

Gravitational-wave astrophysics

Marco Cavaglià, University of Mississippi
for the LIGO Scientific Collaboration and the Virgo Collaboration



2016: One hundred (one) years of General Relativity



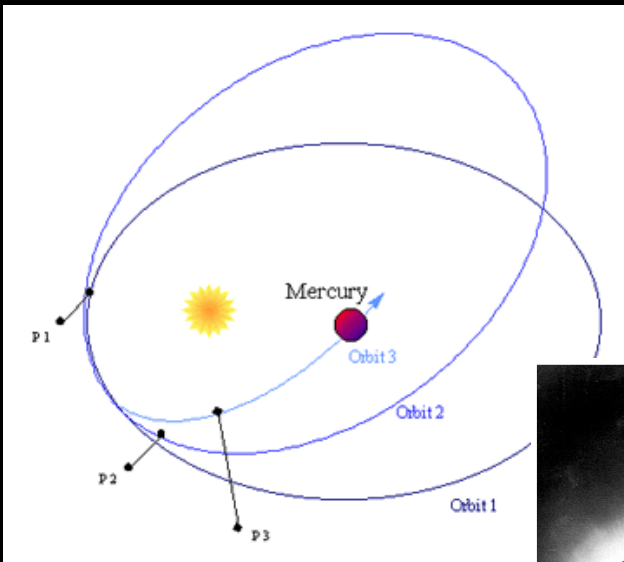
844 Sitzung der physikalisch-mathematischen Klasse vom 25. November 1915

Die Feldgleichungen der Gravitation.

VON A. EINSTEIN.

In zwei vor kurzem erschienenen Mitteilungen¹ habe ich gezeigt, wie man zu Feldgleichungen der Gravitation gelangen kann, die dem Postulat allgemeiner Relativität entsprechen, d. h. die in ihrer allgemeinen Fassung beliebigen Substitutionen der Raumzeitvariablen gegenüber kovariant sind.

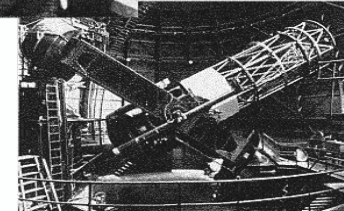
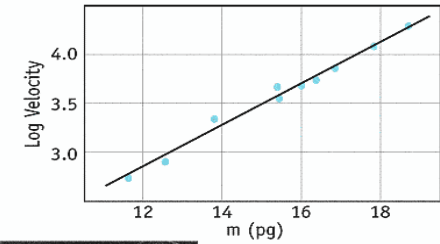
Der Entwicklungsgang war dabei folgender. Zunächst fand ich Gleichungen, welche die NEWTONSCHE Theorie als Näherung enthalten



DISCOVERY OF EXPANDING UNIVERSE



Edwin Hubble



Mt. Wilson
100 Inch
Telescope

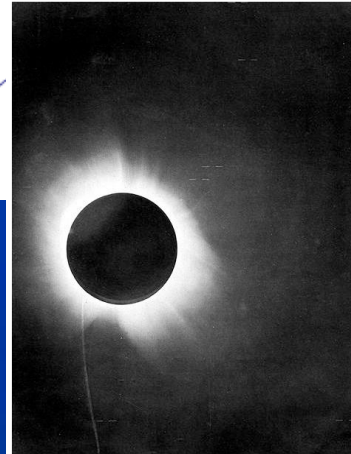


Plate 1, May 29, 1919, Sir Arthur Stanley Eddington Principe Expedition
From:
"A Determination of the Deflection of Light by the Sun's Gravitational Field,
from Observations Made at the Total Eclipse of May 29, 1919" Philosophical
Transactions of the Royal Society of London. Series A, Containing Papers of
a Mathematical or Physical Character (1920): 291-333, on 332, submitted by
F. W. Dyson, A. S. Eddington, and C. Davidson.



Sir Arthur Stanley Eddington (1882 - 1944)



The Gravity Probe B Experiment

...testing Einstein's Universe

Frame-dragging Effect
0.041 arcseconds/year
(0.000011 degrees/year)

Geodetic Effect
6.6 arcseconds/year
(0.0018 degrees/year)

Guide Star
IM Pegasi
(HR 8703)

642 kilometers
(401 miles)

One hundred (one)
years of successes

LIGHTS ALL ASKEW IN THE HEAVENS

Special Cable to THE NEW YORK TIMES.
New York Times 1857; Nov 10, 1919; ProQuest Historical Newspapers The New York Times (1851 - 2004)
pg. 17

LIGHTS ALL ASKEW IN THE HEAVENS

Men of Science More or Less
Agog Over Results of Eclipse
Observations.

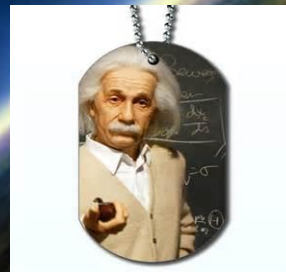
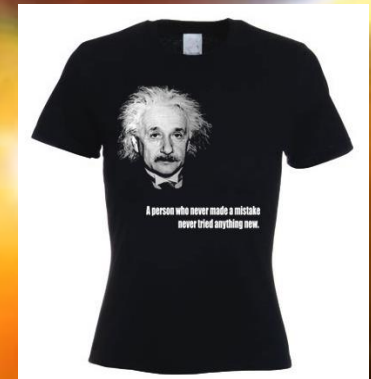
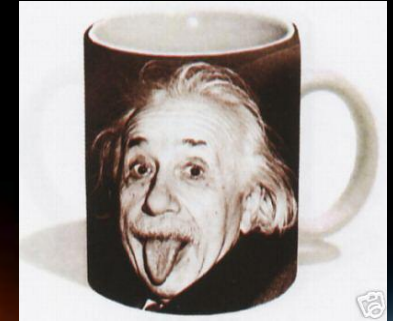
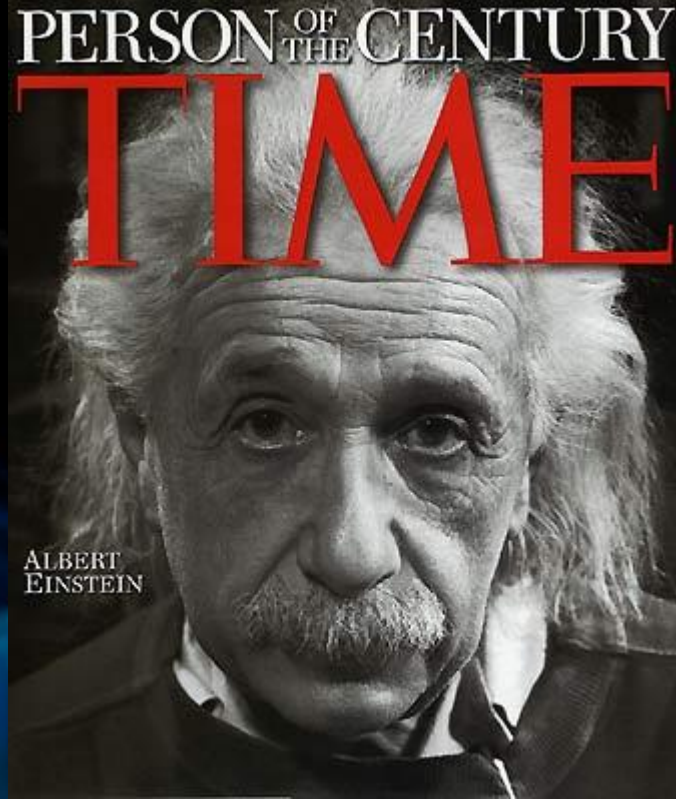
EINSTEIN THEORY TRIUMPHS

Stars Not Where They Seemed
or Were Calculated to be,
but Nobody Need Worry.

A BOOK FOR 12 WISE MEN

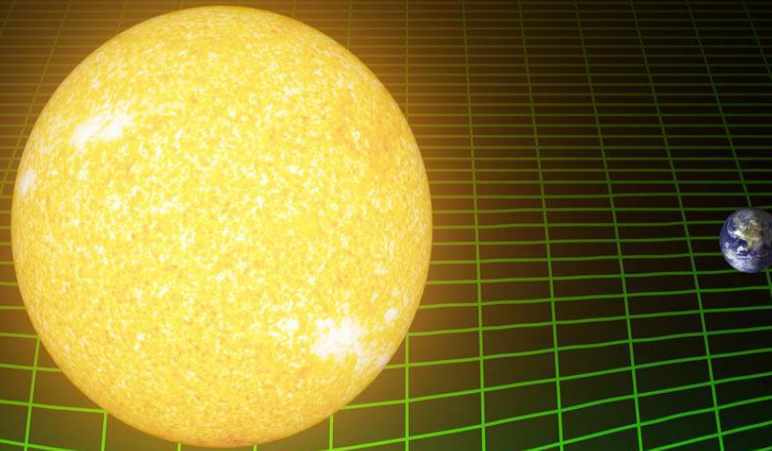
No More in All the World Could
Comprehend It, Said Einstein When
His Daring Publishers Accepted It.

New York Times headline of
November 10, 1919.



“Matter tells space time how to curve, and curved space time tells matter how to move.”

- John Wheeler



$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = -8\pi GT_{\mu\nu}$$

Über Gravitationswellen.

VON A. EINSTEIN.

(Vorgelegt am 31. Januar 1918 [s. oben S. 79].)

Die wichtige Frage, wie die Ausbreitung der Gravitationsfelder erfolgt, ist schon vor anderthalb Jahren in einer Akademiearbeit von mir behandelt worden¹. Da aber meine damalige Darstellung des Gegenstandes nicht genügend durchsichtig und außerdem durch einen bedauerlichen Rechenfehler verunstaltet ist, muß ich hier nochmals auf die Angelegenheit zurückkommen.

Wie damals beschränke ich mich auch hier auf den Fall, daß das betrachtete zeiträumliche Kontinuum sich von einem »galileischen« nur sehr wenig unterscheidet. Um für alle Indizes

$$g_{\mu\nu} = -\delta_{\mu\nu} + \gamma_{\mu\nu} \quad (1)$$

setzen zu können, wählen wir, wie es in der speziellen Relativitätstheorie üblich ist, die Zeitvariable x_4 rein imaginär, indem wir

$$x_4 = it$$

setzen, wobei t die »Lichtzeit« bedeutet. In (1) ist $\delta_{\mu\nu} = 1$ bzw. $\delta_{\mu\nu} = 0$, je nachdem $\mu = \nu$ oder $\mu \neq \nu$ ist. Die $\gamma_{\mu\nu}$ sind gegen 1 kleine Größen, welche die Abweichung des Kontinuums vom feldfreien darstellen; sie bilden einen Tensor vom zweiten Range gegenüber LORENTZ-Transformationen.

§ 1. Lösung der Näherungsgleichungen des Gravitationsfeldes durch retardierte Potentiale.

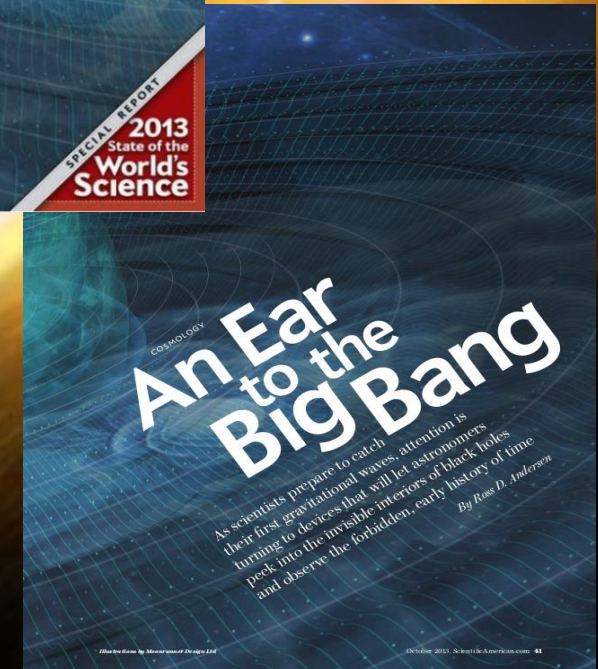
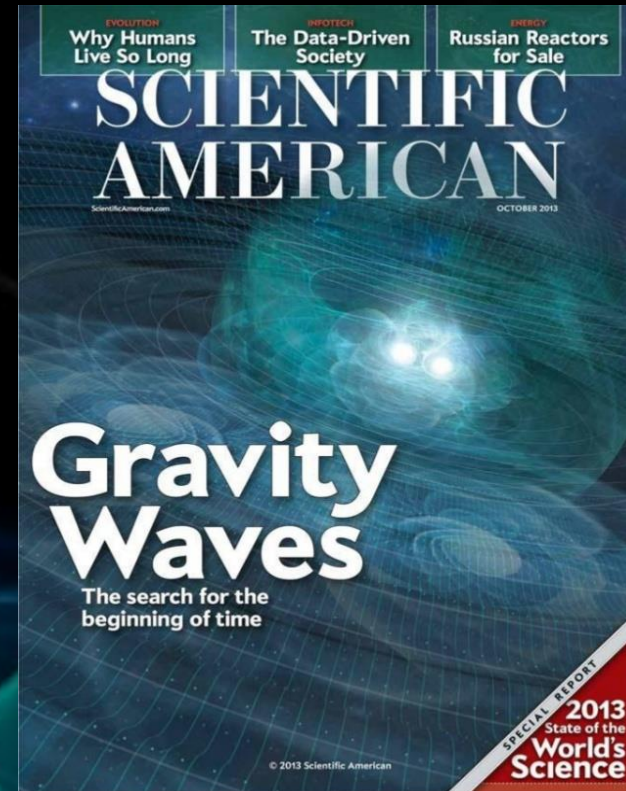
Wir gehen aus von den für ein beliebiges Koordinatensystem gültigen² Feldgleichungen

$$-\sum_{\alpha} \frac{\partial}{\partial x_{\alpha}} \left\{ \frac{\mu\nu}{\alpha} \right\} + \sum_{\alpha} \frac{\partial}{\partial x_{\alpha}} \left\{ \frac{\mu\alpha}{\alpha} \right\} + \sum_{\alpha\beta} \left\{ \frac{\mu\alpha}{\beta} \right\} \left\{ \frac{\nu\beta}{\alpha} \right\} - \sum_{\alpha\beta} \left\{ \frac{\mu\nu}{\alpha} \right\} \left\{ \frac{\alpha\beta}{\beta} \right\} \quad (2)$$

$$= -\kappa \left(T_{\mu\nu} - \frac{1}{2} g_{\mu\nu} T \right).$$

¹ Diese Sitzungsber. 1916, S. 688 ff.

² Von der Einführung des » γ -Gliedes« (vgl. diese Sitzungsber. 1917, S. 142) ist dabei Abstand genommen.

