \varrho \nabla_{x} F  \tag{1.9}\\
\partial_{t}(\varrho \Theta)+\operatorname{div}_{x}(\varrho \Theta \mathbf{u}) & =0  \tag{1.10}\\
\mathbb{S}\left(\nabla_{x} \mathbf{u}\right)=\left(\nabla_{x} \mathbf{u}+\nabla_{x}^{t} \mathbf{u}-\frac{2}{3} \operatorname{div}_{x} \mathbf{u} \mathbb{I}\right) & +\lambda \operatorname{div}_{x} \mathbf{u} \mathbb{I}, \lambda \geq 0, \tag{1.11}
\end{align*}
$$

supplemented with the far-field conditions

$$
\begin{equation*}
\varrho \rightarrow \bar{\varrho}, \mathbf{u} \rightarrow 0, \Theta \rightarrow 1 \text { as }|x| \rightarrow \infty \tag{1.12}
\end{equation*}
$$

where the parameter $\varepsilon \rightarrow 0$, and $\alpha>0$. Equations (1.8-1.10) represent a scaled isentropic Navier-Stokes system written in terms of the density $\varrho$, the velocity $\mathbf{u}$, and the potential temperature $\Theta$. The relevant value of the adiabatic exponent is $\gamma=\frac{5}{3}$ for the monoatomic gas state equation. The singular scaling factors in (1.9) are the Mach number and the Froude number proportional to $\varepsilon^{2}$ and the Reynolds number proportional to $\varepsilon^{-\alpha}$. The relevance of the regime $\varepsilon \rightarrow 0$ in astrophysical models is discussed in [2].

We consider the singular limit $\varepsilon \rightarrow 0$ for the ill-prepared initial data:

$$
\begin{align*}
\varrho(0, \cdot) & =\varrho_{0}+\varepsilon \varrho_{0, \varepsilon}^{(1)},\left\|\varrho_{0, \varepsilon}^{(1)}\right\|_{L^{1} \cap L^{\infty}\left(R^{3}\right)} \leq c, \varrho_{0, \varepsilon}^{(1)} \rightarrow \varrho_{0}^{(1)} \text { in } L^{1}\left(R^{3}\right)  \tag{1.13}\\
\mathbf{u}(0, \cdot) & =\mathbf{u}_{0, \varepsilon},\left\|\mathbf{u}_{0, \varepsilon}\right\|_{L^{2} \cap L^{\infty}\left(R^{3}\right)} \leq c, \mathbf{u}_{0, \varepsilon} \rightarrow \mathbf{v}_{0} \text { in } L^{2}\left(R^{3}\right),  \tag{1.14}\\
\Theta(0, \varepsilon) & =1+\varepsilon^{2} \Theta_{0, \varepsilon}^{(2)},\left\|\Theta_{0, \varepsilon}^{(2)}\right\|_{L^{1} \cap L^{\infty}\left(R^{3}\right)} \leq c, \Theta_{0, \varepsilon}^{(2)} \rightarrow \Theta_{0}^{(2)} \text { in } L^{1}\left(R^{3}\right) . \tag{1.15}
\end{align*}
$$

In particular, we do not require the limit velocity field $\mathbf{v}_{0}$ to satisfy the anelastic constraint (1.1).

In contrast with [10], where the problem for ill-prepared data was considered only in the viscous regime (constant positive Reynolds number), we are able to handle the vanishing viscosity limit in the presence of the disturbing effect of acoustic waves generated by the ill-prepared data. For this kind of problem, the lack of compactness of the velocity must be compensated by the structural stability properties of the system encoded in the relative energy inequality, identified for the present system in [10].

As the underlying physical space is unbounded, the effect of acoustic waves is likely to be annihilated by dispersion. Unfortunately, to our best knowledge, no direct dispersive estimates are available for the acoustic system in question. The problem is that we need rather strong global in time and space estimates of Strichartz's type for a non-constant coefficient wave operator

$$
\mathcal{A}_{\varrho_{0}}: v \mapsto-\frac{p^{\prime}\left(\varrho_{0}\right)}{\varrho_{0}} \operatorname{div}_{x}\left(\varrho_{0} \nabla_{x} v\right)
$$

with $\varrho_{0}$ generated by the long-range potential $F$. Luckily, however, it is enough to have these estimates for a "dense" set of initial data, in particular, we may work on a frequency localized space, where the low and high frequencies are cut-off. Our approach, reminiscent of the decomposition technique used by Smith and Sogge [17], can be summarized as follows.

- Use the result of De Bièvre and Pravica [4], [5] to show that the point spectrum of $\mathcal{A}_{\varrho_{0}}$ is empty and that $\mathcal{A}_{\varrho_{0}}$ satisfies the Limiting absorption principle on a weighted Hilbert space.
- Show frequency and space localized decay estimates by means of the method proposed in [7]. These are estimates on solutions of the wave equation generated by $\mathcal{A}_{\varrho_{0}}$ localized in both the physical and frequency space.
- Combine the local estimates with the global result of Metcalfe and Tataru [15]. As a result we obtain frequency localized Strichartz estimates for $\mathcal{A}_{\varrho_{0}}$ that may be of independent interest, see Section 4.

The paper is organized as follows. We start with some preliminary material in Section 2, introducing the concept of dissipative solution to the primitive system, discussing solvability of the target system, and, finally, stating the main result. In Section

3 , we introduce the relative energy inequality and derive the necessary uniform bounds for solutions of the primitive system independent of the scaling parameters. Section 4 is the heart of the paper. We study the underlying acoustic equation and establish the dispersive estimates of Strichartz's type for frequency localized data. The proof of the main result is completed in Section 5.

## 2 Preliminaries, main result

We introduce the concepts of dissipative weak solution for the primitive system (1.8$1.12)$ and strong solution for the target system (1.1-1.4). Solutions of both systems will be considered on the whole physical space $R^{3}$, with relevant far field conditions for $|x| \rightarrow \infty$.

### 2.1 Dissipative solutions of the primitive system

We consider the weak solutions of system (1.8-1.12) in the class

$$
\begin{align*}
(\varrho-\bar{\varrho}) & \in C_{\text {weak }}\left([0, T] ;\left(L^{2}+L^{\gamma}\right)\left(R^{3}\right)\right)  \tag{2.1}\\
\varrho \mathbf{u} & \in C_{\text {weak }}\left([0, T] ;\left(L^{2}+L^{\frac{2 \gamma}{\gamma+1}}\right)\left(R^{3} ; R^{3}\right)\right)  \tag{2.2}\\
\varrho(\Theta-1) & \in C_{\text {weak }}\left([0, T] ;\left(L^{2}+L^{\gamma}\right)\left(R^{3}\right)\right), \Theta \in L^{\infty}\left((0, T) \times R^{3}\right), \tag{2.3}
\end{align*}
$$

satisfying

$$
\begin{equation*}
\left[\int_{R^{3}} \varrho \varphi \mathrm{~d} x\right]_{t=\tau_{1}}^{t=\tau_{2}}=\int_{\tau_{1}}^{\tau_{2}} \int_{R^{3}}\left[\varrho \partial_{t} \varphi+\varrho \mathbf{u} \cdot \nabla_{x} \varphi\right] \mathrm{d} x \mathrm{~d} t \tag{2.4}
\end{equation*}
$$

for any $0 \leq \tau_{1} \leq \tau_{2} \leq T$, and for any $\varphi \in C_{c}^{1}\left([0, T] \times R^{3}\right)$;

$$
\begin{aligned}
{\left[\int_{R^{3}} \varrho \mathbf{u} \cdot \varphi \mathrm{~d} x\right]_{t=\tau_{1}}^{t=\tau_{2}} } & =\int_{\tau_{1}}^{\tau_{2}} \int_{R^{3}}\left[\varrho \mathbf{u} \cdot \partial_{t} \varphi+\varrho \mathbf{u} \otimes \mathbf{u}: \nabla_{x} \varphi+\frac{1}{\varepsilon^{2}}(\varrho \Theta)^{\gamma} \operatorname{div}_{x} \varphi\right] \mathrm{d} x \mathrm{~d} t \\
& -\int_{\tau_{1}}^{\tau_{2}} \int_{R^{3}}\left[\varepsilon^{\alpha} \mathbb{S}\left(\nabla_{x} \mathbf{u}\right): \nabla_{x} \varphi-\frac{1}{\varepsilon^{2}} \varrho \nabla_{x} F \cdot \varphi\right] \mathrm{d} x \mathrm{~d} t
\end{aligned}
$$

for any $0 \leq \tau_{1} \leq \tau_{2} \leq T$, and for any $\varphi \in C_{c}^{1}\left([0, T] \times R^{3} ; R^{3}\right)$;

$$
\begin{equation*}
\left[\int_{R^{3}} \varrho G(\Theta) \varphi \mathrm{d} x\right]_{t=\tau_{1}}^{t=\tau_{2}}=\int_{\tau_{1}}^{\tau_{2}} \int_{R^{3}}\left[\varrho G(\Theta) \partial_{t} \varphi+\varrho G(\Theta) \mathbf{u} \cdot \nabla_{x} \varphi\right] \mathrm{d} x \mathrm{~d} t \tag{2.5}
\end{equation*}
$$

$$
\begin{equation*}
\text { for any } 0 \leq \tau_{1} \leq \tau_{2} \leq T \text {, and for any } \varphi \in C_{c}^{1}\left([0, T] \times R^{3}\right) \text { and any } G \in C(R) \tag{2.6}
\end{equation*}
$$

Remark 2.1. Equation (2.6), introduced in [10], can be viewed as a renormalized formulation of (1.10).

In this paper, a slightly different form of renormalization of (1.10) is needed, namely

$$
\begin{align*}
{\left[\int_{R^{3}} b(\varrho \Theta) \varphi \mathrm{d} x\right]_{t=\tau_{1}}^{t=\tau_{2}} } & =\int_{\tau_{1}}^{\tau_{2}} \int_{R^{3}}\left[b(\varrho \Theta) \partial_{t} \varphi+b(\varrho \Theta) \mathbf{u} \cdot \nabla_{x} \varphi\right] \mathrm{d} x \mathrm{~d} t \\
& +\int_{\tau_{1}}^{\tau_{2}} \int_{R^{3}}\left[b(\varrho \Theta)-b^{\prime}(\varrho \Theta)(\varrho \Theta)\right] \operatorname{div}_{x} \mathbf{u} \mathrm{~d} x \mathrm{~d} t \tag{2.7}
\end{align*}
$$

for any $\varphi \in C_{c}^{\infty}\left([0, T] \times R^{3}\right)$ and any $b \in C^{1}(R), b^{\prime} \in C_{c}(R)$.
In addition, the dissipative solutions are characterized by the energy inequality

$$
\begin{align*}
& {\left[\int_{R^{3}}\left[\frac{1}{2} \varrho|\mathbf{u}|^{2}+\frac{H(\varrho \Theta)-H^{\prime}\left(\varrho_{0}\right)\left(\varrho-\varrho_{0}\right)-H\left(\varrho_{0}\right)}{\varepsilon^{2}}\right] \mathrm{d} x\right]_{t=0}^{t=\tau}} \\
& +\varepsilon^{\alpha} \int_{0}^{\tau} \int_{R^{3}} \mathbb{S}\left(\nabla_{x} \mathbf{u}\right): \nabla_{x} \mathbf{u} \mathrm{~d} x \mathrm{~d} t \leq 0 \tag{2.8}
\end{align*}
$$

for a.a. $\tau \in(0, T)$, where $H(Z)=\frac{1}{\gamma-1} Z^{\gamma}$.
The existence of dissipative weak solution to problem (1.8-1.12) on a bounded domain with the non-slip boundary conditions was shown by Maltese et al. [13].

