Weak vs. strong solutions of problems in fluid mechanics

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Eduard Feireisl Weak vs. strong

Dynamics of real fluids

No fluid is perfect



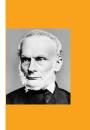
Real fluids are:

- three dimensional
- viscous
- heat conductive
- compressible
- obeying the basic laws of thermodynamics

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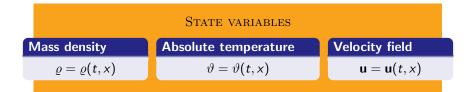
Long time behavior of energetically closed systems

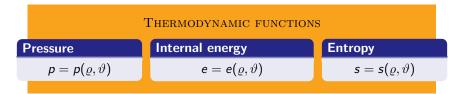


Die Energie der Welt ist constant; Die Entropie der Welt strebt einem Maximum zu

Rudolph Clausius, 1822-1888

Mathematical model







Field equations



Equation of continuity

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = \mathbf{0}$$

Claude Louis Marie Henri Navier [1785-1836]

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) + \nabla_x p(\varrho, \vartheta) = \operatorname{div}_x \mathbb{S} + \varrho \mathbf{f}$$



George Gabriel Stokes [1<mark>819-1903]</mark>

Entropy production

$$\partial_t(\varrho s(\varrho, \vartheta)) + \operatorname{div}_x(\varrho s(\varrho, \vartheta) \mathbf{u}) + \operatorname{div}_x\left(\frac{\mathbf{q}}{\vartheta}\right) = \sigma$$
$$\sigma = (\geq) \frac{1}{\vartheta} \left(\mathbb{S} : \nabla_x \mathbf{u} - \frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} \right)$$

Constitutive relations





Isaac Newton [1643-1727]

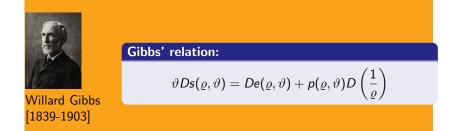
Newton's rheological law

$$\mathbb{S} = \mu(\vartheta) \left(\nabla_{\mathsf{x}} \mathsf{u} + \nabla_{\mathsf{x}}^{t} \mathsf{u} - \frac{2}{3} \mathrm{div}_{\mathsf{x}} \mathsf{u} \right) + \eta(\vartheta) \mathrm{div}_{\mathsf{x}} \mathsf{u} \mathbb{I}$$

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Weak vs. strong

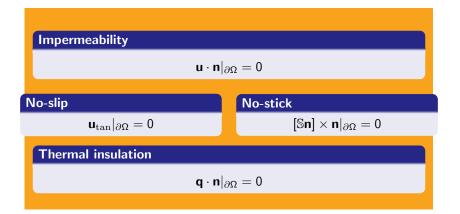
Gibbs' relation



Thermodynamics stability: $\frac{\partial p(\varrho, \vartheta)}{\partial \varrho} > 0, \ \frac{\partial e(\varrho, \vartheta)}{\partial \vartheta} > 0$

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Boundary conditions



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Weak solutions to the complete system

- Equation of continuity holds in the sense of distributions (renormalized equation also satisfied)
- Momentum balance holds in the sense of distributions
- Entropy production equation holds in the sense of distributions, entropy production rate satisfies the inequality
- The system is augmented by

Total energy balance

$$\frac{\mathrm{d}}{\mathrm{d}t} \int_{\Omega} \left[\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho \boldsymbol{e}(\varrho, \vartheta) \right] \, \mathrm{d}x = \int_{\Omega} \varrho \mathbf{f} \cdot \mathbf{u} \, \mathrm{d}x$$
$$\int_{\Omega} \varrho \mathbf{f} \cdot \mathbf{u} \, \mathrm{d}x = \frac{\mathrm{d}}{\mathrm{d}t} \int_{\Omega} \varrho F \, \mathrm{d}x \text{ if } \mathbf{f} = \nabla_x F(x)$$

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Technical hypotheses

Pressure

$$p(\varrho, \vartheta) = \vartheta^{5/2} P\left(\frac{\varrho}{\vartheta^{3/2}}\right) + \frac{a}{3} \vartheta^4$$

$$P(0)=0,\,\,P'(Z)>0,\,\,P(Z)/Z^{5/3}
ightarrow p_{\infty}>0$$
 as $Z
ightarrow\infty$

Internal energy

$$e(\varrho, \vartheta) = rac{3}{2} artheta rac{artheta^{3/2}}{arrho} P\left(rac{arrho}{artheta^{3/2}}
ight) + rac{a}{arrho} artheta^4$$

Transport coefficients

$$\mu(\vartheta) pprox (1+artheta^lpha), \; lpha \in [1/2,1], \; \kappa(artheta) pprox (1+artheta^3)$$

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Weak vs. strong solutions

Local existence of strong solutions

Strong (classical) solutions exist on a (possibly) short time interval for general data and globally for "small" data [Matsumura and Nishida, Valli and Zajaczkowski]

Global existence of weak solutions

Weak (distributional) solutions exist globally in time [EF and A.Novotný]