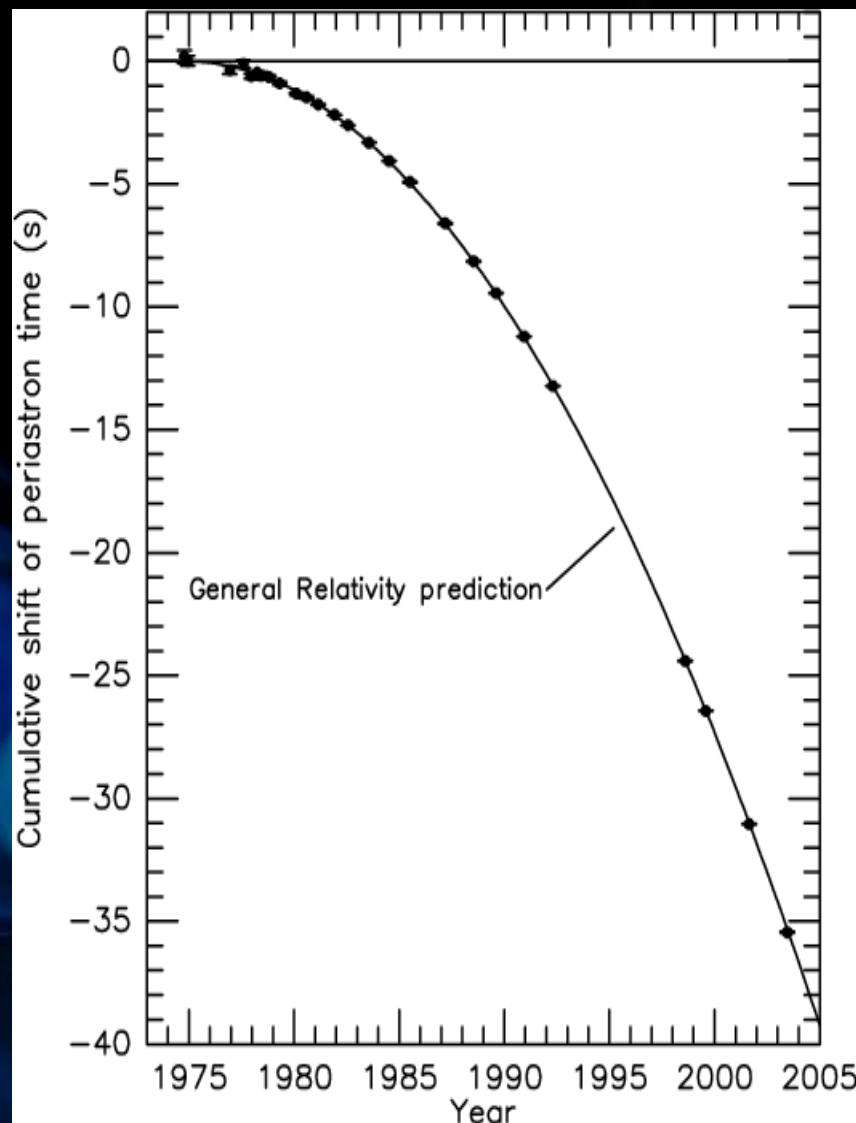


Einstein's Last Great Prediction...

1980's: Indirect evidence for gravitational waves



The Nobel Prize in Physics 1993 was awarded jointly to Russell A. Hulse and Joseph H. Taylor Jr. "for the discovery of a new type of pulsar, a discovery that has opened up new possibilities for the study of gravitation"



J.H. Taylor and J. Weisberg, "A New Test of General Relativity: Gravitational Radiation and the Binary Pulsar PSR 1913+16," *Astrophys. J.* 253, 908 (1982).

...but until now they had never been directly detected.

NAUGY FALLS
UCONN STUNNED
COMMUNITY NEWS

Bill to legalize pot bill
Lawmaker proposes allowing recreational use
EINSTEIN WAS RIGHT!
A CENTURY AFTER HE PREDICTED GRAVITATIONAL WAVES, SCIENTISTS PROVE IT IN A GROUND-BREAKING DISCOVERY

THE POST-STANDARD
COSMIC BREAKTHROUGH
DRIVING WITH SNOW ON YOUR CAR? NY LAWMAKERS PROPOSE FINES
'Einstein Would Be Beaming'

Deal would halt Syria fighting
World powers acted to allow humanitarian access to besieged areas
Clinton, Sanders cordial but firm
Felon steals to the north
Rap sheet to be cited
Music to their ears

POST STAR
FELON STEALS TO THE NORTH
RAP SHEET TO BE CITED
MUSIC TO THEIR EARS

apel grad aided in breakthrough
Discovery opens ears to cosmos

Tri-City Herald
Leap a wvc in W
Gravitational waves detected at LIGO

Pittsburgh Post-Gazette
Picking a major can be daunting
Gravitational waves discovery a 'new window' to universe

THE ADVOCATE
PLAY BALL! LSU softball team begins season with lofty goals
Baton Rouge on tap this weekend
'WE CAN HEAR THE UNIVERSE'
Livingston's LIGO confirms gravitational waves finding

Amtrak
Scientists have detected gravitational waves predicted a century ago
HOLY GRAIL OF SCIENCE!

The Boston Globe
Rivals debate their support for Obama
A whisper from across the universe
Deal set for truce, aid drops in Syria

The New York Times
Long in Clinton's Corner, Blacks Notice Sanders
Deal set for truce, aid drops in Syria

The Columbian Dispatch
With faint chirp, scientists prove Einstein correct

The Bakersfield Californian
Law suit disputes union access to farms
MAKING WAVES
Gravitational waves, that is: New scientific observation affirms Einstein's general theory of relativity

THE MONITOR
DHH investigating two possible cases of Zika
ZIKA VIRUS CASES IN THE U.S.
Council candidates discuss tax hike possibility at forum

SANTA FE NEW MEXICAN
Council candidates discuss tax hike possibility at forum
With faint chirp, scientists prove Einstein correct

USA TODAY WEEKEND
'WHOLE NEW WINDOW ON THE UNIVERSE'
RUSSIA, U.S. REACH DEAL IN SYRIA WAR

The Washington Post
U.S., Russia agree to a halt in Syrian war

The Huntsville Times
Huntsville ready to welcome Uber, Lyft and Zipcar

A FORCE AWAKENS
As Einstein said, scientists detect ripples in gravity

Collaboration proves Einstein correct
Former LIGO director reflects on historic wave detection

Daily Press
CNUT POLL SHOWS GUN RIGHTS WEIGH HEAVILY ON VIRGINIA VOTERS' MINDS
'A WINDOW ONTO THE UNIVERSE'

The Huntsville Times
'AT THE CENTER' OF CONFIRMING EINSTEIN'S THEORY OF RELATIVITY

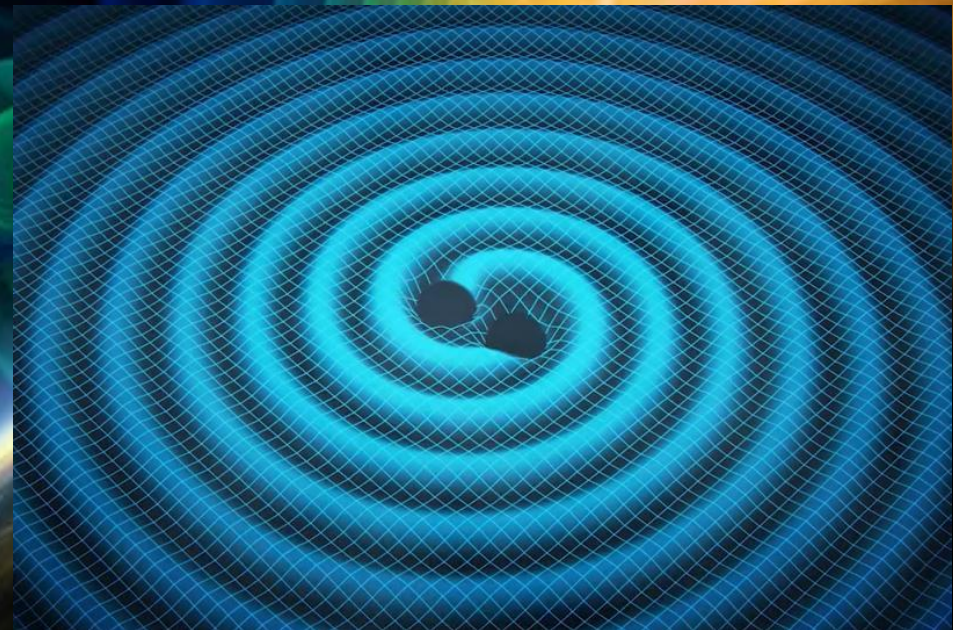
The Washington Post
Gravitational waves Einstein foresaw are detected

The Huntsville Times
Lawmakers hold key to building mega-prisons

A gravitational wave is a propagating disturbance of the space-time

When masses move rapidly, the space-time becomes stirred by their motion:

Gravitational waves start traveling outward with the speed of light



Gravitational waves in a nutshell

- Unavoidable consequence of General Relativity
- Classical phenomenon (as far as we know)
- Theory and phenomenology well understood
- Measured in the weak regime
- Can be used to probe strong gravitational fields
- Can be used to test GR or alternative theories
- Can be used to test fundamental physics
- Information complementary to light and particles

A bit of theory

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = -8\pi GT_{\mu\nu}$$

Weak field limit:

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$



$$\square^2 h_{\mu\nu} = -16\pi GS_{\mu\nu}$$

Vacuum solution

$$h_{\mu\nu}(x) = e_{\mu\nu} e^{ik_\lambda x^\lambda} + e_{\mu\nu}^* e^{-ik_\lambda x^\lambda}$$

where

$$k_\mu k^\mu = 0$$

$$k_\mu e^\mu{}_\nu = \frac{1}{2} k_\nu e^\mu{}_\mu$$



$$k^1 = k^2 = 0, \quad k^3 = k > 0$$

$$e_{11} = -e_{22}$$

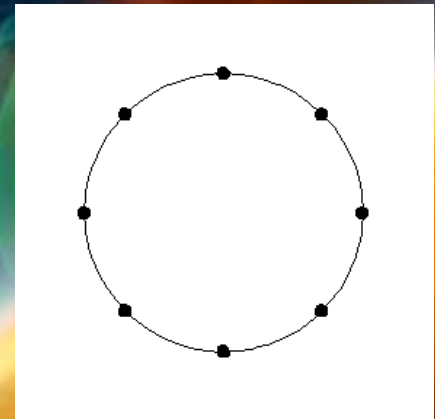
$$e_{12}$$

Transverse, two polarizations, speed of light

What is the effect of a gravitational wave?

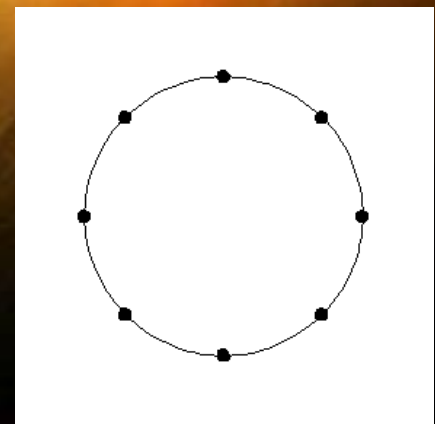
“+” polarization:

$$h_+(t - z) = h_{xx}^{TT} = -h_{yy}^{TT}$$



“x” polarization:

$$h_{\times}(t - z) = h_{xy}^{TT} = h_{yx}^{TT}$$





Gravitational waves are tiny!

For two coalescing compact objects into a black hole:

$$f \sim \frac{1}{M} \sim 10^4 \text{ Hz} \left(\frac{M_{\odot}}{M} \right)$$

$$h \sim \epsilon^{1/2} \frac{M}{r} \sim 10^{-21} \left(\frac{\epsilon}{0.01} \right)^{1/2} \left(\frac{M}{M_{\odot}} \right) \left(\frac{10 \text{ Mpc}}{r} \right)$$

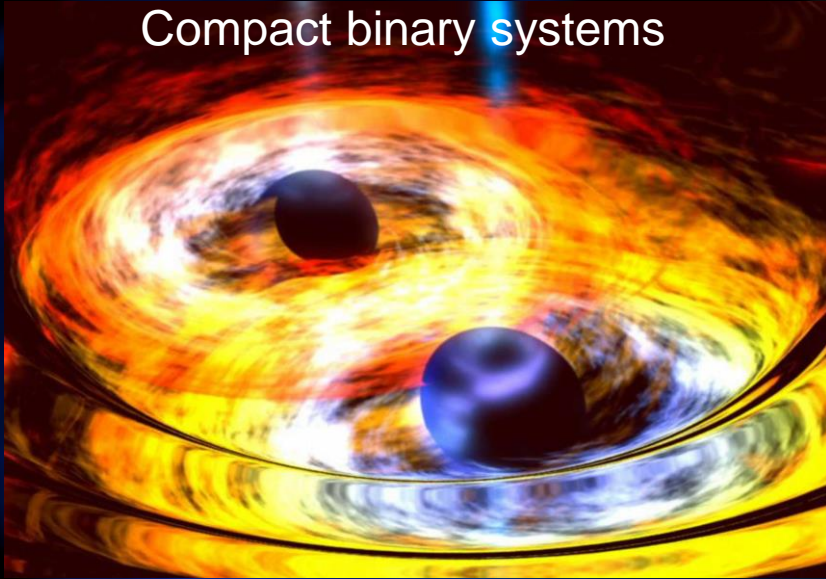
Distance Earth-Sun (1.5×10^7 km)....

...stretches by a fraction of an atom!

How do we detect them?