### 2.2 Smooth solutions of the target problem

Our technique requires the target system to admit a smooth solution. Although there is probably no precise statement concerning solvability of the target system (1.1-1.3) available in the literature, we may anticipate, also in view of the work by Oliver [16] on the anelastic Euler system, the existence of local-in-time smooth solutions in the class:

$$
\begin{align*}
& \mathbf{V} \in C\left(\left[0, T_{\max }\right) ; W^{m, 2}\left(R^{3} ; R^{3}\right)\right)  \tag{2.9}\\
& \Pi \in C\left(\left[0, T_{\max }\right) ; W^{m, 2}\left(R^{3} ; R^{3}\right)\right)  \tag{2.10}\\
& \mathcal{R} \in C\left(\left[0, T_{\max }\right) ; W^{m, 2}\left(R^{3}\right)\right) \tag{2.11}
\end{align*}
$$

for $m>3$, defined on a maximal time interval $\left[0, T_{\max }\right)$.

### 2.3 Main result

In order to state our main result we need a weighted variant of the Helmholtz decomposition,

$$
\mathbf{v}=\mathbf{H}_{\varrho_{0}}[\mathbf{v}]+\nabla_{x} \Phi,
$$

where the potential $\Phi \in D^{1,2}\left(R^{3}\right)$ is uniquely determined by

$$
\operatorname{div}_{x}\left(\varrho_{0} \nabla_{x} \Phi\right)=\operatorname{div}_{x}\left(\varrho_{0} \mathbf{v}\right), \Phi \in D^{1,2}\left(R^{3}\right) \text { for } \mathbf{v} \in L^{2}\left(R^{3}, R^{3}\right)
$$

Our goal is to show the following result.
Theorem 2.2. Let the potential $F$ satisfy (1.5-1.7). Let $\bar{\varrho}>0$ and let $\varrho_{0}$ be the unique solution of (1.4). In addition, suppose

$$
\begin{equation*}
\gamma>\frac{3}{2}, 0<\alpha<\frac{4}{3} . \tag{2.12}
\end{equation*}
$$

Let $\left\{\varrho_{\varepsilon}, \mathbf{u}_{\varepsilon}, \Theta_{\varepsilon}\right\}_{\varepsilon>0}$ be a family of dissipative solutions to the primitive system (1.8-1.12) in $(0, T) \times R^{3}$, with the initial values satisfying (1.13-1.15). Let the target system (1.11.3) admit a strong solution $[\mathbf{V}, \Pi, \mathcal{R}]$ in $(0, T) \times R^{3}$ in the class (2.9-2.11), with the initial values

$$
\mathbf{V}(0, \cdot)=\mathbf{H}_{\varrho_{0}}\left[\mathbf{v}_{0}\right], \mathcal{R}(0, \cdot)=\frac{\varrho_{0}}{\Theta_{0}^{(2)}}>0
$$

Then

$$
\begin{align*}
& \sup _{t \in[0, T]}\left\|\varrho_{\varepsilon}(t, \cdot)-\varrho_{0}\right\|_{\left(L^{2}+L^{\gamma}\right)\left(R^{3}\right)} \rightarrow 0 \\
& \sup _{t \in[0, T]}\left\|\Theta_{\varepsilon}(t, \cdot)-\frac{\varrho_{0}}{\mathcal{R}}(t, \cdot)\right\|_{L^{2}\left(R^{3}\right)} \rightarrow 0,  \tag{2.13}\\
& \int_{0}^{T}\left\|\sqrt{\frac{\varrho_{\varepsilon}}{\varrho_{0}}} \mathbf{u}_{\varepsilon}-\mathbf{V}\right\|_{L^{2}(K)}^{2} \mathrm{~d} t \rightarrow 0 \text { for any compact } K \in R^{3}
\end{align*}
$$

as $\varepsilon \rightarrow 0$.
Remark 2.3. Note that the result is path dependent, meaning the values of the Reynolds and Mach/Froude numbers are interrelated. In particular, the Reynolds number cannot be too large with respect to the Mach number.

The rest of the paper is devoted to the proof of Theorem 2.2.

## 3 Relative energy, uniform bounds

Similarly to [10], we introduce the relative energy functional

$$
\mathcal{E}(\varrho, \Theta, \mathbf{u} \mid r, \mathbf{U})=\int_{R^{3}}\left[\frac{1}{2} \varrho|\mathbf{u}-\mathbf{U}|^{2}+\frac{H(\varrho \Theta)-H^{\prime}(r)(\varrho \Theta-r)-H(r)}{\varepsilon^{2}}\right] \mathrm{d} x
$$

for any pair of "test functions"

$$
\begin{equation*}
(r-\bar{\varrho}) \in C_{c}^{\infty}\left([0, T] \times R^{3}\right), r>0, \quad \mathbf{U} \in C_{c}^{\infty}\left([0, T] \times R^{3} ; R^{3}\right) . \tag{3.1}
\end{equation*}
$$

### 3.1 Relative energy inequality and energy estimates

As shown in [10] (cf. also [9]) any dissipative solution to the primitive system (2.4-2.8) satisfies the relative energy inequality

$$
\begin{align*}
& {[\mathcal{E}(\varrho, \Theta, \mathbf{u} \mid r, \mathbf{U})]_{t=0}^{t=\tau}+\varepsilon^{\alpha} \int_{0}^{\tau} \int_{R^{3}} \mathbb{S}\left(\nabla_{x}(\mathbf{u}-\mathbf{U})\right): \nabla_{x}(\mathbf{u}-\mathbf{U}) \mathrm{d} x \mathrm{~d} t } \\
& \leq \int_{0}^{\tau} \int_{R^{3}} \varrho\left(\partial_{t} \mathbf{U}+\mathbf{u} \cdot \nabla_{x} \mathbf{U}\right) \cdot(\mathbf{U}-\mathbf{u})+\varepsilon^{\alpha} \mathbb{S}\left(\nabla_{x} \mathbf{U}\right): \nabla_{x}(\mathbf{U}-\mathbf{u}) \mathrm{d} x \mathrm{~d} t \\
&+\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}}\left[(r-\varrho \Theta) \partial_{t} H^{\prime}(r)+\nabla_{x} H^{\prime}(r) \cdot(r \mathbf{U}-\varrho \Theta \mathbf{u})\right] \mathrm{d} x \mathrm{~d} t \\
&-\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}}\left[\operatorname{div}_{x} \mathbf{U}\left((\varrho \Theta)^{\gamma}-r^{\gamma}\right)+\varrho \nabla_{x} F \cdot(\mathbf{U}-\mathbf{u})\right] \mathrm{d} x \mathrm{~d} t \tag{3.2}
\end{align*}
$$

for all sufficiently smooth functions belonging to the class (3.1).
Remark 3.1. Here, similarly to ([10]), the class of admissible test functions (3.1) can be considerably extended by means of a density argument.

The relative energy inequality (3.2), together with the renormalized equation (2.6), give rise to a number of uniform bounds for the family $\left\{\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon}\right\}_{\varepsilon>0}$ independent of the scaling parameter. First, let us introduce some notation. Each measurable function $h$ can be decomposed as $h=[h]_{\text {ess }}+[h]_{\text {res }}$, where

$$
[h]_{\mathrm{ess}}=\chi\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right) h,[h]_{\mathrm{res}}=\left(1-\chi\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right) h
$$

where $\chi \in C_{c}^{\infty}(0, \infty), \quad \chi \geq 0, \quad \chi(Y)=1 \quad$ whenever $\frac{1}{2}\left(\min _{x \in R^{3}} \varrho_{0}(x)\right) \leq Y \leq 2\left(\max _{x \in R^{3}} \varrho_{0}(x)\right)$.
Under the hypotheses (1.13-1.15) imposed on the initial data, the following estimates were proved in [10, Section 3.2]:

$$
\begin{align*}
\text { ess } \sup _{t \in[0, T]}\left\|\sqrt{\varrho_{\varepsilon}} \mathbf{u}_{\varepsilon}\right\|_{L^{2}\left(R^{3} ; R^{3}\right)} & \leq c(\text { data }),  \tag{3.3}\\
\varepsilon^{\alpha / 2}\left\|\nabla_{x} \mathbf{u}_{\varepsilon}+\nabla_{x}^{t} \mathbf{u}_{\varepsilon}-\frac{2}{3} \operatorname{div}_{x} \mathbf{u}_{\varepsilon} \mathbb{I}\right\|_{L^{2}\left((0, T) \times R^{3} ; R^{3 \times 3}\right)} & \leq c(\text { data }),  \tag{3.4}\\
\varepsilon^{\alpha / 2} \sqrt{\lambda}\left\|\operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right\|_{L^{2}\left((0, T) \times R^{3}\right)} & \leq c(\text { data }),  \tag{3.5}\\
\operatorname{ess} \sup _{t \in[0, T]}\left(\left\|\frac{\Theta_{\varepsilon}-1}{\varepsilon^{2}}\right\|_{L^{1}\left(R^{3}\right)}+\left\|\frac{\Theta_{\varepsilon}-1}{\varepsilon^{2}}\right\|_{L^{\infty}\left(R^{3}\right)}\right) & \leq c(\text { data })  \tag{3.6}\\
\operatorname{ess} \sup _{t \in[0, T]}\left(\left\|\left[\frac{\varrho_{\varepsilon}-\varrho_{0}}{\varepsilon}\right]_{\operatorname{ess}}\right\|\left\|_{L^{2}\left(R^{3}\right)}+\right\|\left[\frac{\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}}{\varepsilon}\right]_{\mathrm{ess}} \|_{L^{2}\left(R^{3}\right)}\right) & \leq c \text { (data }) \tag{3.7}
\end{align*}
$$

$$
\begin{equation*}
\text { ess } \sup _{t \in[0, T]} \int_{R^{3}}\left([1]_{\mathrm{res}}+\left|\left[\varrho_{\varepsilon}\right]_{\mathrm{res}}\right|^{\gamma}+\left|\left[\varrho_{\varepsilon} \Theta_{\varepsilon}\right]_{\mathrm{res}}\right|^{\gamma}\right) \mathrm{d} x \leq \varepsilon^{2} c(\text { data }) . \tag{3.8}
\end{equation*}
$$

Remark 3.2. In (3.6) and whenever convenient, we set $\Theta_{\varepsilon}=1$ on the vacuum set $\left\{\varrho_{\varepsilon}=0\right\}$.