Weak strong uniqueness

A weak solution coincides with the strong solution emanating from the same initial data as long as the latter exists [EF and A.Novotný]

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Conservative vs. dissipative system

Conservative character

total mass
$$\int_{\Omega} \varrho(t, \cdot) \, \mathrm{d}x = M_0,$$

total energy $\int_{\Omega} \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho, \vartheta) - \varrho F \right)(t, \cdot) \, \mathrm{d}x = E_0$

Dissipative character

total entropy
$$\int_{\Omega} \rho s(\rho, \vartheta) \, \mathrm{d}x = S(t) \nearrow S_{\infty}$$

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Equilibrium solutions

Conservative driving force

 $\mathbf{f} = \nabla_x F, \ F = F(x)$

TOTAL ENERGY CONSERVATION

$$\frac{\mathrm{d}}{\mathrm{d}t}\int_{\Omega}\left(\frac{1}{2}\varrho|\mathbf{u}|^{2}+\varrho\boldsymbol{e}(\varrho,\vartheta)-\varrho\boldsymbol{F}\right) \,\mathrm{d}x=0$$

Static solutions

$$abla_{x} p(ilde{arrho}, \overline{artheta}) = ilde{arrho}
abla_{x} F, \ \overline{artheta} > 0 \ {
m constant}$$

Total mass and energy

$$\int_{\Omega} \tilde{\varrho} \, \mathrm{d} x = M_0, \ \int_{\Omega} \left(\tilde{\varrho} e(\tilde{\varrho}, \overline{\vartheta}) - \tilde{\varrho} F \right) \, \mathrm{d} x = E_0$$

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Total dissipation balance

Ballistic free energy

$$H_{\Theta}(\varrho, \vartheta) = \varrho\Big(e(\varrho, \vartheta) - \Theta s(\varrho, \vartheta)\Big)$$

Relative entropy

 $\mathcal{E}(\varrho, \vartheta, \mathbf{u} | \tilde{\varrho}, \overline{\vartheta})$

$$=\int_{\Omega}\left(\frac{1}{2}\varrho|\mathbf{u}|^{2}+H_{\overline{\vartheta}}(\varrho,\vartheta)-\partial_{\varrho}H_{\overline{\vartheta}}(\tilde{\varrho},\overline{\vartheta})(\varrho-\tilde{\varrho})-H_{\overline{\vartheta}}(\tilde{\varrho},\overline{\vartheta})\right) \, \mathrm{d}x$$

Total dissipation balance

$$\frac{\mathrm{d}}{\mathrm{d}t}\mathcal{E}(\varrho,\vartheta,\mathbf{u}|\tilde{\varrho},\overline{\vartheta}) + \int_{\Omega}\sigma \,\,\mathrm{d}x = \mathbf{0}$$

$$\tilde{\varrho}, \ \overline{artheta} \ -$$
 equilibrium state

Thermodynamic stability

Positive compressibility and specific heat

$$\frac{\partial \boldsymbol{p}(\varrho,\vartheta)}{\partial \varrho} > 0, \ \frac{\partial \boldsymbol{e}(\varrho,\vartheta)}{\partial \vartheta} > 0$$

Coercivity of the ballistic free energy

 $\varrho \mapsto H_{\Theta}(\varrho, \Theta)$ strictly convex

 $\vartheta \mapsto H_{\Theta}(\varrho, \vartheta)$ decreasing for $\vartheta < \Theta$ and increasing for $\vartheta > \Theta$

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Long-time behavior for conservative driving forces

$$\mathbf{f}=\nabla_{\mathbf{x}}F,\ F=F(\mathbf{x})$$

$$\varrho(t,\cdot) \to \tilde{\varrho} \text{ in } L^{5/3}(\Omega) \text{ as } t \to \infty$$

$$v(\iota, \cdot) \rightarrow v \, \text{in } L \, (\Omega) \text{ as } \iota \rightarrow \infty$$

$$(arrho {f u})(t,\cdot)
ightarrow 0$$
 in $L^1(\Omega;R^3)$ as $t
ightarrow \infty$

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Attractors

Hypotheses

$$\begin{split} &\int_{\Omega} \varrho(t,\cdot) \, \mathrm{d} x > M_0, \ t > 0 \\ &\int_{\Omega} \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e(\varrho,\vartheta) - \varrho F \right)(t,\cdot) \, \mathrm{d} x < E_0, \ t > 0 \\ &\int_{\Omega} \varrho s(\varrho,\vartheta)(t,\cdot) \, \mathrm{d} x > S_0, \ t > 0 \end{split}$$

Conclusion

$$egin{aligned} \|arrho(t,\cdot)- ilde{arrho}\|_{L^{5/3}(\Omega)} au(arepsilon)\ \|arrho\mathbf{u}(t,\cdot)\|_{L^1(\Omega;R^3)} au(arepsilon) \end{aligned}$$