Sources we can probe

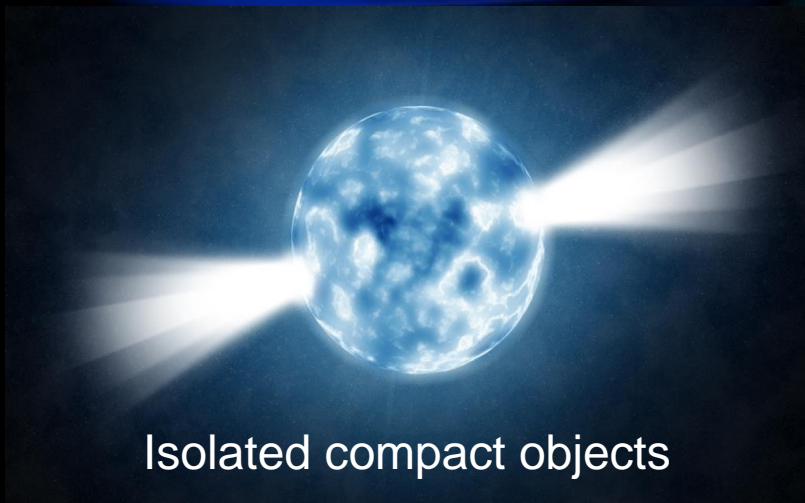
Compact binary systems



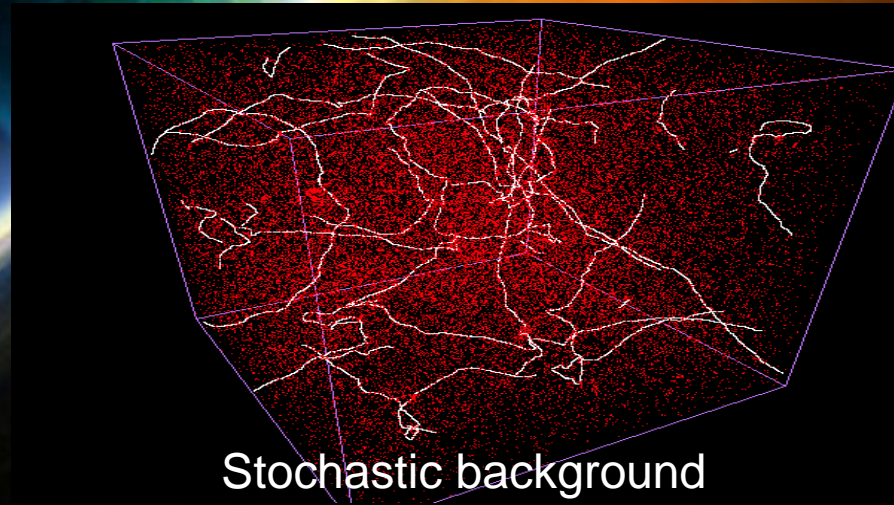
Bursts



Isolated compact objects



Stochastic background

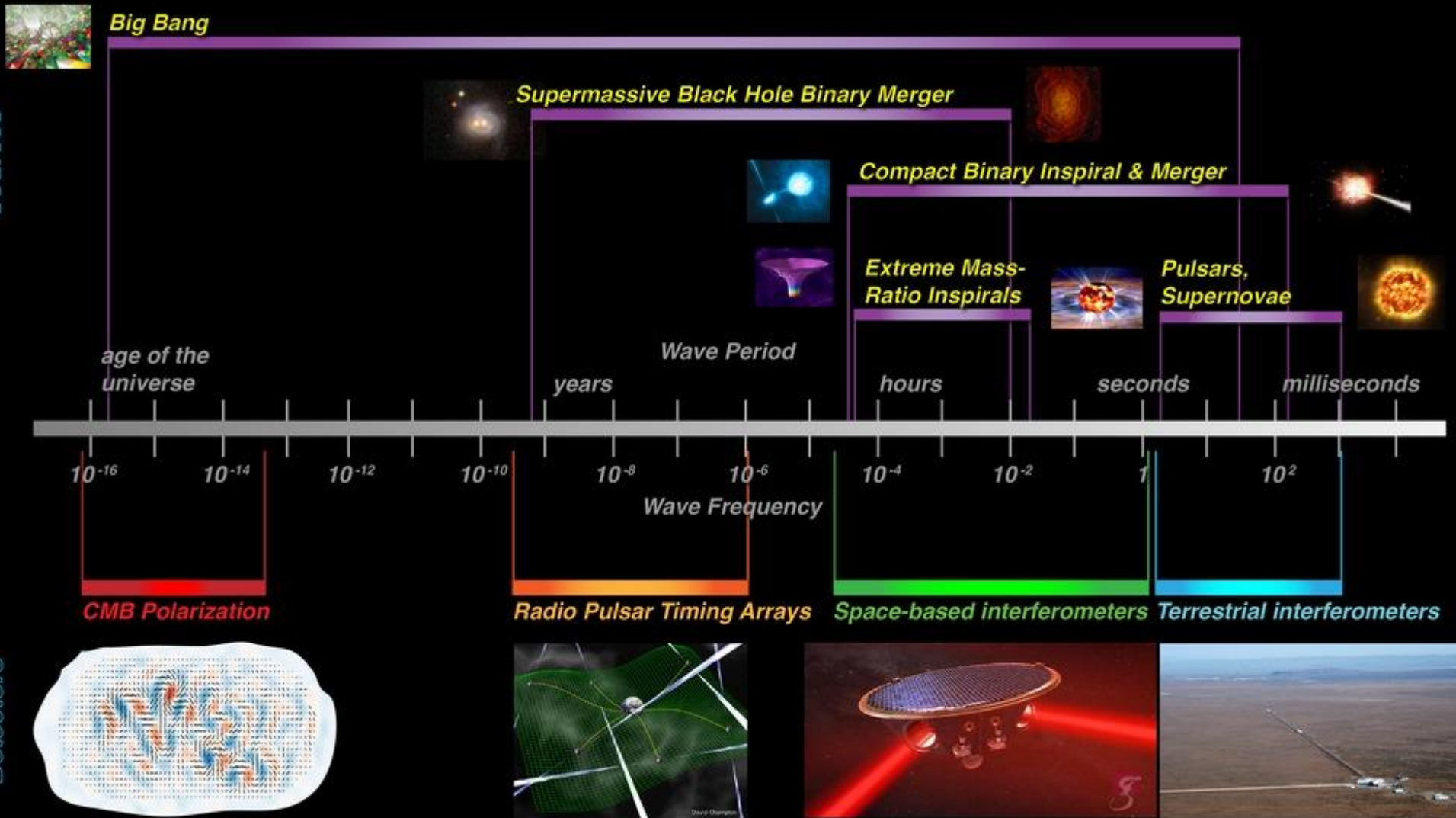


The gravitational-wave spectrum

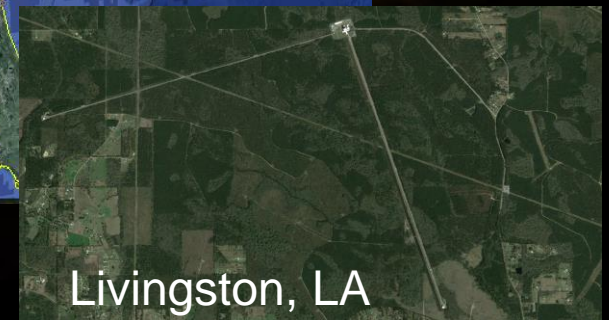
The Gravitational Wave Spectrum

Sources

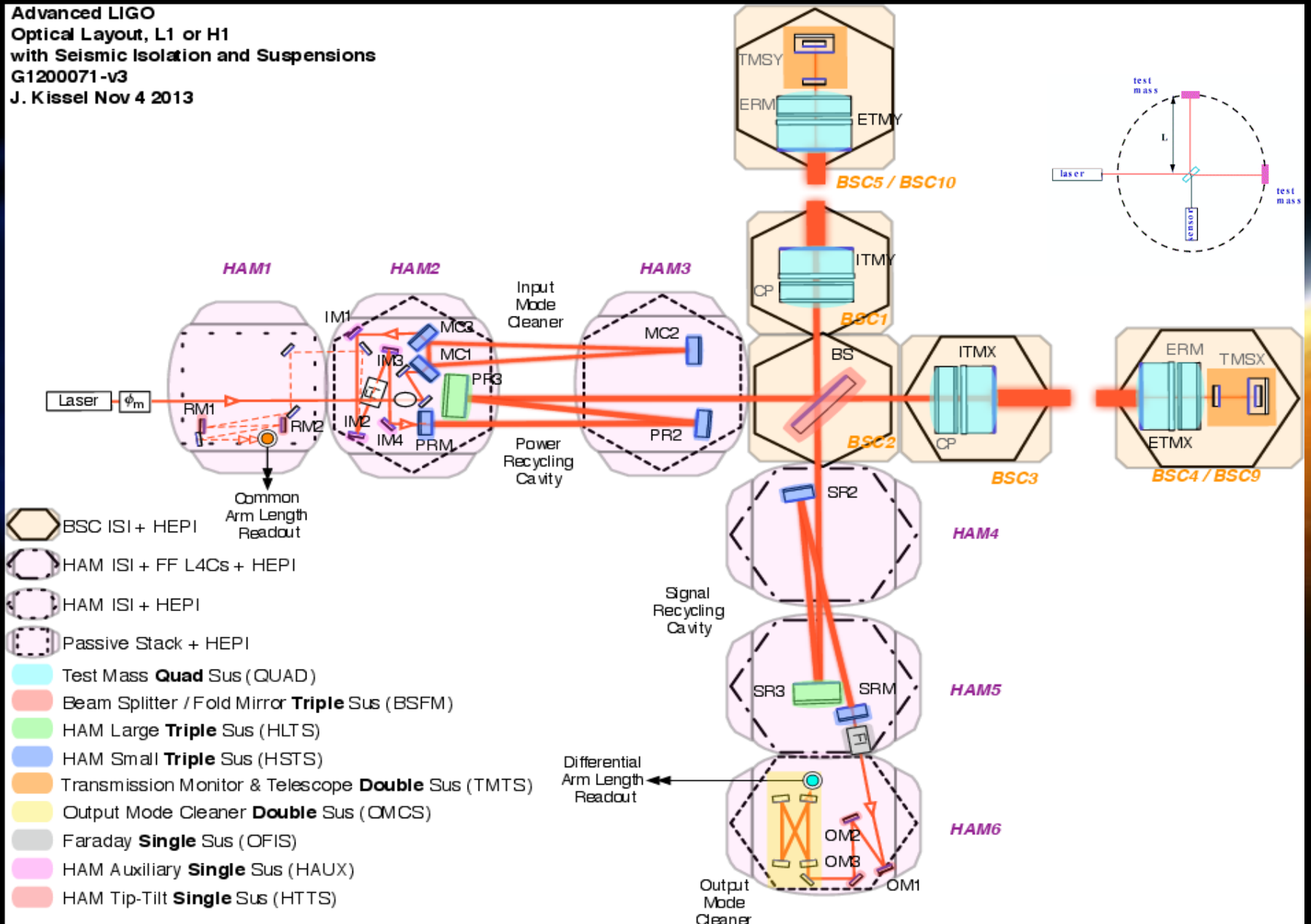
Detectors

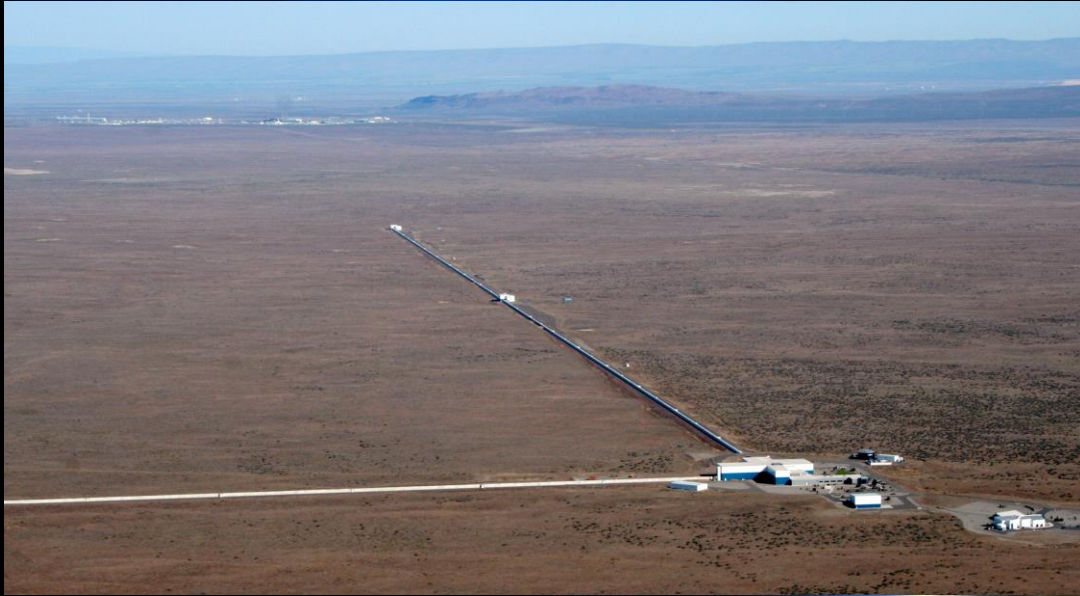


Laser Interferometer Gravitational-wave Observatory



Advanced LIGO
 Optical Layout, L1 or H1
 with Seismic Isolation and Suspensions
 G1200071-v3
 J. Kissel Nov 4 2013





What it looks like from the outside

Photo credits: LIGO Laboratory; Kai Staats.

What it looks like from the inside

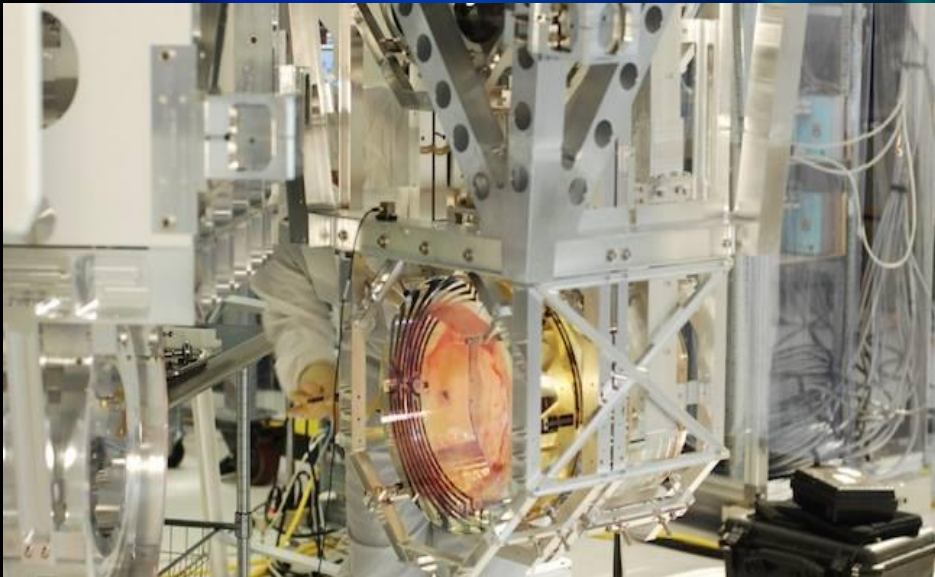
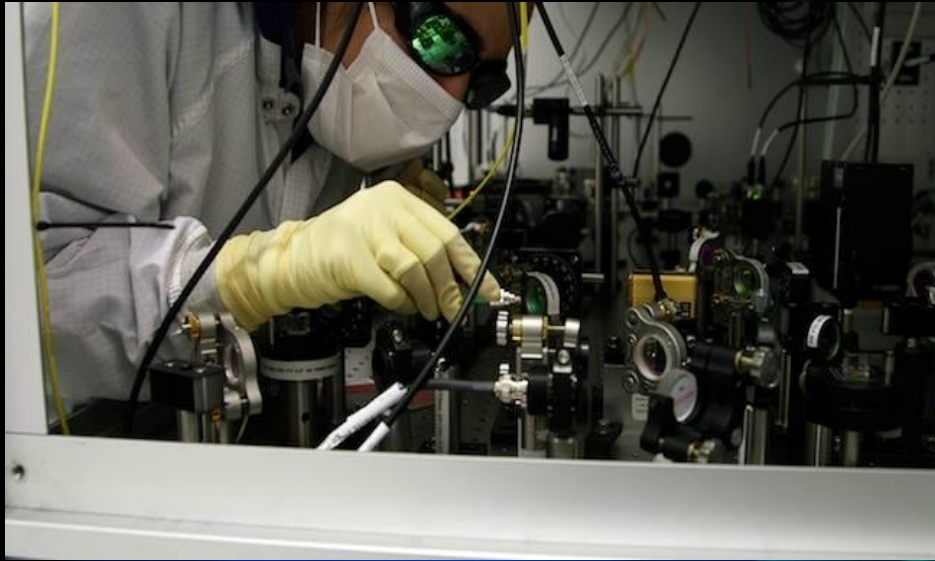


Photo credits: Kai Staats.