Finally, we use (3.3), (3.4), (3.7) to conclude

$$
\begin{equation*}
\left.\left\|\mathbf{u}_{\varepsilon}\right\|_{L^{2}\left(0, T ; W^{1,2}\left(R^{3} ; R^{3}\right)\right)} \leq \varepsilon^{-\alpha / 2} c \text { (data }\right) \tag{3.9}
\end{equation*}
$$

### 3.2 Pressure estimates

The pressure estimates are one of the crucial ingredients of the existence theory for the compressible Navier-Stokes system, see Lions [12]. Similarly to Masmoudi [14] (cf. also [9]), they are also needed to control some terms in the anelastic limit. We use the quantities

$$
\begin{aligned}
\varphi(t, x) & =\phi(x) \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}, \phi \in C_{c}^{\infty}\left(R^{3}\right), \phi \geq 0 \\
b \in C^{\infty}[0, \infty), b & \geq 0, b(Y)=a Y^{\beta}, a \geq 0,0 \leq \beta<\gamma \text { for } Y \gg 1
\end{aligned}
$$

as test functions in the momentum equation (2.5). Here, the inverse $\Delta_{x}^{-1}$ is defined in the standard way

$$
\Delta_{x}^{-1}[v](x)=\frac{1}{4 \pi} \int_{R^{3}} \frac{v(y)}{|x-y|} \mathrm{d} y
$$

Note that (3.8) yields

$$
\begin{equation*}
\text { ess } \sup _{t \in(0, T)}\left\|\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right\|_{L^{q}\left(R^{3}\right)} \leq c \varepsilon^{\frac{2}{q}} \text { for any } 1 \leq q \leq \frac{\gamma}{\beta} . \tag{3.10}
\end{equation*}
$$

Following [9, Section 4.2] and denoting

$$
p(Z)=Z^{\gamma}
$$

we deduce the relation

$$
\begin{equation*}
\frac{1}{\varepsilon^{2}} \int_{0}^{T} \int_{R^{3}} \phi\left(p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p\left(\varrho_{0}\right)\right)\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}} \mathrm{~d} x \mathrm{~d} t=\sum_{j=1}^{7} I_{j, \varepsilon} \tag{3.11}
\end{equation*}
$$

with

$$
\begin{aligned}
& I_{1, \varepsilon}=\frac{1}{\varepsilon^{2}} \int_{0}^{T} \int_{R^{3}}\left(p\left(\varrho_{0}\right)-p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right) \nabla_{x} \phi \cdot \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}} \mathrm{~d} x \mathrm{~d} t \\
& I_{2, \varepsilon}=-\frac{1}{\varepsilon^{2}} \int_{0}^{T} \int_{R^{3}} \phi\left(\varrho_{\varepsilon}-\varrho_{0}\right) \nabla_{x} F \cdot \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}} \mathrm{~d} x \mathrm{~d} t \\
& I_{3, \varepsilon}=\varepsilon^{\alpha} \int_{0}^{T} \int_{R^{3}} \mathbb{S}\left(\nabla_{x} \mathbf{u}_{\varepsilon}\right): \nabla_{x}\left(\phi \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right) \mathrm{d} x \mathrm{~d} t \\
& I_{4, \varepsilon}=-\int_{0}^{T} \int_{R^{3}} \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \otimes \mathbf{u}_{\varepsilon}: \nabla_{x}\left(\phi \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right) \mathrm{d} x \mathrm{~d} t \\
& I_{5, \varepsilon}=\int_{0}^{T} \int_{R^{3}} \phi \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \cdot \nabla_{x} \Delta_{x}^{-1}\left[\operatorname{div}_{x}\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}} \mathbf{u}_{\varepsilon}\right)\right] \mathrm{d} x \mathrm{~d} t \\
& I_{6, \varepsilon}=\int_{0}^{T} \int_{R^{3}} \phi \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \cdot \nabla_{x} \Delta_{x}^{-1}\left[\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right) \mathrm{div}_{x} \mathbf{u}_{\varepsilon}\right] \mathrm{d} x \mathrm{~d} t \\
& I_{7, \varepsilon}=\left[\int_{\Omega} \phi \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \cdot \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}(\tau, \cdot) \mathrm{d} x\right]_{\tau=0}^{\tau=T}
\end{aligned}
$$

where we have also used the renormalized equation (2.7).
We consider two complementary choices of $b$, namely

$$
\begin{equation*}
b(Y)=0 \text { if } 0 \leq Y \leq \min _{x \in R^{3}} \varrho_{0}, \text { or } b(Y)=0 \text { if } Y \geq \max _{x \in R^{3}} \varrho_{0} \tag{3.12}
\end{equation*}
$$

Note that in both cases the integrand on the left-hand side of (3.11) has a definite sign. Now we proceed in several steps.

### 3.2.1 Integrals $I_{1}-I_{3}, I_{7}$

We focus on the (more difficult) former case in (3.12) obtaining

$$
\begin{aligned}
& \left|I_{1, \varepsilon}\right| \leq \frac{c}{\varepsilon^{2}} \int_{0}^{T} \int_{R^{3}}\left|\left[p\left(\varrho_{0}\right)-p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{ess}}\right|\left|\nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right|\left|\nabla_{x} \phi\right| \mathrm{d} x \mathrm{~d} t \\
& \quad+\frac{c}{\varepsilon^{2}} \int_{0}^{T} \int_{R^{3}}\left|\left[p\left(\varrho_{0}\right)-p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right|\left|\nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right|\left|\nabla_{x} \phi\right| \mathrm{d} x \mathrm{~d} t
\end{aligned}
$$

By virtue of the Sobolev embedding relations,

$$
\begin{aligned}
\left\|\nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right\|_{L^{2}\left(R^{3} ; R^{3}\right)} & \leq c\left\|\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right\|_{L^{6 / 5}\left(R^{3}\right)}, \\
\left\|\nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right\|_{L^{\infty}\left(R^{3} ; R^{3}\right)} & \leq\left\|\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right\|_{\left(L^{1} \cap L^{q}\right)\left(R^{3}\right)}, \text { for } q>3 .
\end{aligned}
$$

Thus for $0<\beta<\frac{\gamma}{3}$, we may use (3.10), together with the uniform bounds (3.7), (3.8), to conclude that

$$
\left|I_{1, \varepsilon}\right| \leq \varepsilon^{\omega} c \text { (data) for some } \omega>0
$$

Furthermore, in view of (3.4), (3.7), and (3.8), a similar argument can be applied to $I_{2, \varepsilon}, I_{3, \varepsilon}$, and $I_{7, \varepsilon}$ to obtain

$$
\left|I_{j, \varepsilon}\right| \leq \varepsilon^{\omega} c \text { (data) for some } \omega>0, j=2,3,7
$$

### 3.2.2 Integrals $I_{4}, I_{5}$

To handle $I_{4}$ we have to use the "negative estimate" (3.9) along with the embedding $W^{1,2}\left(R^{3}\right) \hookrightarrow L^{6}\left(R^{3}\right)$. Consequently, we have

$$
\begin{aligned}
I_{4, \varepsilon} & =-\int_{0}^{T} \int_{R^{3}}\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{ess}} \otimes \mathbf{u}_{\varepsilon}: \nabla_{x}\left(\phi \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right) \mathrm{d} x \mathrm{~d} t \\
& -\int_{0}^{T} \int_{R^{3}}\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{res}} \otimes \mathbf{u}_{\varepsilon}: \nabla_{x}\left(\phi \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right) \mathrm{d} x \mathrm{~d} t
\end{aligned}
$$

where, furthermore,

$$
\begin{aligned}
& \left|\int_{R^{3}}\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{ess}} \otimes \mathbf{u}_{\varepsilon}: \nabla_{x}\left(\phi \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right) \mathrm{d} x\right| \\
& \leq c\left\|\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{ess}}\right\|_{L^{2}\left(R^{3} ; R^{3}\right)}\left\|\mathbf{u}_{\varepsilon}\right\|_{L^{6}\left(R^{3} ; R^{3}\right)}\left\|\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right\|_{L^{3}\left(R^{3}\right)},
\end{aligned}
$$

and where, by virtue of (3.3), (3.6), (3.9), and (3.10), the right-hand side is of order $\varepsilon^{\omega}$, $\omega>0$ as $\varepsilon^{2 / 3} \varepsilon^{-\alpha / 2}=\varepsilon^{\omega}, \omega=\frac{2}{3}-\frac{\alpha}{2}>0$ in accordance with hypothesis (2.12).

As for the residual part, we get

$$
\begin{aligned}
& \left|\int_{0}^{T} \int_{R^{3}}\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{res}} \otimes \mathbf{u}_{\varepsilon}: \nabla_{x}\left(\phi \nabla_{x} \Delta_{x}^{-1}\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right) \mathrm{d} x \mathrm{~d} t\right| \\
& \leq\left\|\left[\sqrt{\varrho_{\varepsilon}}\right]_{\mathrm{res}}\right\|_{L^{2 \gamma}\left(R^{3}\right)}\left\|\sqrt{\varrho_{\varepsilon}} \mathbf{u}_{\varepsilon}\right\|_{L^{2}\left(R^{3} ; R^{3}\right)}\left\|\mathbf{u}_{\varepsilon}\right\|_{L^{6}\left(R^{3} ; R^{3}\right)}\left\|\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}\right\|_{L^{q}\left(R^{3}\right)}, q=\frac{6 \gamma}{2 \gamma-3}
\end{aligned}
$$

Thus, in accordance with the bounds (3.8), (3.9), and (3.10), we need

$$
\frac{2 \gamma-3}{3 \gamma}+\frac{1}{\gamma}>\frac{\alpha}{2}, \text { meaning, } 0<\alpha<\frac{4}{3}
$$

exactly as required by (2.12).
The integral $I_{5, \varepsilon}$ can be treated in the same manner.