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Uniform decay of density oscillations

$$\partial_t \varrho_\varepsilon + \mathbf{u}_\varepsilon \cdot \nabla_x \varrho_\varepsilon = -\mathrm{div}_x \mathbf{u}_\varepsilon \ \varrho_\varepsilon$$

$$\varrho_{\varepsilon} \to \varrho, \ \varrho_{\varepsilon} \log(\varrho_{\varepsilon}) \to \overline{\varrho \log(\varrho)}$$
 weakly in L^1

$$d(t) = \int_{\Omega} \left(\overline{\varrho \log(\varrho)} - \varrho \log(\varrho) \right)(t, \cdot) \, \mathrm{d}x$$

Density oscillations decay

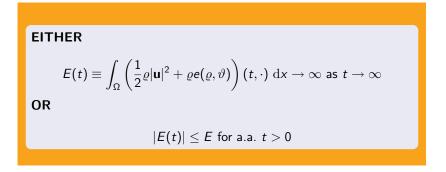
 $\partial_t d(t) + \Psi(d(t)) \leq 0$

$$\Psi(0) = 0, \ \Psi(d) > 0 \ \text{for} \ d > 0.$$

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General time-dependent driving forces

$$\mathbf{f} = \mathbf{f}(t, x), \ |\mathbf{f}(t, x)| \leq \overline{F}$$



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In the case $E(t) \leq E$, each sequence of times $\tau_n \to \infty$ contains a subsequence such that

$$f(au_n + \cdot, \cdot) o
abla_x F$$
 weakly-(*) in $L^{\infty}((0, 1) imes \Omega)$,

where F = F(x) may depend on $\{\tau_n\}$

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STEP 1:

Assume that $E(\tau_n) < E$ for certain $\tau_n \to \infty \Rightarrow$ total entropy remains bounded \Rightarrow integral of entropy production bounded

STEP 2:

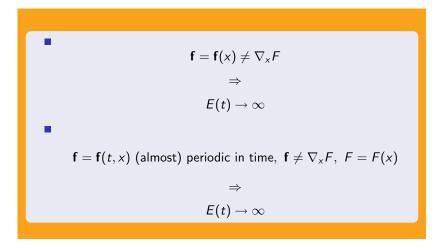
For
$$\tau_n \to \infty$$
 we have $\nabla_x p(\varrho, \vartheta) \approx \varrho \mathbf{f}$, $\vartheta \approx \overline{\vartheta}$, meaning, $\mathbf{f} \approx \nabla_x F$

STEP 3:

The energy cannot "oscillate" since bounded entropy *static solutions* have bounded total energy

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Corollaries



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Rapidly oscillating driving forces

Hypotheses:

$$egin{aligned} &\mathcal{H} = \omega(t^eta) \mathbf{w}(x), \mathbf{w} \in W^{1,\infty}(\Omega;R^3), \; eta > 2 \ &\omega \in L^\infty(R), \; \sup_{ au > 0} \left| \int_0^ au \omega(t) \; \mathrm{d}t
ight| < \infty \end{aligned}$$

Conclusion:

$$(\varrho \mathbf{u})(t, \cdot) \to 0 \text{ in } L^1(\Omega; R^3) \text{ as } t \to \infty$$

 $\varrho(t, \cdot) \to \overline{\varrho} \text{ in } L^{5/3}(\Omega) \text{ as } t \to \infty$
 $\vartheta(t, \cdot) \to \overline{\vartheta} \text{ in } L^4(\Omega) \text{ as } t \to \infty$

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Rapidly oscillating growing driving forces

Hypotheses:

$$\begin{aligned} \mathbf{f} &= t^{\delta} \omega(t^{\beta}) \mathbf{w}(x), \mathbf{w} \in W^{1,\infty}(\Omega; R^3) \\ \hline \delta &> 0, \ \beta - 2\delta > 2 \text{ or } \delta \leq 0, \ \beta - \delta > 2 \\ \hline \omega \in L^{\infty}(R), \ \sup_{\tau > 0} \left| \int_0^{\tau} \omega(t) \ \mathrm{d}t \right| < \infty \end{aligned}$$

Conclusion:

$$(\varrho \mathbf{u})(t, \cdot) \to 0 \text{ in } L^1(\Omega; \mathbb{R}^3) \text{ as } t \to \infty$$

 $\varrho(t, \cdot) \to \overline{\varrho} \text{ in } L^{5/3}(\Omega) \text{ as } t \to \infty$
 $\vartheta(t, \cdot) \to \overline{\vartheta} \text{ in } L^4(\Omega) \text{ as } t \to \infty$

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Time-periodic solutions and boundary dissipation

Dissipative boundary conditions

$$|\mathbf{u}|_{\partial\Omega} = 0, \ \mathbf{q} \cdot \mathbf{n} = d(x)(\vartheta - \tilde{\vartheta})$$

Time periodic forcing

$$\mathbf{f}(t+\omega,\cdot)=\mathbf{f}(t,\cdot)$$

Time periodic solutions

$$\varrho(t+\omega,\cdot) = \varrho(t,\cdot), \ \vartheta(t+\omega,\cdot) = \vartheta(t,\cdot), \ \mathbf{u}(t+\omega,\cdot) = \mathbf{u}(t,\cdot)$$

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