LIGO Scientific Collaboration

Abilene Christian University
 Albert-Einstein Institut
 Andrews University
 American University
 California Institute of Technology
 California State Univ., Fullerton
 Canadian Inst. Th. Astrophysics
 Carleton College
 Chinese University of Hong Kong
 College of William and Mary
 Columbia U. in the City of New York
 Embry-Riddle Aeronautical Univ.
 Eötvös Loránd University
 Georgia Institute of Technology
 Goddard Space Flight Center
 Hobart & William Smith Colleges
 ICTP-SAIFR
 IndIGO
 IAP-Russian Acad. of Sciences
 Inst. Nacional Pesquisas Espaciais
 Kenyon College
 Korean Gravitational-Wave Group
 Louisiana State University
 Montana State University
 Montclair State University
 Moscow State University
 National Tsinghua University
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Rochester Institute of Technology
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 University of Oregon
 University of Sannio
 Univ. of Texas-Rio Grande Valley
 University of Tokyo
 University of Washington
 University of Wisconsin-Milwaukee
 Washington State University
 West Virginia University
 Whitman College

LIGO Laboratory: California Institute of Technology, Massachusetts Institute of Technology, LIGO Hanford Observatory, LIGO Livingston Observatory

Australian Consortium for Interferometric Gravitational Astronomy (ACIGA):

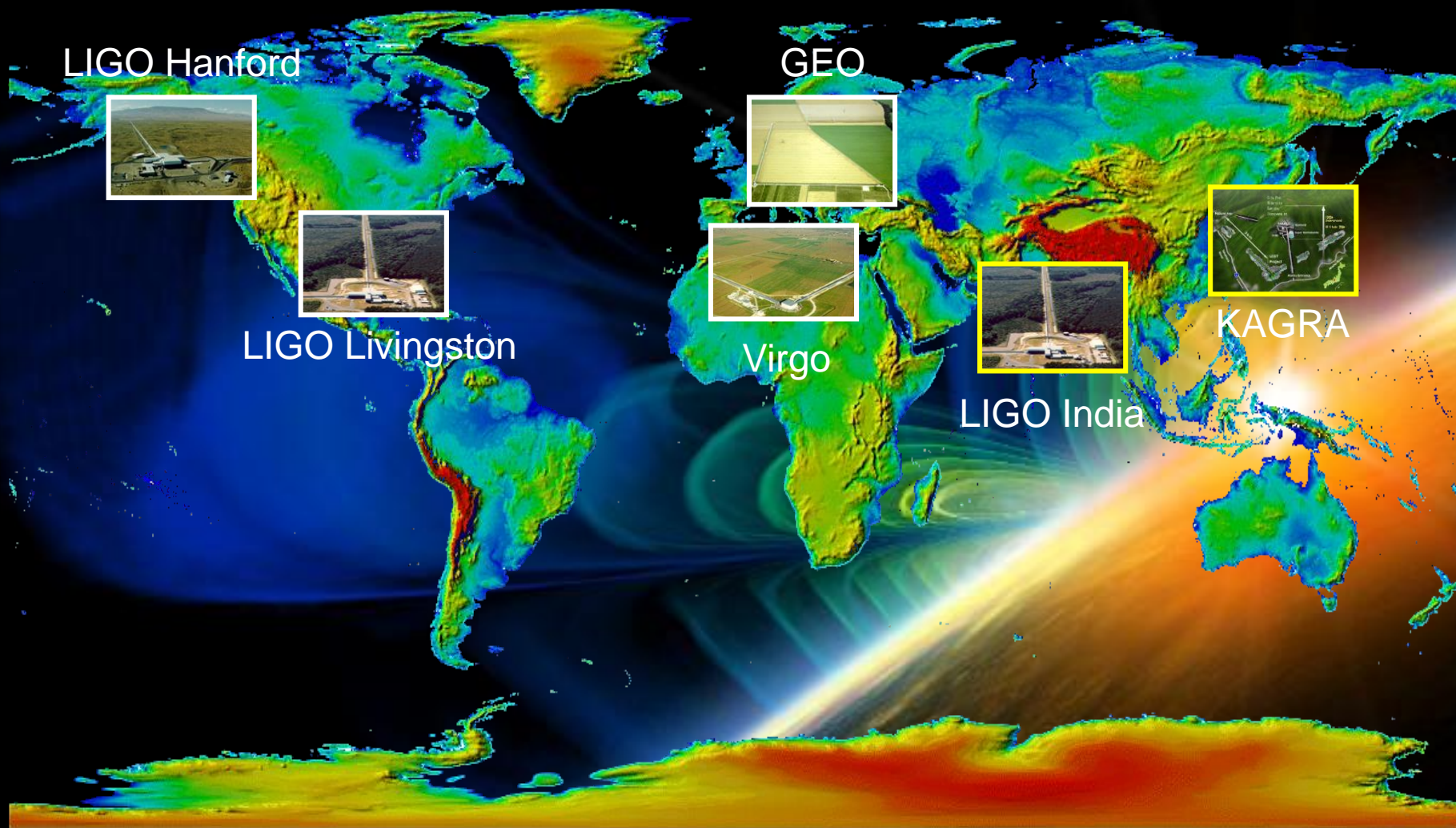
Australian National University, Charles Sturt University, Monash University, University of Adelaide, University of Melbourne, University of Western Australia

German/British Collaboration for the Detection of Gravitational Waves (GEO600):

Cardiff University, Leibniz Universität Hannover, Albert-Einstein Institut, Hannover, King's College London, University of Birmingham, University of Cambridge, University of Glasgow, University of Hamburg, University of Sheffield, University of Southampton, University of Strathclyde, University of the West of Scotland



Photo credits: Kai Staats.



LIGO Hanford



GEO



LIGO Livingston



Virgo



LIGO India



KAGRA

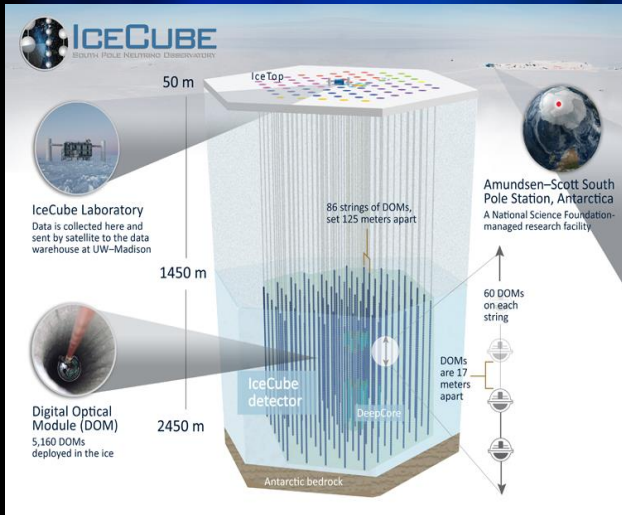
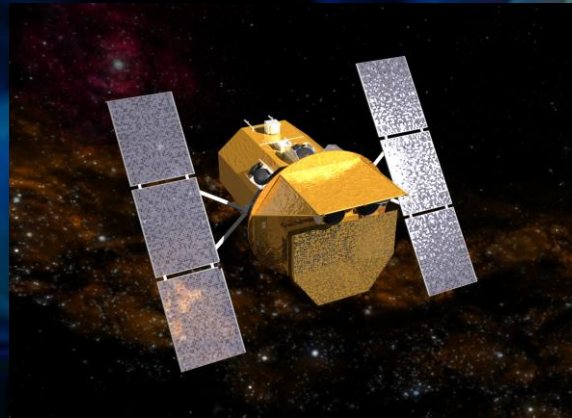


International network

Partners

75+ agreements with astronomers
for electromagnetic follow-up

MOUs with Icecube, Antares for
neutrino follow-up



The gravitational strain

Laser light travels on light cones:

$$ds^2 = -c^2 dt^2 + (\delta_{ij} + h_{ij}) dx_i dx_j = 0$$

On the x -axis, integrate from $x=0$ to $x=L$

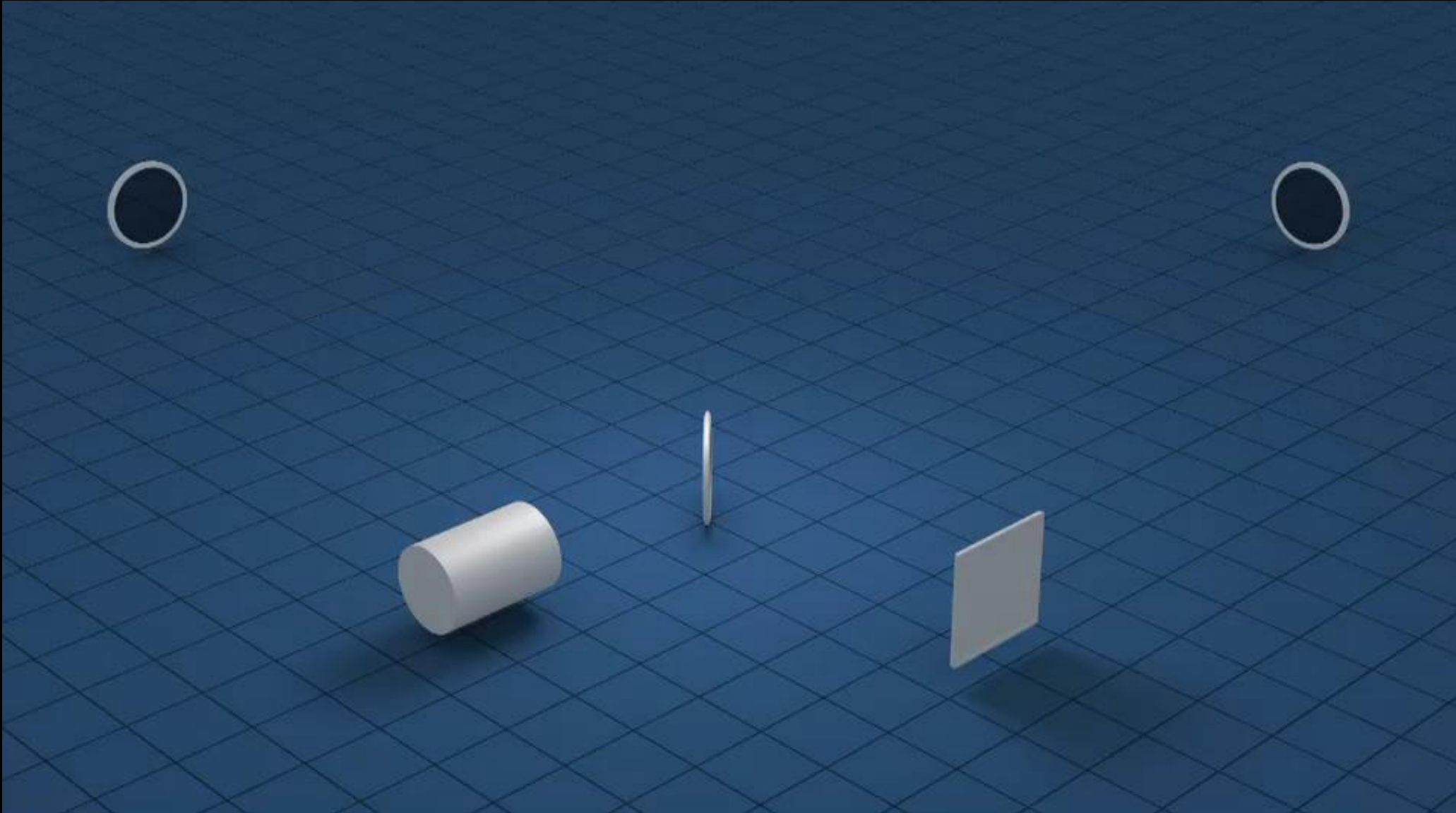
$$\Delta t = h_{11} L / 2c$$

Repeat for y -arm. Difference is ($h_{11} = -h_{22} = h$):

$$\Delta t = 2hL/c \xrightarrow[N \text{ trips}]{} \Delta t = 2hNL/c$$

$$\frac{\Delta L}{L} = \frac{h}{2}$$

← “strain”



Required sensitivity for these sources

$$\frac{\Delta L}{L} \sim 10^{-21}$$

Can we reach this precision?

If we look at on/off fringes, with a 4-km arm:

$$\Delta x \sim \lambda \sim 1 \mu\text{m} \quad \rightarrow \quad \frac{\Delta x}{L} \sim 2.5 \times 10^{-10}$$

but...

Required sensitivity for these sources

$$\frac{\Delta L}{L} \sim 10^{-21}$$

Can we reach this precision?

If we look at on/off fringes, with a 4-km arm:

$$\Delta x \sim \lambda \sim 1 \mu\text{m} \quad \rightarrow \quad \frac{\Delta x}{L} \sim 2.5 \times 10^{-10}$$

...first build up power in the arms

Required sensitivity for these sources

$$\frac{\Delta L}{L} \sim 10^{-21}$$

Can we reach this precision?

If we look at on/off fringes, with a 4-km arm:

$$\Delta x \sim \lambda \sim 1 \mu\text{m} \rightarrow \frac{\Delta x}{L_{\text{eff}}} \sim 10^{-12}$$

then...

Average flux of photons

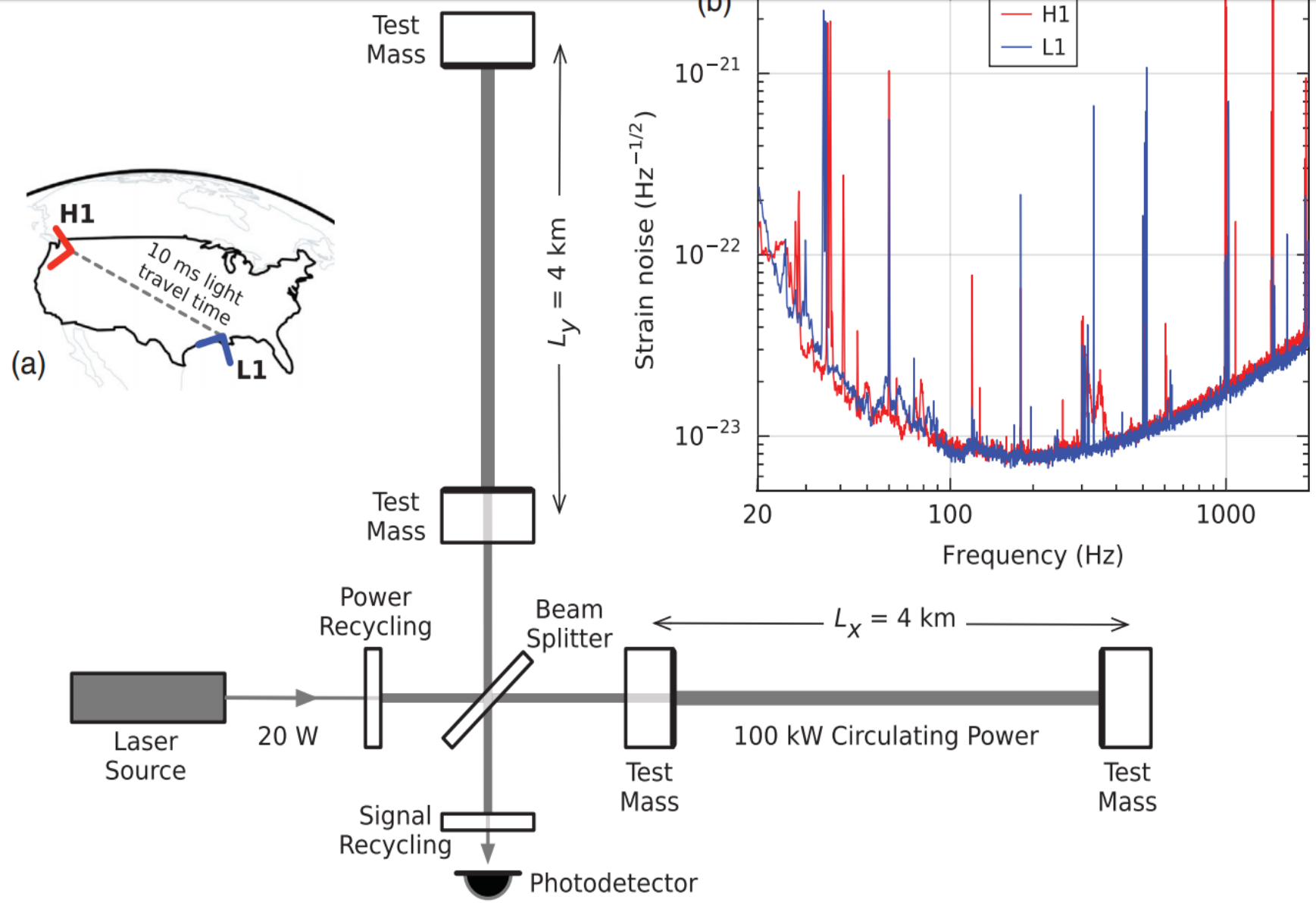
$$\bar{N} = \frac{\lambda}{2\pi\hbar c} P$$