### 3.2.3 Integral $I_{6}$

Finally,

$$
\begin{aligned}
\left|I_{6, \varepsilon}\right| & \leq\left|\int_{0}^{T} \int_{R^{3}} \phi\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{ess}} \cdot \nabla_{x} \Delta_{x}^{-1}\left[\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right) \operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right] \mathrm{d} x \mathrm{~d} t\right| \\
& +\left|\int_{0}^{T} \int_{R^{3}} \phi\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{res}} \cdot \nabla_{x} \Delta_{x}^{-1}\left[\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right) \operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right] \mathrm{d} x \mathrm{~d} t\right|,
\end{aligned}
$$

where, by means of the Sobolev embedding $W^{1,6 / 5} \hookrightarrow L^{2}$, we get

$$
\begin{aligned}
& \left|\int_{0}^{T} \int_{R^{3}} \phi\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{ess}} \cdot \nabla_{x} \Delta_{x}^{-1}\left[\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right) \operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right] \mathrm{d} x \mathrm{~d} t\right| \\
& \leq\left\|\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{ess}}\right\|_{L^{2}\left(R^{3} ; R^{3}\right)}\left\|\nabla_{x} \mathbf{u}\right\|_{L^{2}\left(R^{3} ; R^{3 \times 3}\right)}\left\|\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right\|_{L^{3}\left(R^{3}\right)}
\end{aligned}
$$

where the right-hand side can be estimated exactly is its counterpart in $I_{4, \varepsilon}$.
To conclude

$$
\begin{aligned}
& \left|\int_{0}^{T} \int_{R^{3}} \phi\left[\varrho_{\varepsilon} \mathbf{u}_{\varepsilon}\right]_{\mathrm{res}} \cdot \nabla_{x} \Delta_{x}^{-1}\left[\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right) \operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right] \mathrm{d} x \mathrm{~d} t\right| \\
& \leq\left\|\left[\sqrt{\varrho_{\varepsilon}}\right]_{\mathrm{res}}\right\|_{L^{2 \gamma}\left(R^{3}\right)}\left\|\sqrt{\varrho_{\varepsilon}} \mathbf{u}_{\varepsilon}\right\|_{L^{2}\left(R^{3} ; R^{3}\right)} \times \\
& \times\left\|\nabla_{x} \Delta_{x}^{-1}\left[\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right) \operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right]\right\|_{L^{q}\left(R^{3} ; R^{3}\right)}, q=\frac{2 \gamma}{\gamma-1},
\end{aligned}
$$

where, furthermore,

$$
\begin{aligned}
& \left\|\nabla_{x} \Delta_{x}^{-1}\left[\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right) \operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right]\right\|_{L^{2 \gamma /(\gamma-1)\left(R^{3} ; R^{3}\right)}} \\
& \leq\left\|\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right) \operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right\|_{L^{6 \gamma /(5 \gamma-3)\left(R^{3}\right)}}
\end{aligned}
$$

and, by Hölder's inequality,

$$
\begin{aligned}
& \left\|\left(\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right) \operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right\|_{L^{6 \gamma /(5 \gamma-3)}\left(R^{3}\right)} \\
& \leq\left\|\operatorname{div}_{x} \mathbf{u}_{\varepsilon}\right\|_{L^{2}\left(R^{3} ; R^{3}\right)}\left\|\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}-\left[b\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)\right]_{\mathrm{res}}^{\prime} \varrho_{\varepsilon} \Theta_{\varepsilon}\right\|_{L^{6 \gamma /(2 \gamma-3)}\left(R^{3}\right)}
\end{aligned}
$$

whence the required bound follows again from hypothesis (2.12).

### 3.2.4 Pressure estimates - conclusion

We conclude that, under hypothesis (2.12),
$\int_{0}^{T} \int_{K}\left(\left[\varrho_{\varepsilon} \Theta_{\varepsilon}\right]_{\mathrm{res}}\right)^{\gamma+\beta} \mathrm{d} x \mathrm{~d} t \leq c(K) \varepsilon^{2+\omega}$ for certain $\beta, \omega>0$ and any compact $K \subset R^{3}$.

Remark 3.3. Note that estimate (3.13) is only local in space. It is only at this moment of the proof, where we effectively use the restriction

$$
0<\alpha<\frac{4}{3}
$$

imposed by (2.12).

## 4 Acoustic waves

The initial data can be decomposed as

$$
\begin{aligned}
\varrho(0, \cdot) & =\varrho_{0}+\varepsilon\left(\varrho_{0, \varepsilon}^{(1)}-\varrho_{0}^{(1)}\right)+\varepsilon \varrho_{0}^{(1)} \\
\mathbf{u}(0, \varepsilon) & =\mathbf{u}_{0, \varepsilon}-\mathbf{v}_{0}+\mathbf{H}_{\varrho 0}\left[\mathbf{v}_{0}\right]+\mathbf{v}_{0}-\mathbf{H}_{\varrho 0}\left[\mathbf{v}_{0}\right]=\mathbf{u}_{0, \varepsilon}-\mathbf{v}_{0}+\mathbf{V}_{0}+\nabla_{x} \Phi_{0}
\end{aligned}
$$

where, by virtue of (1.13), (1.14), the terms $\varrho_{0}$ and $\mathbf{V}_{0}$ represent the zero-th order approximation, the terms $\left(\varrho_{0, \varepsilon}^{(1)}-\varrho_{0}^{(1)}\right)$ and $\mathbf{u}_{0, \varepsilon}-\mathbf{v}_{0}$ vanish in the asymptotic limit, and

$$
\varrho_{0}^{(1)}, \nabla_{x} \Phi_{0}, \text { where } \operatorname{div}_{x}\left(\varrho_{0} \nabla_{x} \Phi_{0}\right)=\operatorname{div}_{x}\left(\varrho_{0} \mathbf{v}_{0}\right)
$$

create the acoustic waves governed by the system of equations

$$
\begin{align*}
\varepsilon \partial_{t} s_{\varepsilon}+\operatorname{div}_{x}\left[\varrho_{0} \nabla_{x} \Phi_{\varepsilon}\right] & =0  \tag{4.1}\\
\varepsilon \varrho_{0} \partial_{t} \nabla_{x} \Phi_{\varepsilon}+\varrho_{0} \nabla_{x}\left[\frac{p^{\prime}\left(\varrho_{0}\right)}{\varrho_{0}} s_{\varepsilon}\right] & =0 \tag{4.2}
\end{align*}
$$

with the initial data

$$
\begin{equation*}
s_{\varepsilon}(0, \cdot)=\varrho_{0}^{(1)}, \nabla_{x} \Phi_{\varepsilon}(0, \cdot)=\nabla_{x} \Phi_{0} \tag{4.3}
\end{equation*}
$$

Note that (4.1-4.3) is exactly the same system as obtained in [9], where the analysis has been performed under a highly simplifying assumption of $F$ being compactly supported. In such a case, the density profile $\varrho_{0}$ given by (1.4) coincides with the constant $\bar{\varrho}$ outside a bounded ball and the problem can be treated as a compact perturbation of the standard wave equation generated by "flat" Laplacian. In the present setting, $\varrho_{0}$ shares the asymptotic properties of the long range potential $F$ specified in (1.5-1.7), more precisely

$$
\begin{equation*}
\varrho_{0}(x)=Q^{-1}(F(x)+Q(\bar{\varrho})), \text { where } Q^{\prime}(r)=\frac{p^{\prime}(r)}{r}=\gamma r^{\gamma-2} . \tag{4.4}
\end{equation*}
$$

We introduce the differential operator

$$
\mathcal{A}_{\varrho_{0}}: v \mapsto-\frac{p^{\prime}\left(\varrho_{0}\right)}{\varrho_{0}} \operatorname{div}_{x}\left(\varrho_{0} \nabla_{x} v\right)
$$

that can be interpreted as a non-negative, self-adjoint operator on the weighted space $L^{2}\left(R^{3}\right)$ space endowed with the scalar product

$$
\langle u ; v\rangle=\int_{R^{3}} u v \frac{\varrho_{0}}{p^{\prime}\left(\varrho_{0}\right)} \mathrm{d} x
$$

see DeBièvre and Pravica [4]. Moreover, as $\varrho_{0}$ is given by (4.4), where $F$ satisfies (1.5-1.7), we may infer that

- the point spectrum of $\mathcal{A}_{\varrho_{0}}$ is empty, see DeBièvre and Pravica [5, Theorem 2.1(a)];
- the operator $\mathcal{A}_{\varrho_{0}}$ satisfies the Limiting absorption principle, see DeBièvre and Pravica [4, Theorem 1.1].
Finally, we may use the same arguments as in [7, Section 6.3] (see also [6]), in particular the abstract theorem of Kato [11], to show the frequency localized decay estimates:

$$
\begin{equation*}
\int_{-\infty}^{\infty}\left\|\varphi G\left(\mathcal{A}_{\varrho_{0}}\right) \exp \left( \pm \mathrm{i} \sqrt{\mathcal{A}_{\varrho_{0}}} t\right)[h]\right\|_{L^{2}\left(R^{3}\right)}^{2} \mathrm{~d} t \leq c(\varphi, G)\|h\|_{L^{2}\left(R^{3}\right)}^{2} \tag{4.5}
\end{equation*}
$$

for any $\varphi \in C_{c}^{\infty}\left(R^{3}\right)$ and any $G \in C_{c}^{\infty}(0, \infty)$. Note that $\varphi$ represents a cut-off in the physical space while $G$ a cut-off in the frequency space. Such an estimate is too weak to be used in combination with the relative energy inequality requiring stronger estimates of Strichartz's type, cf. [9]. They will be derived in the remaining part of the section.

### 4.1 Strichartz estimates by Metcalfe and Tataru

Metcalfe and Tataru [15] proved global Strichartz estimates for solutions of the wave equation

$$
\begin{equation*}
\partial_{t, t}^{2} V-\operatorname{div}_{x}\left(\tilde{A}(x) \nabla_{x} V\right)+\tilde{\mathbf{B}}(x) \cdot \nabla_{x} V=Z, V(0, \cdot)=V_{0}, \partial_{t} V(0, \cdot)=V_{1} \tag{4.6}
\end{equation*}
$$

in the form

$$
\begin{equation*}
\|V\|_{L^{p}\left(-\infty, \infty ; L^{q}\left(R^{3}\right)\right)} \leq\left(\left\|\nabla_{x} V_{0}\right\|_{L^{2}\left(R^{3} ; R^{3}\right)}+\left\|V_{1}\right\|_{L^{2}\left(R^{3}\right)}+\|Z\|_{L^{r}\left(-\infty, \infty ; L^{s}\left(R^{3}\right)\right)}\right), \tag{4.7}
\end{equation*}
$$

whenever

$$
\begin{equation*}
\frac{1}{p}+\frac{3}{q}=\frac{1}{2}=\frac{1}{r}+\frac{3}{s}-2 . \tag{4.8}
\end{equation*}
$$

The coefficients $\tilde{A}(x), \tilde{\mathbf{B}}(x)$ must be "asymptotically flat" in the sense that

$$
\begin{align*}
\sum_{j \in Z} \sup _{x \in A_{j}}\left(|x|^{2}\left|\nabla_{x}^{2} \tilde{A}(x)\right|+|x||\tilde{A}(x)|+|\tilde{A}(x)-\bar{A}|\right) & \leq \delta, \bar{A}>0  \tag{4.9}\\
\sum_{j \in Z} \sup _{x \in A_{j}}\left(|x|^{2}\left|\nabla_{x} \tilde{\mathbf{B}}(x)\right|+|x||\tilde{\mathbf{B}}(x)|\right) & \leq \delta \tag{4.10}
\end{align*}
$$

where $A_{j}=\left\{2^{j} \leq|x| \leq 2^{j+1}\right\}$ are the dyadic regions covering $R^{3}$, and $\delta>0$ is sufficiently small, see Metcalfe and Tataru [15, Theorem 2].