Fluctuations (shot noise):

$$\frac{\Delta N}{N} = \frac{1}{\sqrt{N}}$$

200 W of laser light carries 10^{19} photons per second, giving a sensitivity of

$$\frac{\Delta N}{N} \sim 3 \times 10^{-10} \quad \rightarrow \quad h \sim \frac{\Delta x}{L_{\text{eff}}} \frac{\Delta N}{N} \sim 3 \times 10^{-22}$$

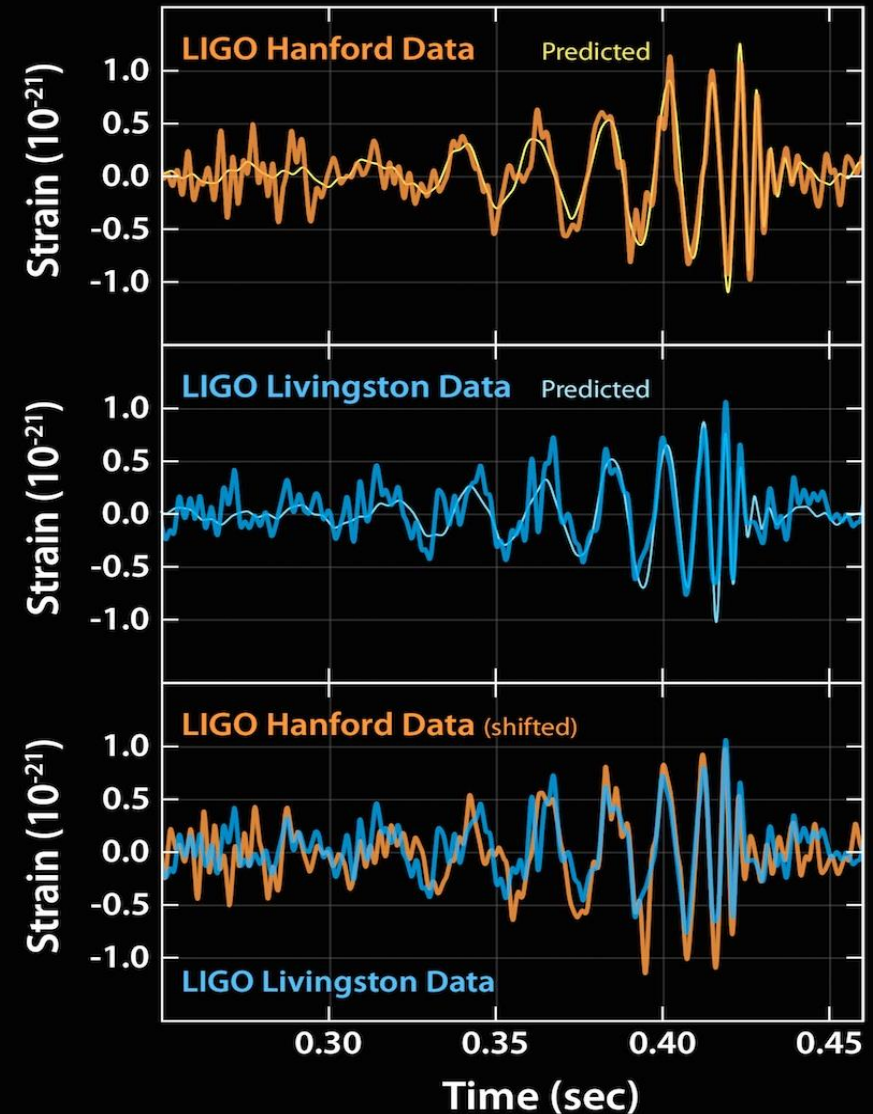


B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration) Phys. Rev. Lett. **116**, 061102

LIGO's first detection

GW150914: September 14, 2015, 9:50:45 UTC

- 7 ms delay between arrival times at Livingston and Hanford
- Filtered to remove low frequency fluctuations and single frequency lines
- Hanford's data inverted to account for orientation of detectors



PRL 116, 061102 (2016)

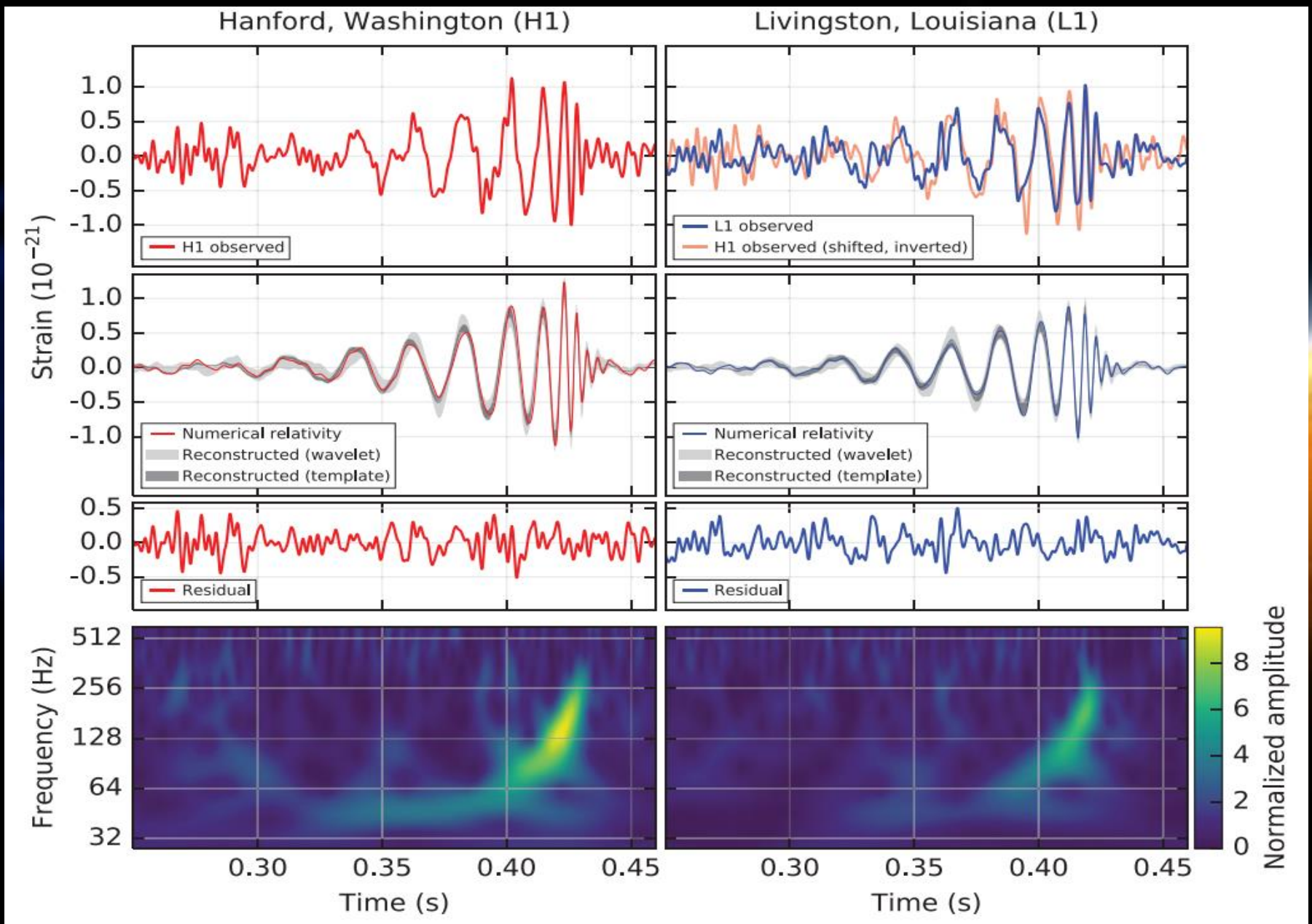
PHYSICAL REVIEW LETTERS



Observation of Gravitational Waves from a Binary Black Hole Merger

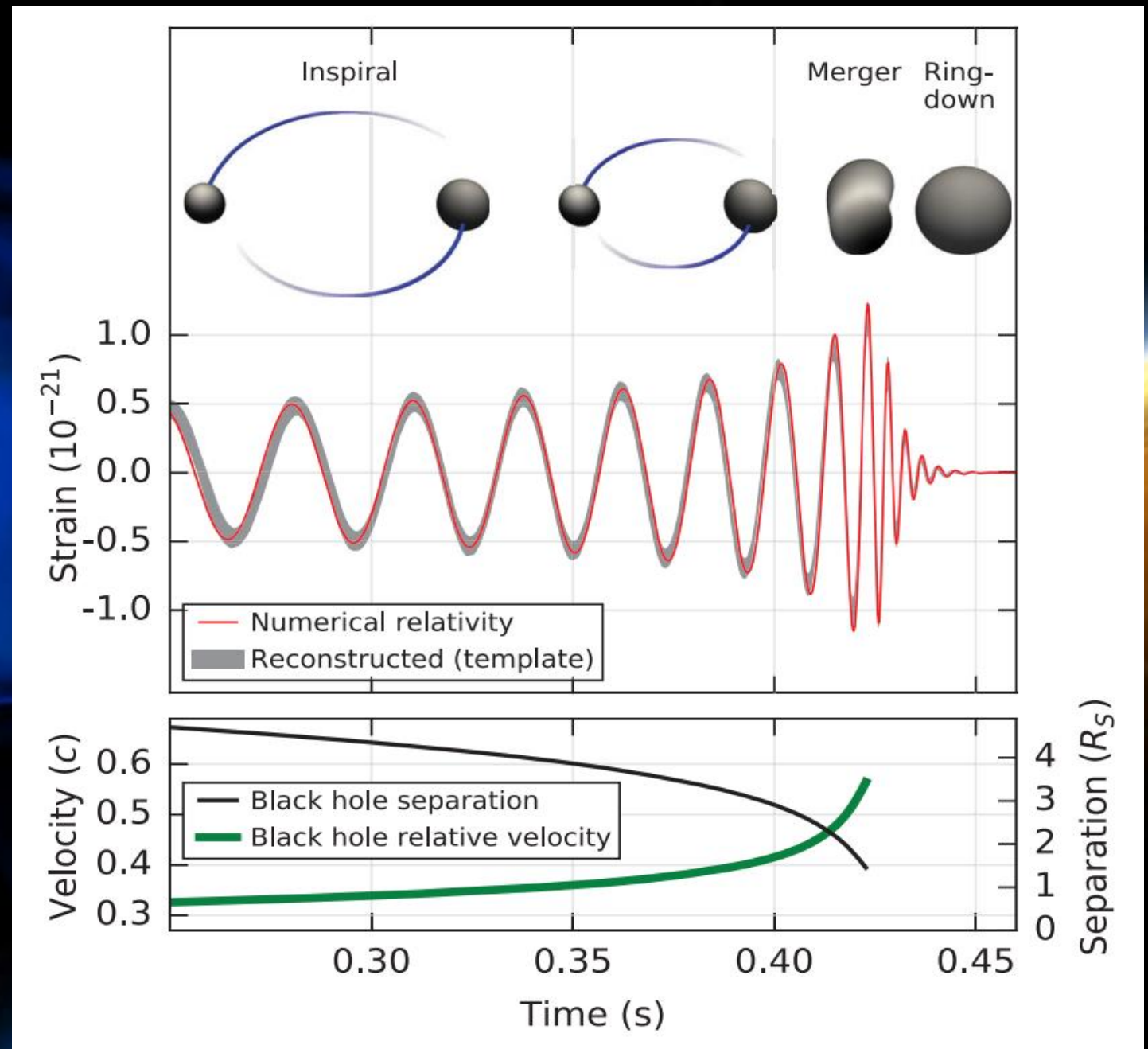
B. P. Abbott *et al.**

(LIGO Scientific Collaboration and Virgo Collaboration)
(Received 21 January 2016; published 11 February 2016)

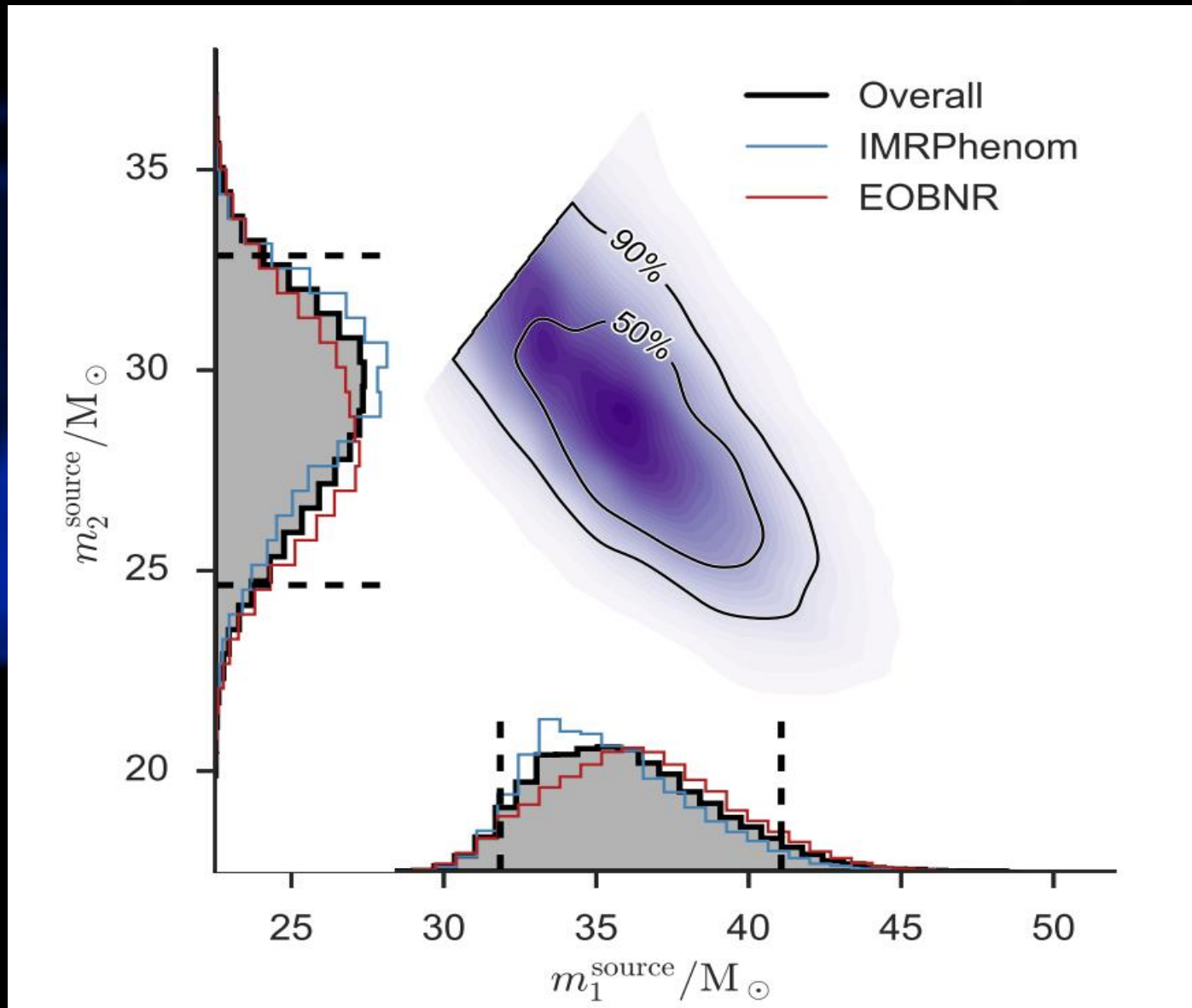


B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration) Phys. Rev. Lett. **116**, 061102