Now, we rewrite the operator $\mathcal{A}_{\varrho_{0}}$ in the form used by Metcalfe and Tataru, specifically,

$$
\begin{aligned}
\mathcal{A}_{\varrho_{0}}[v] & =-\operatorname{div}_{x}\left(p^{\prime}\left(\varrho_{0}\right) \nabla_{x} v\right)+\varrho_{0} \nabla_{x}\left(\frac{p^{\prime}\left(\varrho_{0}\right)}{\varrho_{0}}\right) \cdot \nabla_{x} v \\
& =-\operatorname{div}_{x}\left(p^{\prime}\left(\varrho_{0}\right) \nabla_{x} v\right)+\varrho_{0} Q^{\prime \prime}\left(\varrho_{0}\right) \nabla_{x} \varrho_{0} \cdot \nabla_{x} v
\end{aligned}
$$

and set

$$
\begin{equation*}
A(x)=p^{\prime}\left(\varrho_{0}\right)(x), \mathbf{B}(x)=\varrho_{0} Q^{\prime \prime}\left(\varrho_{0}\right) \nabla_{x} \varrho_{0} \tag{4.11}
\end{equation*}
$$

We easily compute that

$$
\begin{equation*}
\left|\nabla_{x} A(x)\right|+|\mathbf{B}(x)| \leq c\left|\nabla_{x} F(x)\right| \leq \frac{c}{\left(1+|x|^{2}\right)} \tag{4.12}
\end{equation*}
$$

and

$$
\begin{equation*}
\left|\nabla_{x}^{2} A(x)\right|+\left|\nabla_{x} \mathbf{B}(x)\right| \leq c\left(\left|\nabla_{x}^{2} F(x)\right|+\left|\nabla_{x} F(x)\right|^{2}\right) \leq \frac{c}{\left(1+|x|^{3}\right)} \tag{4.13}
\end{equation*}
$$

Of course, the result of Metcalfe and Tataru [15] may not be applicable to the operator $\mathcal{A}_{\varrho_{0}}$ as $A$ and $\mathbf{B}$ may fail to comply with (4.9), (4.10) for small $\delta$. Instead, we consider a modified density profile $\check{\varrho}$,

$$
\begin{equation*}
\tilde{\varrho}(x)=Q^{-1}\left(H_{R}(F(x))+Q(\bar{\varrho})\right), \tag{4.14}
\end{equation*}
$$

where

$$
\begin{align*}
H_{R} \in C^{\infty}(R), H_{R}(z) & =z \text { on an open interval containing }(-\delta, \delta) \\
\left|H_{R}^{\prime}(z)\right| & \leq 1 \text { for all } z,\left|H_{R}^{\prime \prime}(z)\right| \leq \frac{c}{R}, H_{R}(z)=0 \text { for all } z \geq R \tag{4.15}
\end{align*}
$$

with a constant $c$ independent of $R>0$.
Next, we introduce a wave operator $\mathcal{A}_{\tilde{\varrho}}$,

$$
\mathcal{A}_{\tilde{\varrho}}[v]=\operatorname{div}_{x}\left(p^{\prime}(\tilde{\varrho}) \nabla_{x} v\right)+\tilde{\varrho} Q^{\prime \prime}(\tilde{\varrho}) \nabla_{x} \tilde{\varrho} \cdot \nabla_{x} v,
$$

with

$$
\tilde{A}(x)=p^{\prime}(\tilde{\varrho})(x), \tilde{\mathbf{B}}(x)=\tilde{\varrho} Q^{\prime \prime}(\tilde{\varrho}) \nabla_{x} \tilde{\varrho} .
$$

Observe that, since $H_{R}$ is linear on a neighborhood of zero and $F(x) \rightarrow 0$ as $|x| \rightarrow \infty$,

$$
\begin{equation*}
\mathcal{A}_{\tilde{\varrho}}=\mathcal{A}_{\varrho_{0}} \text { on the set }|x|>M, M=M(R) \text { large enough. } \tag{4.16}
\end{equation*}
$$

Our goal is to show that for any $\delta>0$, there is $R>0$ sufficiently small such that $\tilde{A}, \tilde{\mathbf{B}}$ satisfy (4.9), (4.10), in other words, we have the Strichartz estimates (4.7) for equation (4.6). Indeed, as $F>0$, we have that for any $L>0$, there is $R=R(L)$ such that

$$
\begin{equation*}
\tilde{A}(x)=p^{\prime}(\bar{\varrho}) \text { for all }|x|<L,\left|\tilde{A}(x)-p^{\prime}(\bar{\varrho})\right| \leq H_{R}(F(x)) \leq \frac{1}{|x|} \text { for }|x| \geq L . \tag{4.17}
\end{equation*}
$$

By the same token, using (4.12), we get

$$
\begin{align*}
\nabla_{x} \tilde{A}, \tilde{\mathbf{B}} & =0 \text { for all }|x|<L \\
\left|\nabla_{x} \tilde{A}(x)\right|+|\tilde{\mathbf{B}}(x)| & \leq c\left|\nabla_{x} H_{R}(F)(x)\right| \leq c\left|\nabla_{x} F\right| \leq \frac{c}{|x|^{2}} \text { for }|x| \geq L \tag{4.18}
\end{align*}
$$

Finally, by virtue of (4.13), we also have

$$
\begin{align*}
\nabla_{x}^{2} \tilde{A}, \nabla_{x} \tilde{\mathbf{B}} & =0 \text { for all }|x|<L, \\
\left|\nabla_{x}^{2} \tilde{A}(x)\right|+\left|\nabla_{x} \tilde{\mathbf{B}}(x)\right| & \leq c\left(\left|\nabla_{x}^{2} H_{R}(F)(x)\right|+\left|\nabla_{x} H_{R}(F)(x)\right|^{2}\right) \\
& \leq c\left[\left(1+H_{R}^{\prime \prime}(F(x))\right)\left|\nabla_{x} F(x)\right|^{2}+\left|\nabla_{x}^{2} F(x)\right|^{2}\right]  \tag{4.19}\\
& \leq c\left[\frac{1}{|x|^{3}}+H_{R}^{\prime \prime}(F(x)) \frac{1}{|x|^{4}}\right] \text { for }|x| \geq L .
\end{align*}
$$

On the other hand, however,

$$
H_{R}^{\prime \prime}(F) \neq 0 \Rightarrow 0 \leq F(x) \leq R
$$

and, as $F(x) \approx \frac{1}{|x|}$ for $|x| \rightarrow \infty$ and by virtue of (4.15),

$$
\left|H_{R}^{\prime \prime}(F)(x)\right| \leq \frac{c}{R} \leq c|x| ;
$$

whence (4.19) yields

$$
\begin{align*}
& \nabla_{x} \tilde{A}, \nabla_{x}^{2} \tilde{A}, \quad \nabla_{x} \tilde{\mathbf{B}}=0 \text { for all }|x|<L, \\
&\left|\nabla_{x}^{2} \tilde{A}(x)\right|+\left|\nabla_{x} \tilde{\mathbf{B}}(x)\right| \leq \frac{c}{|x|^{3}} \text { for }|x| \geq L . \tag{4.20}
\end{align*}
$$

Thus if $R=R(\delta)>0$ is taken small enough, the coefficients $\tilde{A}, \tilde{\mathbf{B}}$ satisfy (4.9), (4.10) and the result of Metcalfe and Tataru [15] applies yielding (4.6-4.8).

### 4.2 Decomposition method of Smith and Sogge

Our final goal is to combine the localized decay estimates (4.5) with the global ones established in the preceding section to derive frequency localized Strichartz estimates for the operator $\mathcal{A}_{\varrho_{0}}$ in the form

$$
\begin{equation*}
\int_{-\infty}^{\infty}\left\|G\left(\mathcal{A}_{\varrho_{0}}\right) \exp \left( \pm \mathrm{i} \sqrt{\mathcal{A}_{\varrho_{0}}} t\right)[h]\right\|_{L^{q}\left(R^{3}\right)}^{p} \mathrm{~d} t \leq c(G)\|h\|_{L^{2}\left(R^{3}\right)}^{p} \tag{4.21}
\end{equation*}
$$

for any any $G \in C_{c}^{\infty}(0, \infty)$ provided

$$
\frac{1}{p}+\frac{3}{q}=\frac{1}{2}
$$

Remark 4.1. The main difference between (4.5) and (4.21) is the absence of the spatial cut-off in the latter. As the functions $G\left(-\mathcal{A}_{\varrho_{0}}\right) \exp \left( \pm \mathrm{i} \sqrt{\mathcal{A}_{\varrho_{0}}} t\right)[h]$ are smooth, estimate (4.21) yields uniform decay in $x$ of spatial derivatives of arbitrary order.

To deduce (4.21), we use the decomposition method by Smith and Sogge [17]. We consider

$$
U=G\left(\mathcal{A}_{\varrho_{0}}\right) \exp \left( \pm \mathrm{i} \sqrt{\mathcal{A}_{\varrho_{0}}} t\right)[h]
$$

- the solution of the problem

$$
\begin{aligned}
& \partial_{t, t}^{2} U-\operatorname{div}_{x}\left(A(x) \nabla_{x} U\right)+\mathbf{B}(x) \cdot \nabla_{x} U=0, \\
& U(0, \cdot)=G\left(\mathcal{A}_{\varrho_{0}}\right)[h], \partial_{t} U(0, \cdot)= \pm \mathrm{i} \sqrt{\mathcal{A}_{\varrho_{0}}} G\left(-\mathcal{A}_{\varrho_{0}}\right)[h],
\end{aligned}
$$

where $A$ and $\mathbf{B}$ have been introduced in (4.11).
Consider a cut-off function

$$
\chi \in C_{c}^{\infty}(R), \chi(z)=\left\{\begin{array}{l}
1 \text { for }|z|<M+1 \\
0 \text { for }|z|>M+2
\end{array}\right.
$$

where $M$ is from (4.16). Now, we decompose

$$
U=\chi(|x|) U+(1-\chi(|x|)) U,
$$

where, by virtue of (4.16), the function $V=(1-\chi(|x|)) U$ satisfies

$$
\begin{align*}
\partial_{t, t}^{2} V-\operatorname{div}_{x}\left(\tilde{A}(x) \nabla_{x} V\right) & +\tilde{\mathbf{B}}(x) \cdot \nabla_{x} V=Z \\
Z & =\tilde{A} \nabla_{x} \chi \cdot \nabla_{x} U+\operatorname{div}_{x}\left(\tilde{A} \nabla_{x} \chi U\right)-\tilde{\mathbf{B}} \cdot \nabla_{x} \chi U  \tag{4.22}\\
V(0, \cdot) & =(1-\chi) U_{0}, \partial_{t} V(0, \cdot)=(1-\chi) U_{1}
\end{align*}
$$

Now, as $U$ satisfies the local estimates (4.5), the desired decay properties for $\chi U$ follow. As for $V$, we can apply the Strichartz estimates (4.7) for $r=2$, $s=\frac{3}{2}$, estimating $Z$ by means of (4.5). Note that all components of $Z$ have compact support in $R^{3}$ and $U$ as well as $\tilde{A}, \tilde{\mathbf{B}}$, and $\chi$ are smooth. Thus we have proved (4.21).