Inspiral: low velocity and weak gravitational field.
 Late inspiral/plunge: high velocity and strong gravitational field.
 Merger: nonlinear and non perturbative effects.
 Ringdown: excitation of quasinormal modes

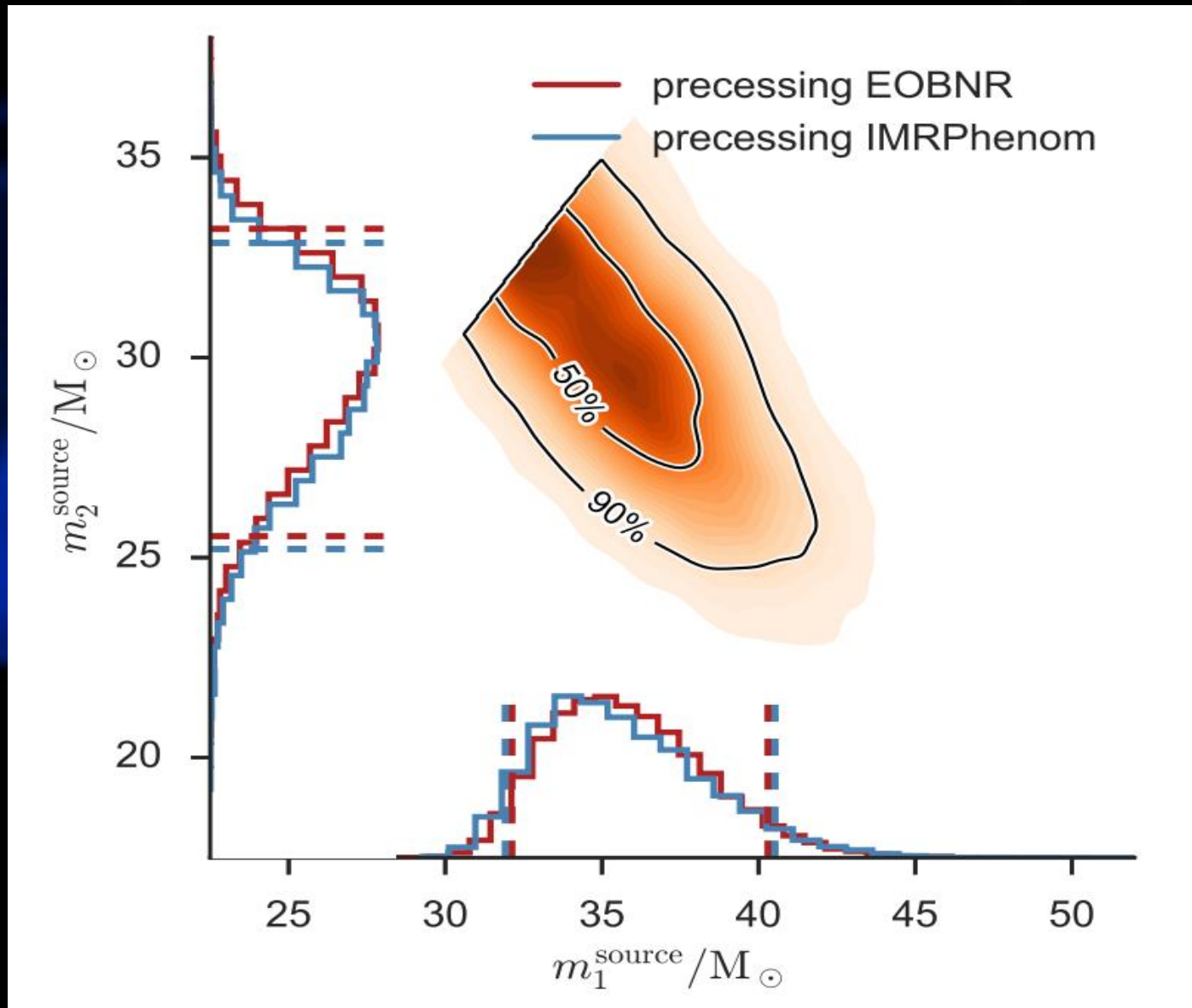


Black hole masses



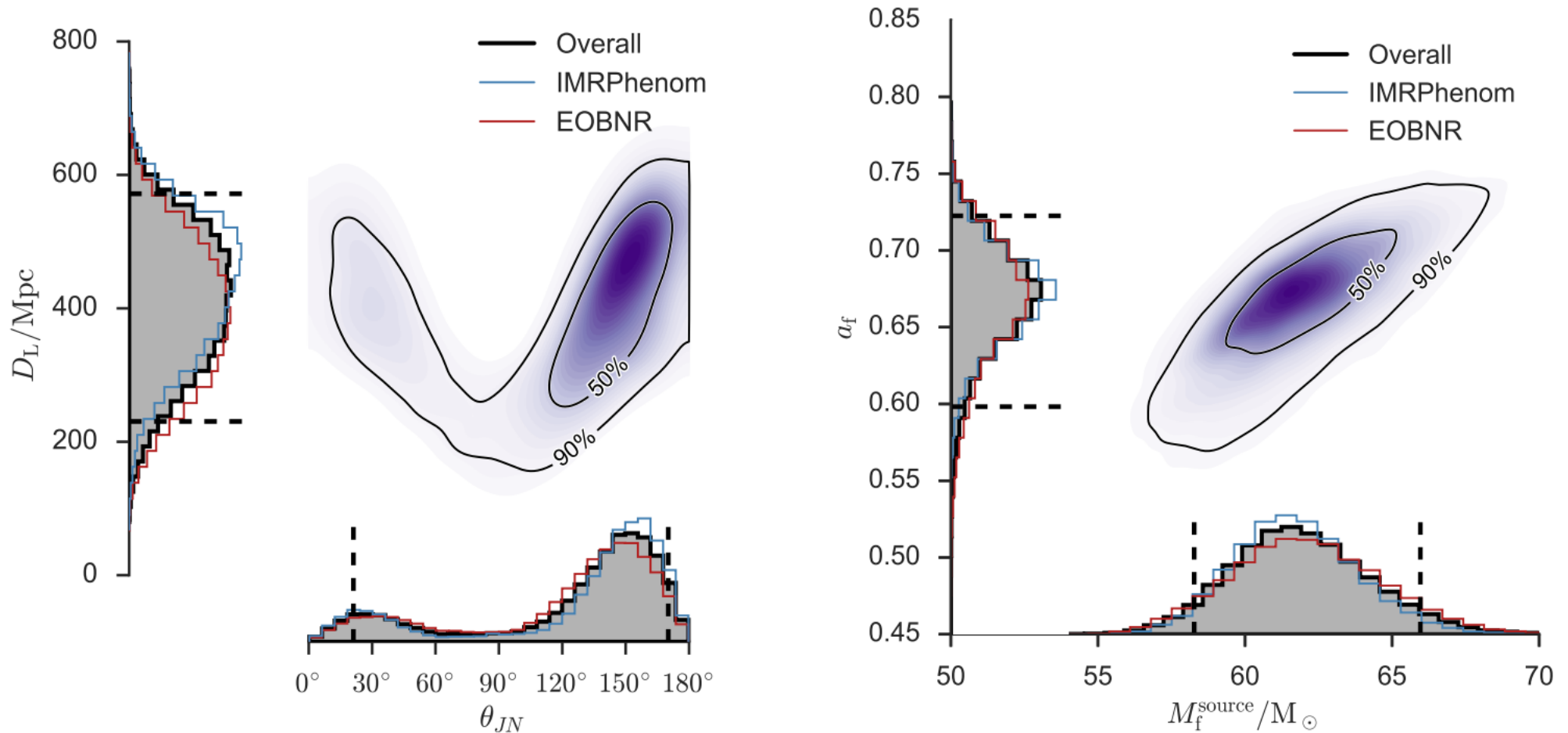
B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration) Phys. Rev. Lett. **116**, 061102

Black hole masses



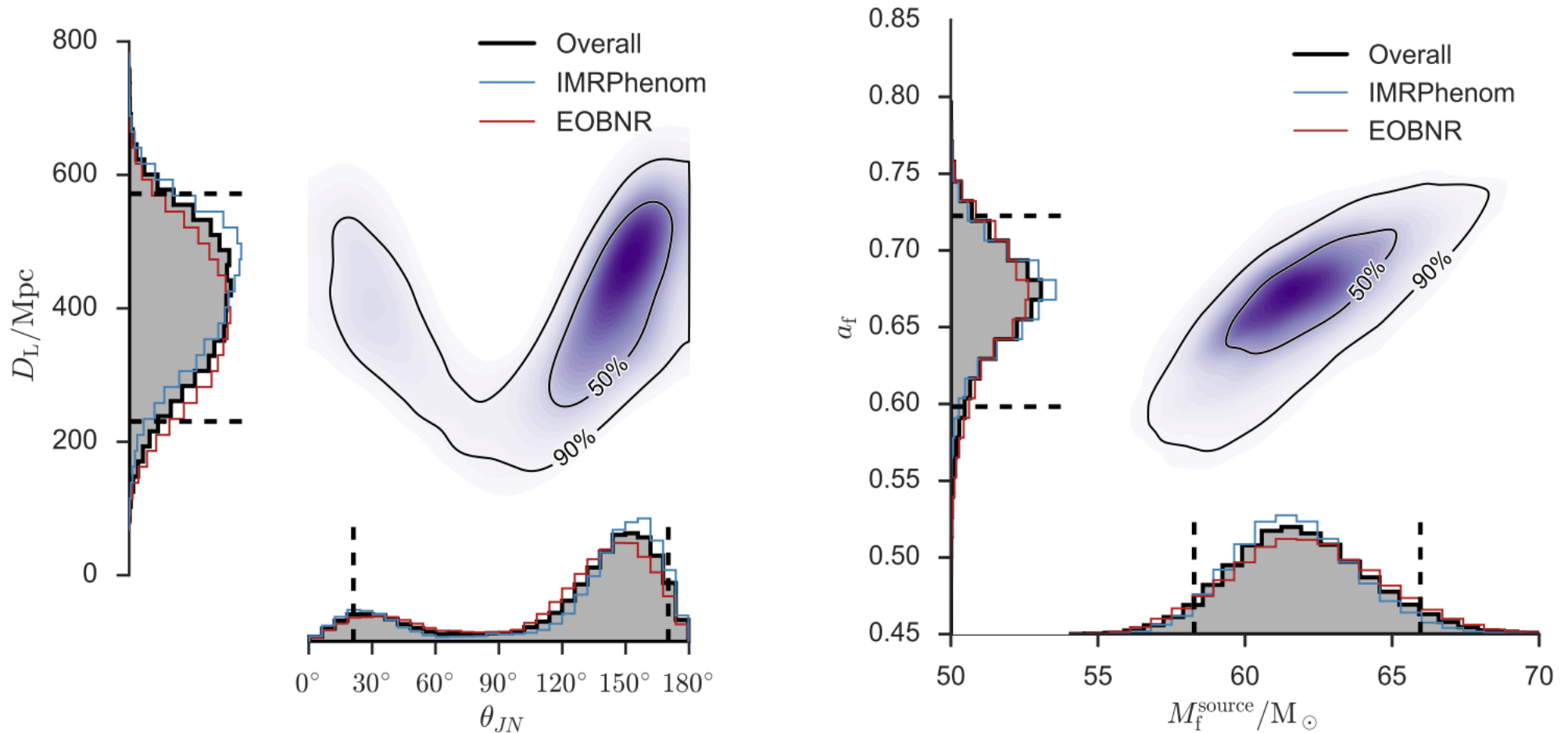
B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration) arXiv 1606.01210.

Distance and final black hole



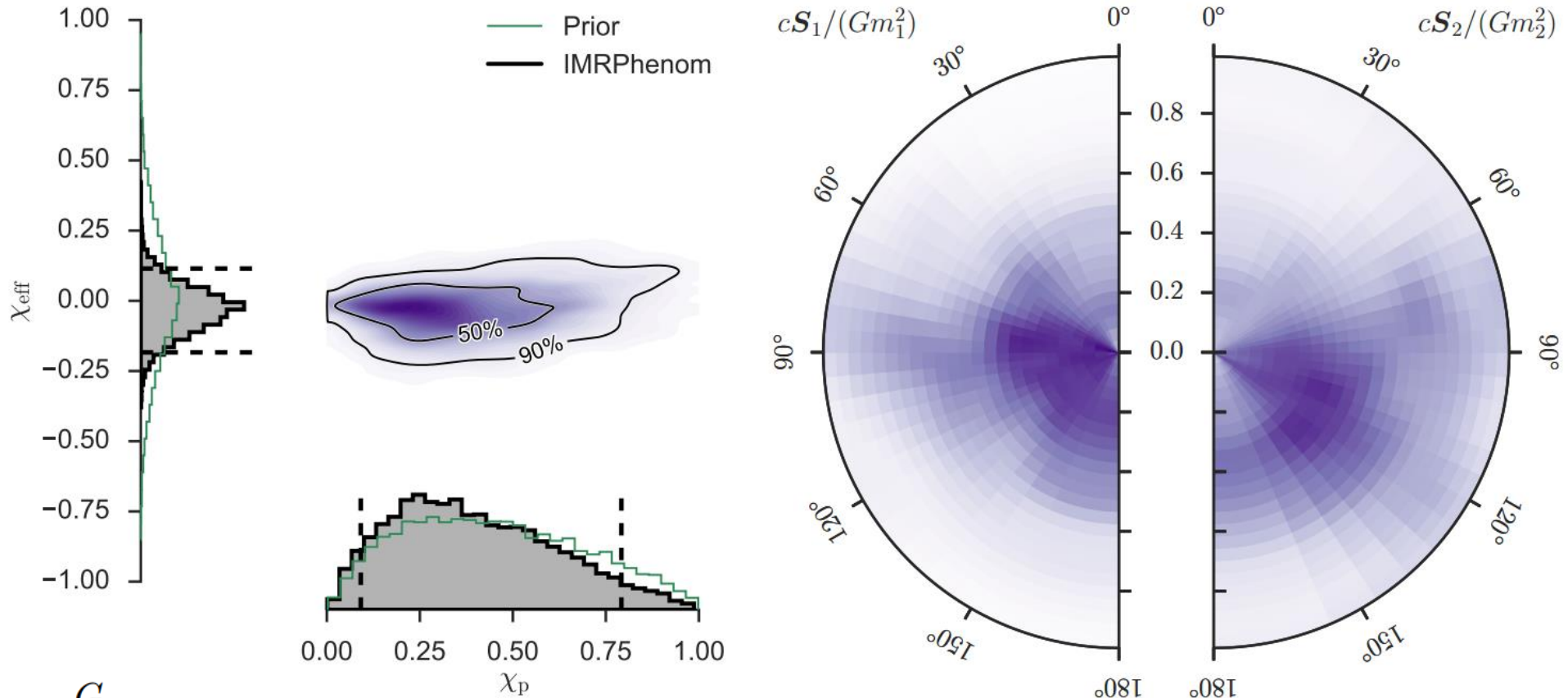
B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration) Phys. Rev. Lett. **116**, 061102

Distance and final black hole



B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration) Phys. Rev. Lett. **116**, 061102

Spins



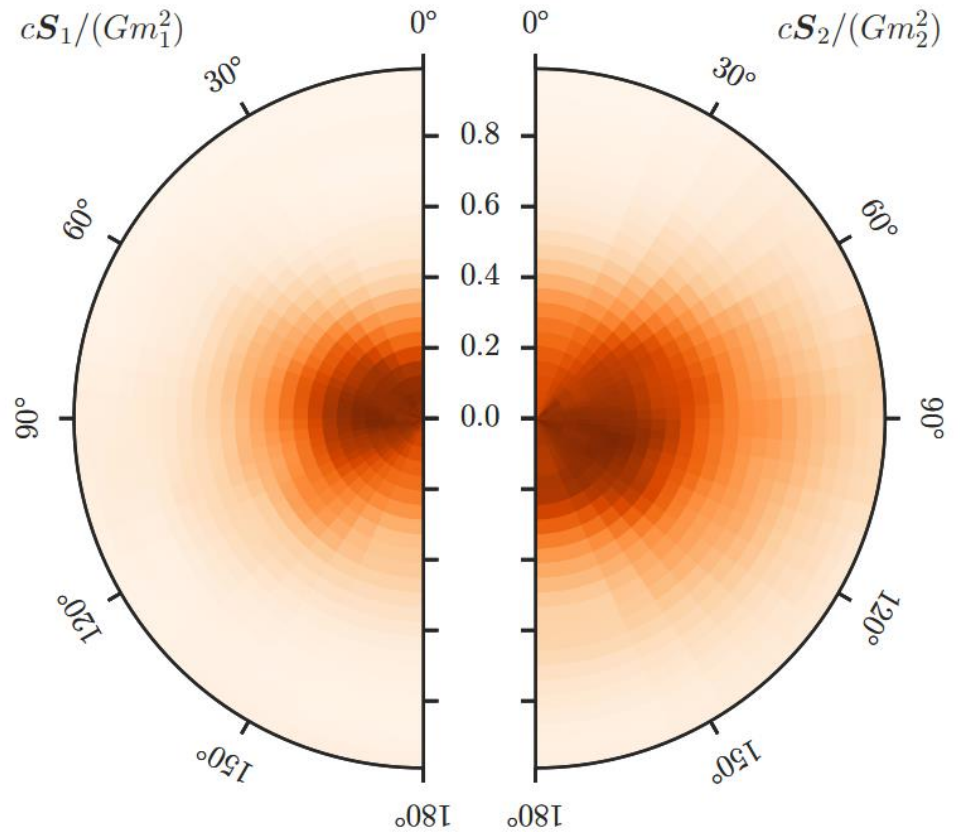
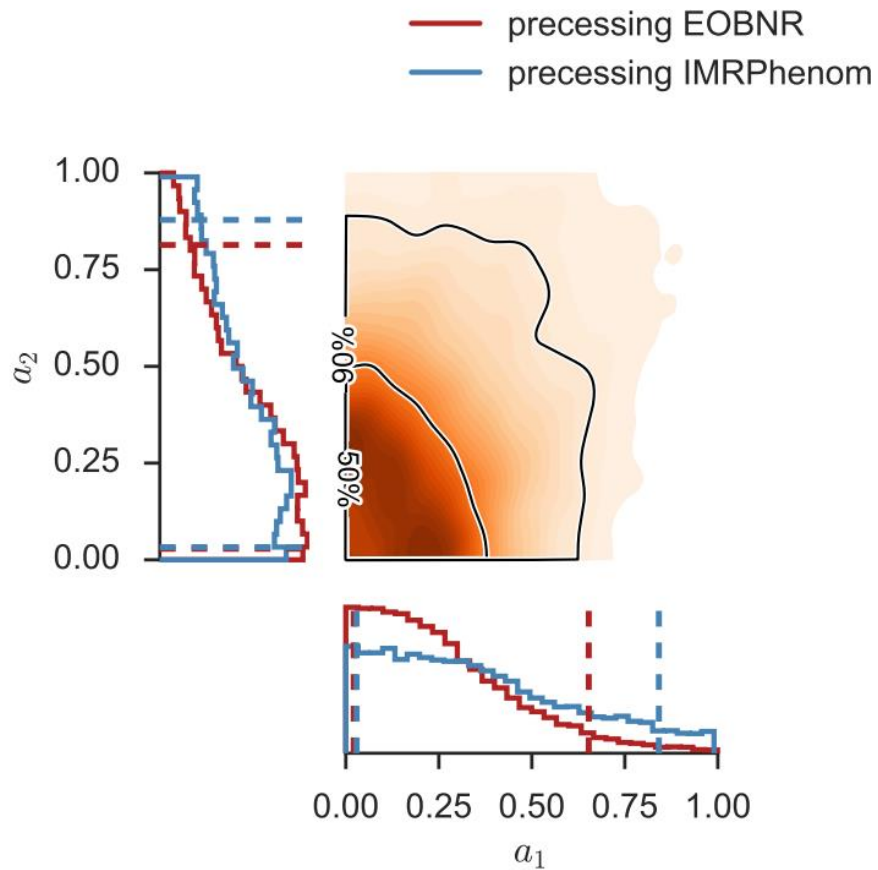
$$\dot{\mathbf{L}} = \frac{G}{c^2 r^3} (B_1 \mathbf{S}_{1\perp} + B_2 \mathbf{S}_{2\perp}) \times \mathbf{L}$$

$$\dot{\mathbf{S}}_i = \frac{G}{c^2 r^3} B_i \mathbf{L} \times \mathbf{S}_i,$$

$$\chi_{\text{eff}} = \frac{c}{G} \left(\frac{\mathbf{S}_1}{m_1} + \frac{\mathbf{S}_2}{m_2} \right) \cdot \frac{\hat{\mathbf{L}}}{M} \quad \chi_p = \frac{c}{B_1 G m_1^2} \max(B_1 S_{1\perp}, B_2 S_{2\perp}) > 0$$

B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration) Phys. Rev. Lett. **116**, 061102

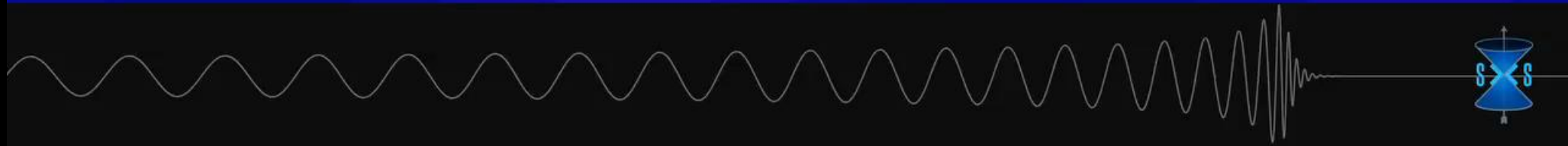
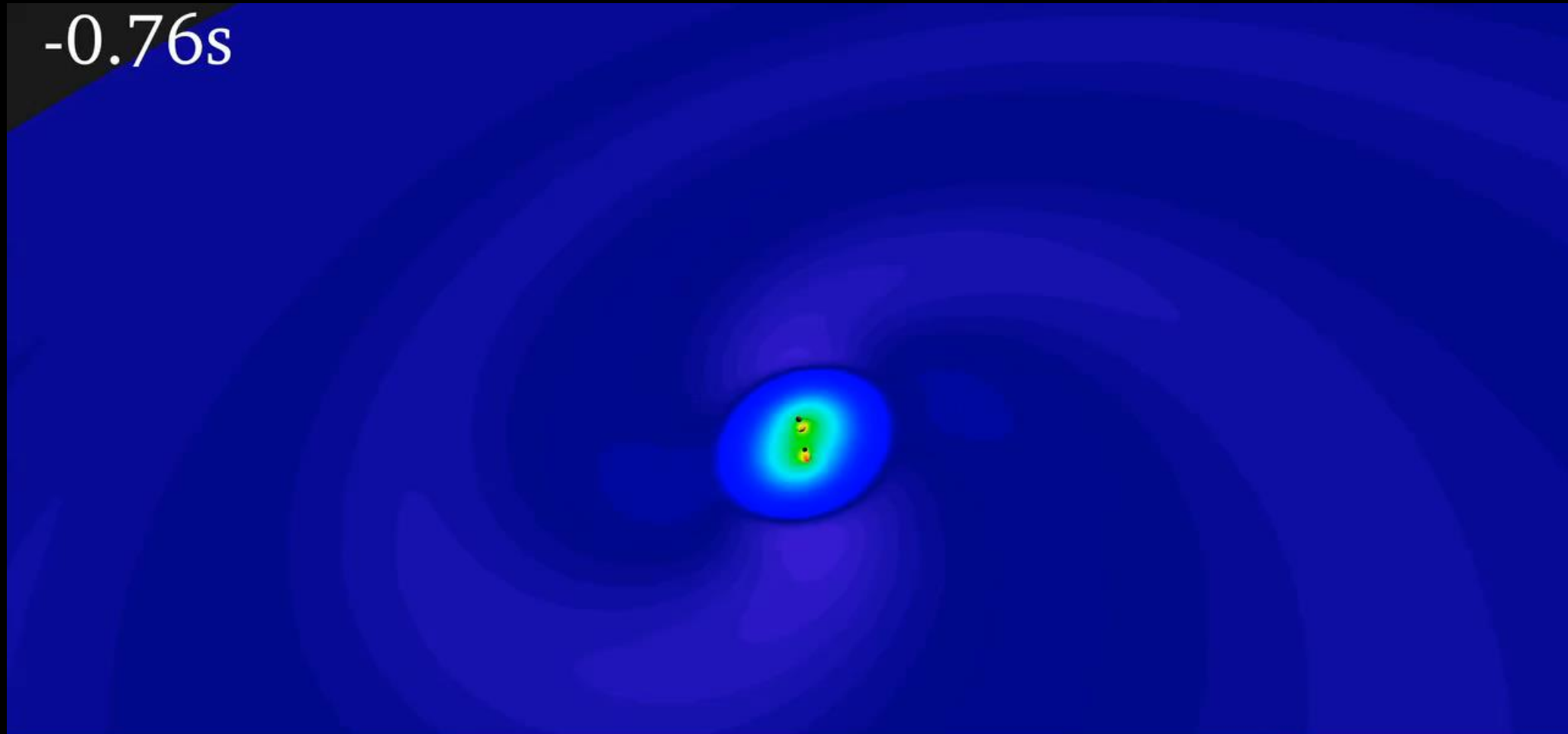
Spins



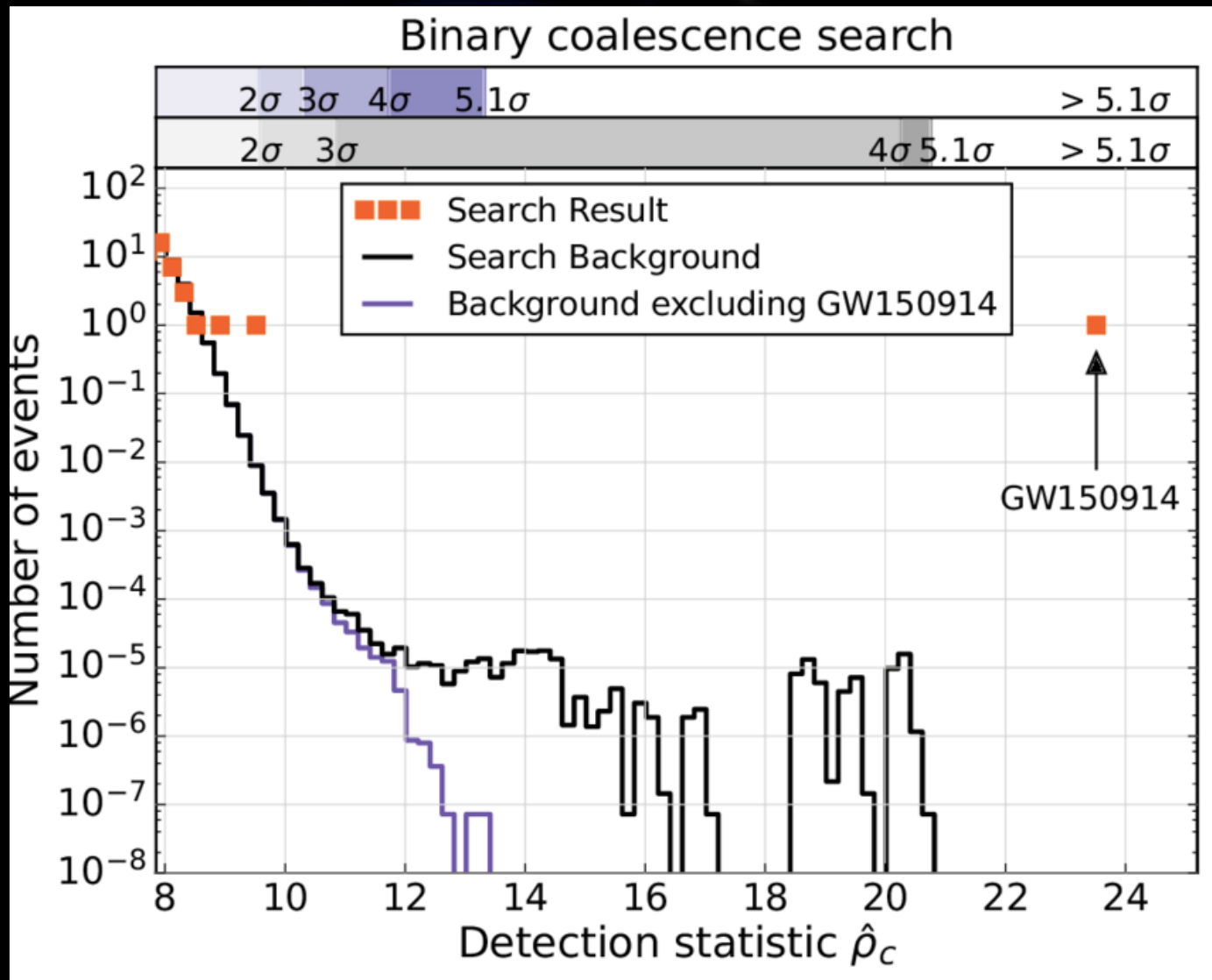
B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration) arXiv 1606.01210.

observed by	LIGO L1, H1	duration from 30 Hz	~ 200 ms
source type	black hole (BH) binary	# cycles from 30 Hz	~10
date	14 Sept 2015	peak GW strain	1×10^{-21}
time	09:50:45 UTC	peak displacement of interferometers arms	± 0.002 fm
likely distance	0.75 to 1.9 Gly 230 to 570 Mpc	frequency/wavelength at peak GW strain	150 Hz, 2000 km
redshift	0.054 to 0.136	peak speed of BHs	~ 0.6 c
signal-to-noise ratio	24	peak GW luminosity	3.6×10^{56} erg s ⁻¹
false alarm prob.	< 1 in 5 million	radiated GW energy	2.5-3.5 M _⊙
false alarm rate	< 1 in 200,000 yr	remnant ringdown freq.	~ 250 Hz
Source Masses	M _⊙	remnant damping time	~ 4 ms
total mass	60 to 70	remnant size, area	180 km, 3.5×10^5 km ²
primary BH	32 to 41	consistent with general relativity?	passes all tests performed
secondary BH	25 to 33	graviton mass bound	< 1.2×10^{-22} eV
remnant BH	58 to 67	coalescence rate of binary black holes	2 to 400 Gpc ⁻³ yr ⁻¹
mass ratio	0.6 to 1	online trigger latency	~ 3 min
primary BH spin	< 0.7	# offline analysis pipelines	5
secondary BH spin	< 0.9	CPU hours consumed	~ 50 million (=20,000 PCs run for 100 days)
remnant BH spin	0.57 to 0.72	papers on Feb 11, 2016	13
signal arrival time delay	arrived in L1 7 ms before H1	# researchers	~1000, 80 institutions in 15 countries
likely sky position	Southern Hemisphere		
likely orientation resolved to	face-on/off ~600 sq. deg.		