### 4.3 Dispersive estimates for the rescaled equation

Going back to the original rescaled system (4.1-4.3) we first regularize the initial data in the way similar to [8], [9] taking

$$
\begin{equation*}
s_{\varepsilon}(0, \cdot)=s_{0, \delta}=\frac{\varrho_{0}}{p^{\prime}\left(\varrho_{0}\right)}\left[\frac{p^{\prime}\left(\varrho_{0}\right)}{\varrho_{0}} \varrho_{0}^{(1)}\right]_{\delta}, \Phi_{\varepsilon}(0 \cdot)=\Phi_{0, \delta}=\left[\Phi_{0}\right]_{\delta} \tag{4.23}
\end{equation*}
$$

where the regularization operator $[h]_{\delta}$ is defined as

$$
\begin{gathered}
{[h]_{\delta}=G_{\delta}\left(\sqrt{\mathcal{A}_{\varrho_{0}}}\right)\left[\psi_{\delta} h\right]} \\
\psi_{\delta} \in C_{c}^{\infty}\left(R^{3}\right), 0 \leq \psi_{\delta} \leq 1, \psi_{\delta}(x)=1 \text { for }|x|<\frac{1}{\delta}, \psi_{\delta}(x)=0 \text { for }|x|>\frac{2}{\delta},
\end{gathered}
$$

$$
\begin{aligned}
G_{\delta} & \in C_{c}^{\infty}(R), 0 \leq G_{\delta} \leq 1, G_{\delta}(-z)=G_{\delta}(z) \\
G_{\delta}(z) & =1 \text { for } z \in\left(-\frac{1}{\delta},-\delta\right) \cup\left(\delta, \frac{1}{\delta}\right) \\
G_{\delta}(z) & =0 \text { for } z \in\left(-\infty,-\frac{2}{\delta}\right) \cup\left(-\frac{\delta}{2}, \frac{\delta}{2}\right) \cup\left(\frac{2}{\delta}, \infty\right) .
\end{aligned}
$$

Finally, exactly as in [9], we may combine the frequency localized estimate (4.21), together with the standard energy estimates for the rescaled system (4.1-4.3), to conclude that

$$
\begin{align*}
& \sup _{t \in[0, T]}\left(\left\|\Phi_{\varepsilon}(t, \cdot)\right\|_{W^{m, 2}\left(R^{3}\right)}+\left\|s_{\varepsilon}(t, \cdot)\right\|_{W^{m, 2}\left(R^{3}\right)}\right) \\
& \quad \leq c_{E}(m, \delta)\left(\left\|\Phi_{0, \delta}\right\|_{L^{2}\left(R^{3}\right)}+\left\|s_{0, \delta}\right\|_{L^{2}\left(R^{3}\right)}\right) \\
& \quad \int_{0}^{T}\left(\left\|\Phi_{\varepsilon}(t, \cdot)\right\|_{W^{m, \infty}\left(R^{3}\right)}+\left\|s_{\varepsilon}(t, \cdot)\right\|_{W^{m, \infty}\left(R^{3}\right)}\right) \mathrm{d} t  \tag{4.24}\\
& \quad \leq c_{D}(\varepsilon, m, \delta)\left(\left\|\nabla_{x} \Phi_{0, \delta}\right\|_{L^{2}\left(R^{3} ; R^{3}\right)}+\left\|s_{0, \delta}\right\|_{L^{2}\left(R^{3}\right)}\right),
\end{align*}
$$

for any $m$ and $\delta>0$, where

$$
c_{D}(\varepsilon, m, \delta) \rightarrow 0 \text { as } \varepsilon \rightarrow 0 \text { for } m, \delta \text { fixed. }
$$

## 5 Conclusion, proof of the main result

The proof of Theorem 2.2 will be concluded by taking the limit solution in place of the test functions in the relative energy inequality (3.2). To this end, it is convenient to rewrite (1.1-1.3) in terms of the "temperature"

$$
\mathcal{T}=\frac{\varrho_{0}}{\mathcal{R}}
$$

namely

$$
\begin{align*}
\operatorname{div}_{x}\left(\varrho_{0} \mathbf{V}\right) & =0  \tag{5.1}\\
\partial_{t} \mathbf{V}+\mathbf{V} \cdot \nabla_{x} \mathbf{V}+\nabla_{x} \Pi & =-\mathcal{T} \nabla_{x} F  \tag{5.2}\\
\partial_{t} \mathcal{T}+\mathbf{V} \cdot \nabla_{x} \mathcal{T} & =0 \tag{5.3}
\end{align*}
$$

with the initial data

$$
\begin{equation*}
\mathbf{V}(0, \cdot)=\mathbf{H}_{\varrho_{0}}\left[\mathbf{v}_{0}\right], \mathcal{T}(0, \cdot)=\Theta_{0}^{(2)} \tag{5.4}
\end{equation*}
$$

### 5.1 Relative energy ansatz

We consider

$$
r=r_{\varepsilon}=\varrho_{0}+\varepsilon s_{\varepsilon}, \mathbf{U}=\mathbf{U}_{\varepsilon}=\mathbf{V}+\nabla_{x} \Phi_{\varepsilon}
$$

where $\left[s_{\varepsilon}, \Phi_{\varepsilon}\right]$ solve the acoustic wave equation (4.1), (4.2), and $\mathbf{V}$ is the velocity in the limit problem (5.1-5.4), as test functions in the relative energy inequality (3.2). Note that the initial data can be replaced by their regularizations (4.23) as

$$
s_{0, \delta} \rightarrow \varrho_{0}^{(1)} \text { in } L^{2}\left(R^{3}\right), \nabla_{x} \Phi_{0, \delta} \rightarrow \nabla_{x} \Phi_{0} \text { in } L^{2}\left(R^{3} ; R^{3}\right) \text { as } \delta \rightarrow 0
$$

With $\delta>0$ fixed, our ultimate goal will be to show that all terms on the right-hand side of the relative energy inequality (3.2) can be either "absorbed" by the left-hand side by means of a Gronwall type argument or vanish in the asymptotic limit for $\varepsilon \rightarrow 0$. This will be done in several steps.

### 5.1.1 Step 1 - rewriting the pressure terms

To begin, observe that

$$
\partial_{t} r_{\varepsilon}+\operatorname{div}_{x}\left(r_{\varepsilon} \mathbf{U}_{\varepsilon}\right)=\varepsilon \partial_{t} s_{\varepsilon}+\operatorname{div}_{x}\left(r_{\varepsilon} \mathbf{V}+r_{\varepsilon} \nabla_{x} \Phi_{\varepsilon}\right)=\varepsilon \operatorname{div}_{x}\left(s_{\varepsilon} \mathbf{U}_{\varepsilon}\right)
$$

Using this fact we get, after a straightforward manipulation,

$$
\begin{aligned}
& \left(r_{\varepsilon}-\varrho_{\varepsilon} \Theta_{\varepsilon}\right) \partial_{t} H^{\prime}\left(r_{\varepsilon}\right)+\nabla_{x} H^{\prime}\left(r_{\varepsilon}\right) \cdot\left(r_{\varepsilon} \mathbf{U}_{\varepsilon}-\varrho_{\varepsilon} \Theta_{\varepsilon} \mathbf{u}_{\varepsilon}\right)-\operatorname{div}_{x} \mathbf{U}_{\varepsilon}\left(\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)^{\gamma}-r_{\varepsilon}^{\gamma}\right) \\
& =-\operatorname{div}_{x} \mathbf{U}_{\varepsilon}\left[p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(r_{\varepsilon}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-r_{\varepsilon}\right)-p\left(r_{\varepsilon}\right)\right]+\varepsilon\left(r_{\varepsilon}-\varrho_{\varepsilon} \Theta_{\varepsilon}\right) H^{\prime \prime}\left(r_{\varepsilon}\right) \operatorname{div}_{x}\left(s_{\varepsilon} \mathbf{U}_{\varepsilon}\right) \\
& +\varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x} H^{\prime}\left(r_{\varepsilon}\right) \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right)
\end{aligned}
$$

Furthermore,

$$
\begin{aligned}
& \varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x} H^{\prime}\left(r_{\varepsilon}\right) \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right)+\varrho_{\varepsilon} \nabla_{x} F \cdot\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right) \\
& =\varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x}\left[H^{\prime}\left(r_{\varepsilon}\right)-H^{\prime \prime}\left(\varrho_{e}\right)\left(r_{\varepsilon}-\varrho_{0}\right)-H^{\prime}\left(\varrho_{0}\right)\right] \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \\
& +\varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x}\left[H^{\prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)+H^{\prime}\left(\varrho_{0}\right)\right] \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right)+\varrho_{\varepsilon} \nabla_{x} H^{\prime}\left(\varrho_{0}\right) \cdot\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right) \\
& =\varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x}\left[H^{\prime}\left(r_{\varepsilon}\right)-H^{\prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)-H^{\prime}\left(\varrho_{0}\right)\right] \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \\
& +\varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x}\left[H^{\prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)\right] \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right)+\varrho_{\varepsilon}\left(1-\Theta_{\varepsilon}\right) \nabla_{x} H^{\prime}\left(\varrho_{0}\right) \cdot\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right) .
\end{aligned}
$$

Finally, using (4.2), we conclude

$$
\varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x}\left[H^{\prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)\right] \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right)=-\varepsilon^{2} \varrho_{\varepsilon} \Theta_{\varepsilon} \partial_{t} \nabla_{x} \Phi_{\varepsilon} \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right)
$$

Summarizing the previous relations we may rewrite (3.2) in the form

$$
\begin{align*}
& {\left[\mathcal{E}\left(\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon} \mid r_{\varepsilon}, \mathbf{U}_{\varepsilon}\right)\right]_{t=0}^{t=\tau}+\varepsilon^{\alpha} \int_{0}^{\tau} \int_{R^{3}} \mathbb{S}\left(\nabla_{x}\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right)\right): \nabla_{x}\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t} \\
& \leq \int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(\partial_{t} \mathbf{V}+\mathbf{u}_{\varepsilon} \cdot \nabla_{x} \mathbf{U}_{\varepsilon}\right) \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right)+\varepsilon^{\alpha} \mathbb{S}\left(\nabla_{x} \mathbf{U}_{\varepsilon}\right): \nabla_{x}\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& -\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}} \operatorname{div}_{x} \mathbf{U}_{\varepsilon}\left[p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(r_{\varepsilon}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-r_{\varepsilon}\right)-p\left(r_{\varepsilon}\right)\right] \mathrm{d} x \mathrm{~d} t \\
& +\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x}\left[H^{\prime}\left(r_{\varepsilon}\right)-H^{\prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)-H^{\prime}\left(\varrho_{0}\right)\right] \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& +\frac{1}{\varepsilon} \int_{0}^{\tau} \int_{R^{3}}\left(r_{\varepsilon}-\varrho_{\varepsilon} \Theta_{\varepsilon}\right) H^{\prime \prime}\left(r_{\varepsilon}\right) \operatorname{div}_{x}\left(s_{\varepsilon} \mathbf{U}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& +\int_{0}^{\tau} \int_{R^{3}}\left[\varrho_{\varepsilon}\left(1-\Theta_{\varepsilon}\right) \partial_{t} \nabla_{x} \Phi_{\varepsilon} \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right)+\frac{\varrho_{\varepsilon}}{\varepsilon^{2}}\left(1-\Theta_{\varepsilon}\right) \nabla_{x} H^{\prime}\left(\varrho_{0}\right) \cdot\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right)\right] \mathrm{d} x \mathrm{~d} t \tag{5.5}
\end{align*}
$$

### 5.1.2 Step 2 - the diffusion term

It follows from estimate (3.9) that

$$
\varepsilon^{\alpha} \int_{0}^{\tau} \int_{R^{3}} \mathbb{S}\left(\nabla_{x} \mathbf{U}_{\varepsilon}\right): \nabla_{x}\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \rightarrow 0 \text { as } \varepsilon \rightarrow 0
$$

### 5.1.3 Step 3 - the velocity term

We write

$$
\begin{aligned}
\int_{0}^{\tau} & \int_{R^{3}} \varrho_{\varepsilon}\left(\partial_{t} \mathbf{V}+\mathbf{u}_{\varepsilon} \cdot \nabla_{x} \mathbf{U}_{\varepsilon}\right) \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& =\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right) \cdot \nabla_{x} \mathbf{U}_{\varepsilon} \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& +\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(\mathbf{V} \cdot \nabla_{x}^{2} \Phi_{\varepsilon}+\nabla_{x} \Phi_{\varepsilon} \cdot \nabla_{x} \mathbf{U}_{\varepsilon}\right) \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& +\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(\partial_{t} \mathbf{V}+\mathbf{V} \cdot \nabla_{x} \mathbf{V}\right) \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t
\end{aligned}
$$

where the first integral on the right-hand side can be controlled by a similar term in $\mathcal{E}$.
Next, the second integral can be handled as

$$
\begin{aligned}
& \left|\int_{R^{3}} \varrho_{\varepsilon}\left(\mathbf{V} \cdot \nabla_{x}^{2} \Phi_{\varepsilon}+\nabla_{x} \Phi_{\varepsilon} \cdot \nabla_{x} \mathbf{U}_{\varepsilon}\right) \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x\right| \\
& \leq\left\|\nabla_{x} \Phi_{\varepsilon}\right\|_{W^{1, \infty}\left(R^{3} ; R^{3}\right)}^{2}\left(\|\mathbf{V}\|_{L^{\infty}\left(R^{3} ; R^{3}\right)}+\left\|\nabla_{x} \mathbf{U}_{\varepsilon}\right\|_{L^{\infty}\left(R^{3} ; R^{3 \times 3}\right)}\right)^{2}+c \mathcal{E},
\end{aligned}
$$

where the first term vanishes in the asymptotic limit thanks to the dispersive estimates (4.24).

The last term reads

$$
\begin{aligned}
& \int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(\partial_{t} \mathbf{V}+\mathbf{V} \cdot \nabla_{x} \mathbf{V}\right) \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& -\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon} \nabla_{x} \Pi \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t-\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon} \mathcal{T} \nabla_{x} F \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t
\end{aligned}
$$

where

$$
\begin{aligned}
& \int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon} \nabla_{x} \Pi \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& =-\int_{0}^{\tau} \int_{\Omega} \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \cdot \nabla_{x} \Pi \mathrm{~d} x \mathrm{~d} t+\int_{0}^{\tau} \int_{\Omega} \varrho_{\varepsilon}\left(\mathbf{V}+\nabla_{x} \Phi_{\varepsilon}\right) \cdot \nabla_{x} \Pi \mathrm{~d} x \mathrm{~d} t .
\end{aligned}
$$

In view of (3.3), (3.12),

$$
\left.\varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \rightarrow \overline{\varrho \mathbf{u}} \text { weakly-( }{ }^{*}\right) \text { in } L^{\infty}\left(0, T ;\left(L^{2}+L^{2 \gamma /(\gamma+1)}\left(R^{3} ; R^{3}\right)\right),\right.
$$

where, as can be deduced from the equation of continuity and (3.12),

$$
\operatorname{div}_{x} \bar{\varrho} \overline{\mathbf{u}}=0 .
$$

Consequently,

$$
\int_{0}^{\tau} \int_{\Omega} \varrho_{\varepsilon} \mathbf{u}_{\varepsilon} \cdot \nabla_{x} \Pi \mathrm{~d} x \mathrm{~d} t \rightarrow 0 \text { as } \varepsilon \rightarrow 0
$$

Similarly, we may combine the dispersive estimates (4.24) with the anelastic constraint (1.1) to deduce that

$$
\int_{0}^{\tau} \int_{\Omega} \varrho_{\varepsilon}\left(\mathbf{V}+\nabla_{x} \Phi_{\varepsilon}\right) \cdot \nabla_{x} \Pi \mathrm{~d} x \mathrm{~d} t \rightarrow 0 \text { as } \varepsilon \rightarrow 0
$$

Thus (5.5) can be reduced to

$$
\begin{align*}
& {\left[\mathcal{E}\left(\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon} \mid r_{\varepsilon}, \mathbf{U}_{\varepsilon}\right)\right]_{t=0}^{t=\tau}} \\
& \leq \frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x}\left[H^{\prime}\left(r_{\varepsilon}\right)-H^{\prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)-H^{\prime}\left(\varrho_{0}\right)\right] \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& -\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}} \operatorname{div}_{x} \mathbf{U}_{\varepsilon}\left[p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(r_{\varepsilon}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-r_{\varepsilon}\right)-p\left(r_{\varepsilon}\right)\right] \mathrm{d} x \mathrm{~d} t \\
& +\frac{1}{\varepsilon} \int_{0}^{\tau} \int_{R^{3}}\left(r_{\varepsilon}-\varrho_{\varepsilon} \Theta_{\varepsilon}\right) H^{\prime \prime}\left(r_{\varepsilon}\right) \operatorname{div}_{x}\left(s_{\varepsilon} \mathbf{U}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t  \tag{5.6}\\
& +\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(1-\Theta_{\varepsilon}\right) \partial_{t} \nabla_{x} \Phi_{\varepsilon} \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& +\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(\frac{1-\Theta_{\varepsilon}}{\varepsilon^{2}}-\mathcal{T}\right) \nabla_{x} F \cdot\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& +c \int_{0}^{\tau} \mathcal{E}\left(\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon} \mid r_{\varepsilon}, \mathbf{U}_{\varepsilon}\right) \mathrm{d} t+o(\varepsilon),
\end{align*}
$$

where

$$
o(\varepsilon) \rightarrow 0 \text { as } \varepsilon \rightarrow 0
$$

### 5.1.4 Step 4 - controlling pressure, I

We compute

$$
\begin{aligned}
\nabla_{x}\left(H^{\prime}\left(r_{\varepsilon}\right)\right. & \left.-H^{\prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)-H^{\prime}\left(\varrho_{0}\right)\right) \\
& =\varepsilon\left(H^{\prime \prime}\left(r_{\varepsilon}\right)-H^{\prime \prime}\left(\varrho_{0}\right)\right) \nabla_{x} s_{\varepsilon}+\left(P^{\prime \prime}\left(r_{\varepsilon}\right)-P^{\prime \prime}\left(\varrho_{0}\right)-P^{\prime \prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)\right) \nabla_{x} \varrho_{0}
\end{aligned}
$$

therefore we may use Taylor's formula to deduce

$$
\left|\nabla_{x}\left(H^{\prime}\left(r_{\varepsilon}\right)-H^{\prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)-H^{\prime}\left(\varrho_{0}\right)\right)\right| \leq \varepsilon^{2} c\left(s_{\varepsilon}\left|\nabla_{x} s_{\varepsilon}\right|+s_{\varepsilon}^{2}\right)
$$

Thus we conclude

$$
\begin{aligned}
& \frac{1}{\varepsilon^{2}}\left|\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon} \Theta_{\varepsilon} \nabla_{x}\left[H^{\prime}\left(r_{\varepsilon}\right)-H^{\prime \prime}\left(\varrho_{0}\right)\left(r_{\varepsilon}-\varrho_{0}\right)-H^{\prime}\left(\varrho_{0}\right)\right] \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t\right| \\
& \quad \leq c \int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(s_{\varepsilon}\left|\nabla_{x} s_{\varepsilon}\right|+s_{\varepsilon}^{2}\right)\left|\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right| \mathrm{d} x \mathrm{~d} t
\end{aligned}
$$

where, by virtue of the dispersive estimates (4.24), the last integral is dominated by

$$
o(\varepsilon)+c \int_{0}^{\tau} \mathcal{E}\left(\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon} \mid r_{\varepsilon}, \mathbf{U}_{\varepsilon}\right) \mathrm{d} t
$$

Relation (5.6) reduces to

$$
\begin{align*}
& {\left[\mathcal{E}\left(\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon} \mid r_{\varepsilon}, \mathbf{U}_{\varepsilon}\right)\right]_{t=0}^{t=\tau}} \\
& -\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}} \operatorname{div}_{x} \mathbf{U}_{\varepsilon}\left[p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(r_{\varepsilon}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-r_{\varepsilon}\right)-p\left(r_{\varepsilon}\right)\right] \mathrm{d} x \mathrm{~d} t \\
& +\frac{1}{\varepsilon} \int_{0}^{\tau} \int_{R^{3}}\left(r_{\varepsilon}-\varrho_{\varepsilon} \Theta_{\varepsilon}\right) H^{\prime \prime}\left(r_{\varepsilon}\right) \operatorname{div}_{x}\left(s_{\varepsilon} \mathbf{U}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& +\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(1-\Theta_{\varepsilon}\right) \partial_{t} \nabla_{x} \Phi_{\varepsilon} \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t  \tag{5.7}\\
& +\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(\frac{1-\Theta_{\varepsilon}}{\varepsilon^{2}}-\mathcal{T}\right) \nabla_{x} F \cdot\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t \\
& +c \int_{0}^{\tau} \mathcal{E}\left(\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon} \mid r_{\varepsilon}, \mathbf{U}_{\varepsilon}\right) \mathrm{d} t+o(\varepsilon)
\end{align*}
$$

### 5.1.5 Step 5 - controlling pressure, II

Probably the most delicate step is to handle the term

$$
\begin{align*}
& \frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}} \operatorname{div}_{x} \mathbf{U}_{\varepsilon}\left[p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(r_{\varepsilon}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-r_{\varepsilon}\right)-p\left(r_{\varepsilon}\right)\right] \mathrm{d} x \mathrm{~d} t \\
& \quad=\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}}\left(\operatorname{div}_{x} \mathbf{V}+\Delta_{x} \Phi_{\varepsilon}\right)\left[p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(\varrho_{0}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right)-p\left(\varrho_{0}\right)\right] \mathrm{d} x \mathrm{~d} t  \tag{5.8}\\
& \quad+\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}}\left(\operatorname{div}_{x} \mathbf{V}+\Delta_{x} \Phi_{\varepsilon}\right)\left[p\left(\varrho_{0}\right)-p\left(\varrho_{0}+\varepsilon s_{\varepsilon}\right)+\varepsilon p^{\prime}\left(\varrho_{0}+\varepsilon s_{\varepsilon}\right) s_{\varepsilon}\right] \mathrm{d} x \mathrm{~d} t \\
& \quad+\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}}\left(\operatorname{div}_{x} \mathbf{V}+\Delta_{x} \Phi_{\varepsilon}\right)\left[p^{\prime}\left(\varrho_{0}\right)-p^{\prime}\left(\varrho_{0}+\varepsilon s_{\varepsilon}\right)\right]\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right) \mathrm{d} x \mathrm{~d} t
\end{align*}
$$

where

$$
\left|\frac{p\left(\varrho_{0}\right)-p\left(\varrho_{0}+\varepsilon s_{\varepsilon}\right)+\varepsilon p^{\prime}\left(\varrho_{0}+\varepsilon s_{\varepsilon}\right) s_{\varepsilon}}{\varepsilon^{2}}\right| \leq c\left|s_{\varepsilon}\right|^{2}
$$

and, similarly,

$$
\left|\frac{p^{\prime}\left(\varrho_{0}\right)-p^{\prime}\left(\varrho_{0}+\varepsilon s_{\varepsilon}\right)}{\varepsilon}\right|\left|\frac{\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right)}{\varepsilon}\right| \leq c\left|s_{\varepsilon}\right|\left|\frac{\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right)}{\varepsilon}\right| .
$$

Thus the last two integrals in (5.8) can be controlled by means of (3.7) and the dispersive estimates (4.24).

Next, observing that $\Delta_{x} \Phi_{\varepsilon}$ vanishes for $\varepsilon \rightarrow 0$ because of (4.24) we have

$$
\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}} \Delta_{x} \Phi_{\varepsilon}\left(p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(\varrho_{0}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right)-p\left(\varrho_{0}\right)\right) \mathrm{d} x \mathrm{~d} t \rightarrow 0 \text { in } L^{1}(0, T) \text { as } \varepsilon \rightarrow 0
$$

The next step is to handle

$$
\begin{aligned}
& \frac{1}{\varepsilon^{2}}\left|\int_{0}^{\tau} \int_{R^{3}} \operatorname{div}_{x} \mathbf{V}\left(p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(\varrho_{0}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right)-p\left(\varrho_{0}\right)\right) \mathrm{d} x \mathrm{~d} t\right| \\
& \quad \leq \frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}}^{\tau}\left|\left[p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(\varrho_{0}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right)-p\left(\varrho_{0}\right)\right]_{\operatorname{ess}}\right|\left|\operatorname{div}_{x} \mathbf{V}\right| \mathrm{d} x \mathrm{~d} t \\
& \frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{R^{3}}^{\tau}\left|\left[p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(\varrho_{0}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right)-p\left(\varrho_{0}\right)\right]_{\text {res }}\right|\left|\operatorname{div}_{x} \mathbf{V}\right| \mathrm{d} x \mathrm{~d} t
\end{aligned}
$$

As the limit velocity field $\mathbf{V}$ obeys the anelastic constraint (1.1), we get

$$
\operatorname{div}_{x} \mathbf{V}=-\frac{\nabla_{x} \varrho_{0}}{\varrho_{0}} \cdot \mathbf{V}
$$

Since $\varrho_{0}$ is given by (1.4), we may find a compact set $K \subset R^{3}$ such that $\operatorname{div}_{x} \mathbf{V}$ is small in $R^{3} \backslash K$. Thus we may use the local pressure estimates established in (3.13) to obtain

$$
\frac{1}{\varepsilon^{2}} \int_{0}^{\tau} \int_{K}\left|\left[p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(\varrho_{0}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right)-p\left(\varrho_{0}\right)\right]_{\mathrm{res}}\right|\left|\operatorname{div}_{x} \mathbf{V}\right| \mathrm{d} x \mathrm{~d} t \rightarrow 0 \text { as } \varepsilon \rightarrow 0
$$

while the integral over the complement $R^{3} \backslash K$ is small because of (3.8).
Remark 5.1. This is the only moment in the proof where the pressure estimates are needed.

Finally,

$$
\begin{aligned}
& \left|\int_{0}^{\tau} \int_{R^{3}}\left[\frac{p\left(\varrho_{\varepsilon} \Theta_{\varepsilon}\right)-p^{\prime}\left(\varrho_{0}\right)\left(\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}\right)-p\left(\varrho_{0}\right)}{\varepsilon^{2}}\right]_{\mathrm{ess}}\right| \operatorname{div}_{x} \mathbf{V}|\mathrm{~d} x \mathrm{~d} t| \\
& \leq\left. c \int_{0}^{\tau} \int_{R^{3}}\left|\left[\frac{\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}}{\varepsilon}\right]_{\mathrm{ess}}\right|\right|^{2}\left|\operatorname{div}_{x} \mathbf{V}\right| \mathrm{d} x \mathrm{~d} t \\
& \leq c \int_{0}^{\tau} \int_{R^{3}}\left(\left[\frac{\varrho_{\varepsilon} \Theta_{\varepsilon}-\varepsilon s_{\varepsilon}-\varrho_{0}}{\varepsilon}\right]_{\mathrm{ess}}^{2}+\left|s_{\varepsilon}\right|^{2}\right)\left|\operatorname{div}_{x} \mathbf{V}\right| \mathrm{d} x \mathrm{~d} t \\
& \leq c \int_{0}^{\tau} \mathcal{E}\left(\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon} \mid r_{\varepsilon} \cdot \mathbf{U}_{\varepsilon}\right)+\int_{0}^{\tau} \int_{R^{3}}\left|s_{\varepsilon}\right|^{2}\left|\operatorname{div}_{x} \mathbf{V}\right| \mathrm{d} x \mathrm{~d} t .
\end{aligned}
$$

### 5.1.6 Step 6 - conclusion

To conclude, we write

$$
\begin{aligned}
& \frac{1}{\varepsilon}\left|\int_{0}^{\tau} \int_{R^{3}}\left(r_{\varepsilon}-\varrho_{\varepsilon} \Theta_{\varepsilon}\right) H^{\prime \prime}\left(r_{\varepsilon}\right) \operatorname{div}_{x}\left(s_{\varepsilon} \mathbf{U}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t\right| \\
& \leq c \int_{0}^{\tau} \int_{R^{3}}\left(\left|\frac{\varrho_{\varepsilon} \Theta_{\varepsilon}-\varrho_{0}}{\varepsilon}\right|+\left|s_{\varepsilon}\right|\right)\left(\left|s_{\varepsilon}\right|+\left|\nabla_{x} s_{\varepsilon}\right|\right)\left(|\mathbf{V}|+\left|\operatorname{div}_{x} \mathbf{V}\right|+\left|\nabla_{x} \Phi_{\varepsilon}\right|+\left|\Delta_{x} \Phi_{\varepsilon}\right|\right) \mathrm{d} x \mathrm{~d} t
\end{aligned}
$$

where again the right-hand side vanishes for $\varepsilon \rightarrow 0$ due to the dispersive estimates (4.24).

Finally, the integral

$$
\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(1-\Theta_{\varepsilon}\right) \partial_{t} \nabla_{x} \Phi_{\varepsilon} \cdot\left(\mathbf{U}_{\varepsilon}-\mathbf{u}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t
$$

is small because of (3.6), and

$$
\int_{0}^{\tau} \int_{R^{3}} \varrho_{\varepsilon}\left(\frac{1-\Theta_{\varepsilon}}{\varepsilon^{2}}-\mathcal{T}\right) \nabla_{x} F \cdot\left(\mathbf{u}_{\varepsilon}-\mathbf{U}_{\varepsilon}\right) \mathrm{d} x \mathrm{~d} t
$$

is controlled by the initial data satisfying (1.15).
Thus we may infer that

$$
\begin{array}{r}
{\left[\mathcal{E}\left(\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon} \mid r_{\varepsilon}, \mathbf{U}_{\varepsilon}\right)\right]_{t=0}^{t=\tau} \leq c \int_{0}^{\tau} \mathcal{E}\left(\varrho_{\varepsilon}, \Theta_{\varepsilon}, \mathbf{u}_{\varepsilon} \mid r_{\varepsilon}, \mathbf{U}_{\varepsilon}\right) \mathrm{d} t+o(\varepsilon),} \\
o(\varepsilon) \rightarrow 0 \text { as } \varepsilon \rightarrow 0
\end{array}
$$

and use Gronwall's lemma to complete the proof of Theorem 2.2.

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[^0]:    *The research of D.D. was supported by European Union MSCA-ITN-2014-ETN - Marie Sklodowska-Curie Innovative Training Networks (ITN-ETN) under grant agreement 642768 ModCompShock
    ${ }^{\dagger}$ The research of E.F. leading to these results has received funding from the European Research Council under the European Union's Seventh Framework Programme (FP7/2007-2013)/ ERC Grant Agreement 320078. The Institute of Mathematics of the Academy of Sciences of the Czech Republic is supported by RVO:67985840. E.F. thanks Gran Sasso Science Institute for hospitality and support during his stay in L'Aquila.