-0.76s

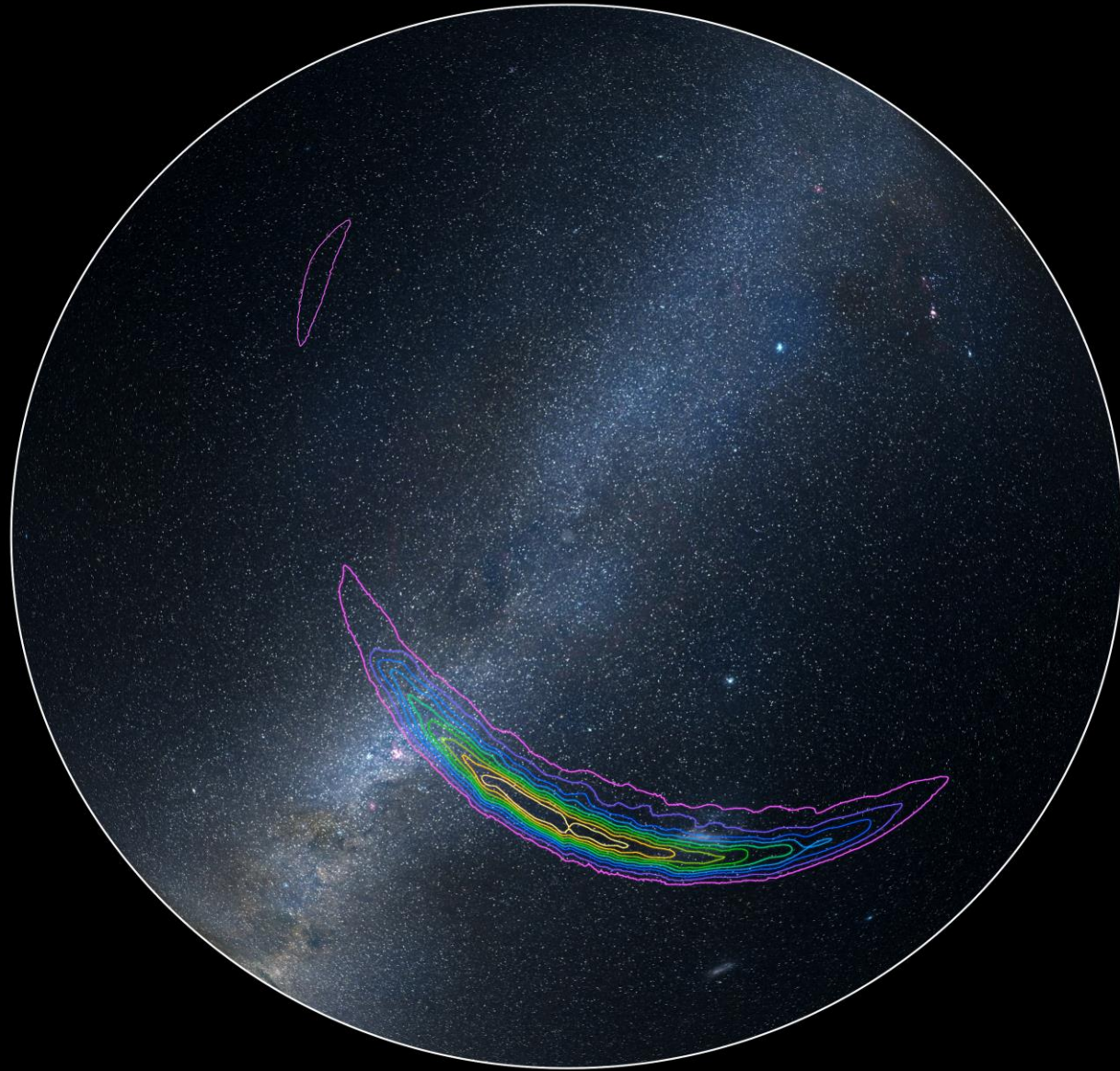


GW150914 statistical significance

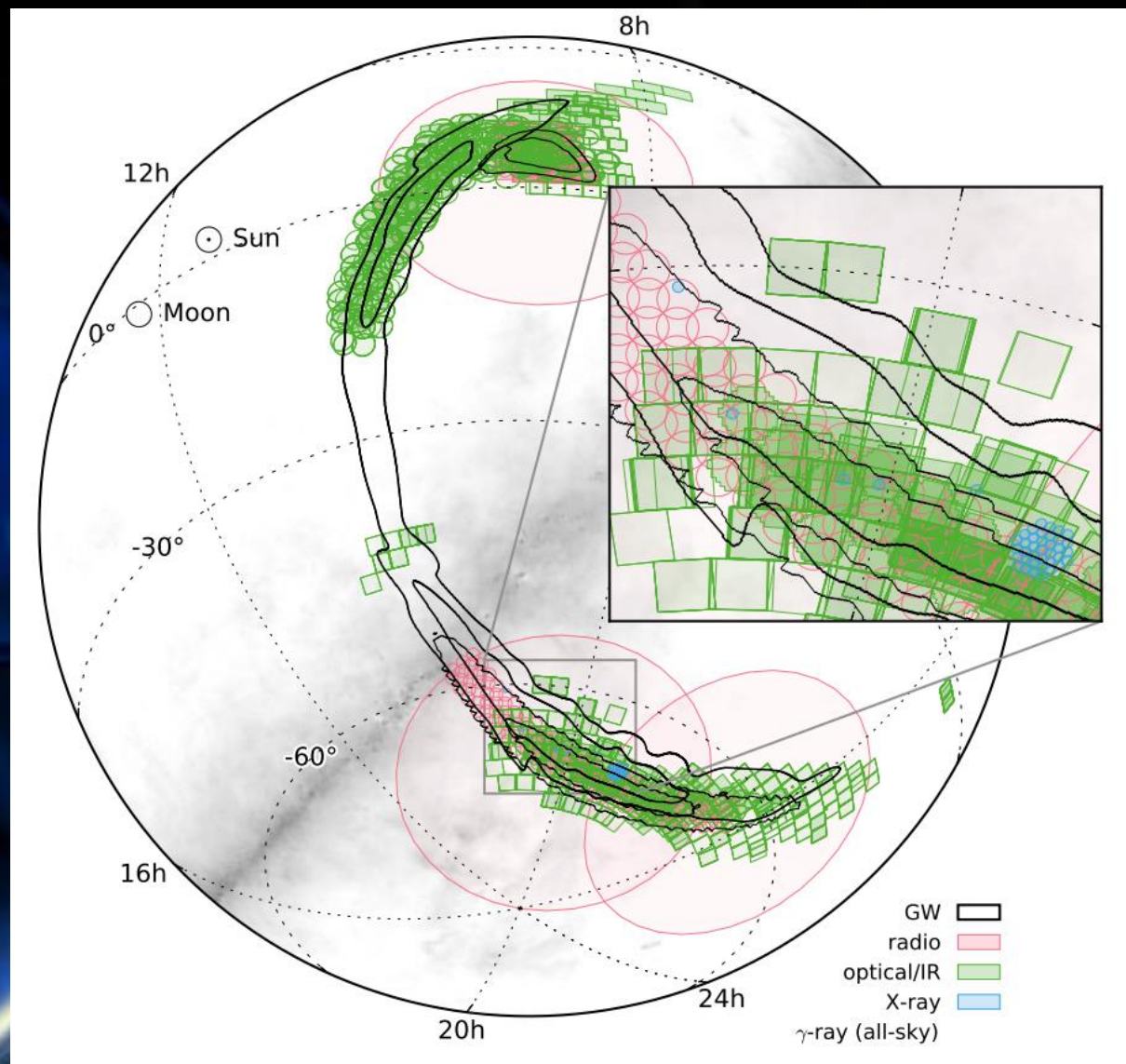
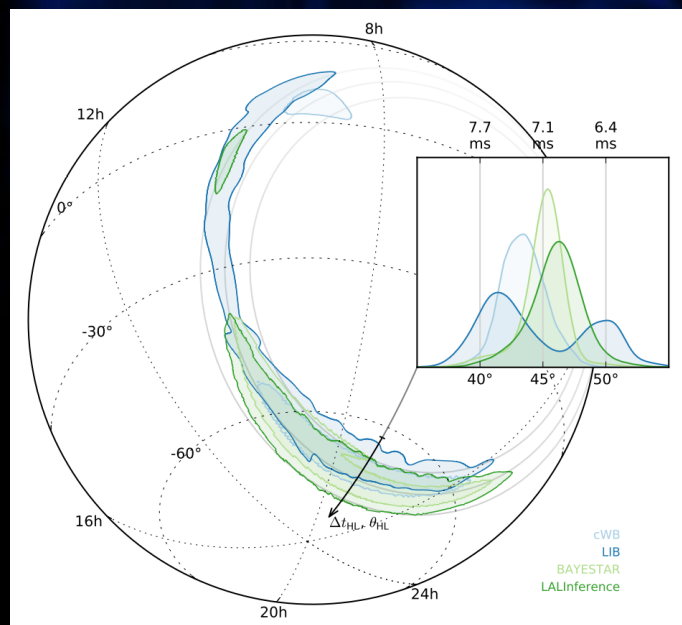


- False alarm rate is less than 1 event per 203,000 years
- A significance of > 5.1 sigma

Where did it come from?



Skymap sent for EM follow-up



B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration + Astronomers) arXiv:1602.08492

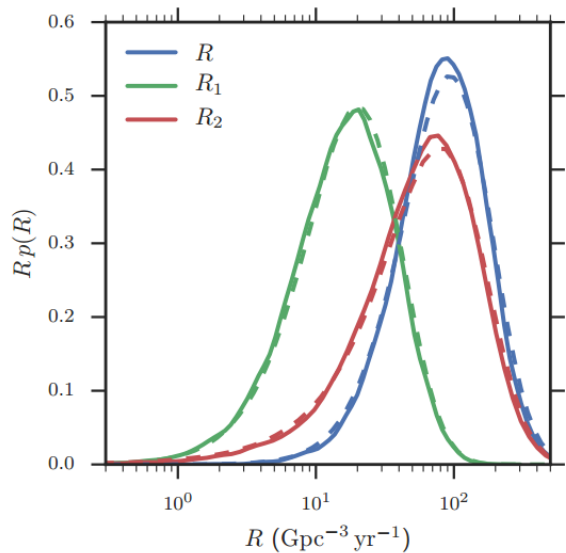
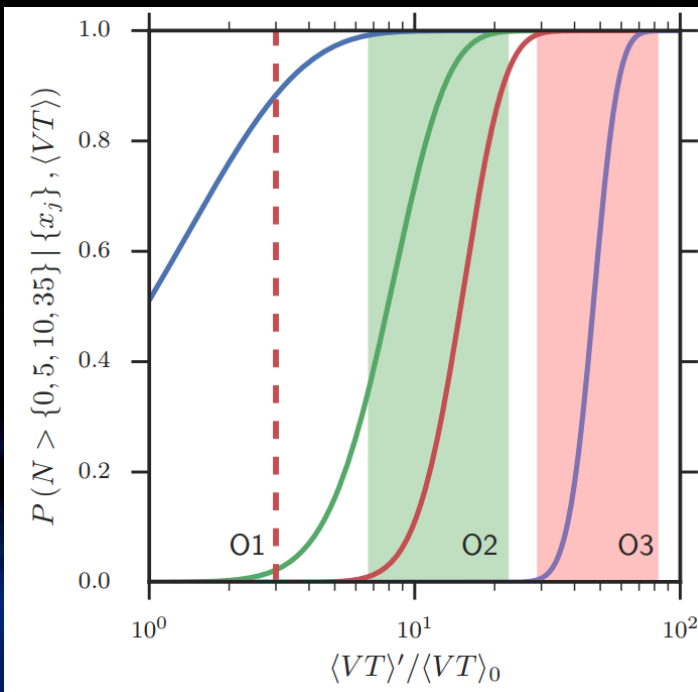
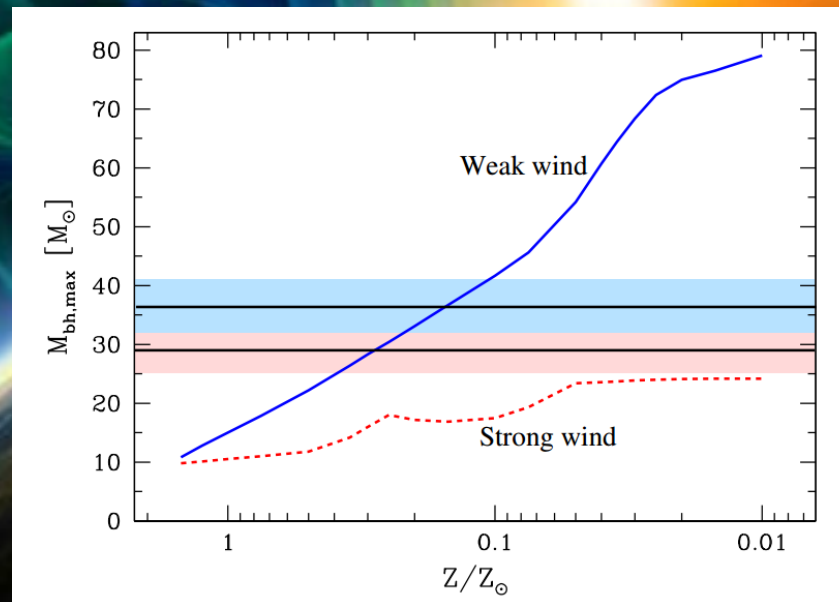


Figure 4. The posterior density on the rate of GW150914-like BBH inspirals, R_1 (green), LVT151012-like BBH inspirals, R_2 (red), and the inferred total rate, $R = R_1 + R_2$ (blue). The median and 90% credible levels are given in Table 1. Solid lines give the rate inferred from the pycbc trigger set, while dashed lines give the rate inferred from the gstlal trigger set.

Mass Distribution	$R / (\text{Gpc}^{-3} \text{yr}^{-1})$		
	pycbc	gstlal	Combined
GW150914	16^{+38}_{-13}	17^{+39}_{-14}	17^{+39}_{-13}
LVT151012	61^{+152}_{-53}	62^{+164}_{-55}	62^{+165}_{-54}
Both	82^{+155}_{-61}	84^{+172}_{-64}	83^{+168}_{-63}
Astrophysical			
Flat	33^{+64}_{-26}	32^{+65}_{-25}	33^{+62}_{-26}
Power Law	102^{+198}_{-79}	99^{+203}_{-79}	100^{+201}_{-79}

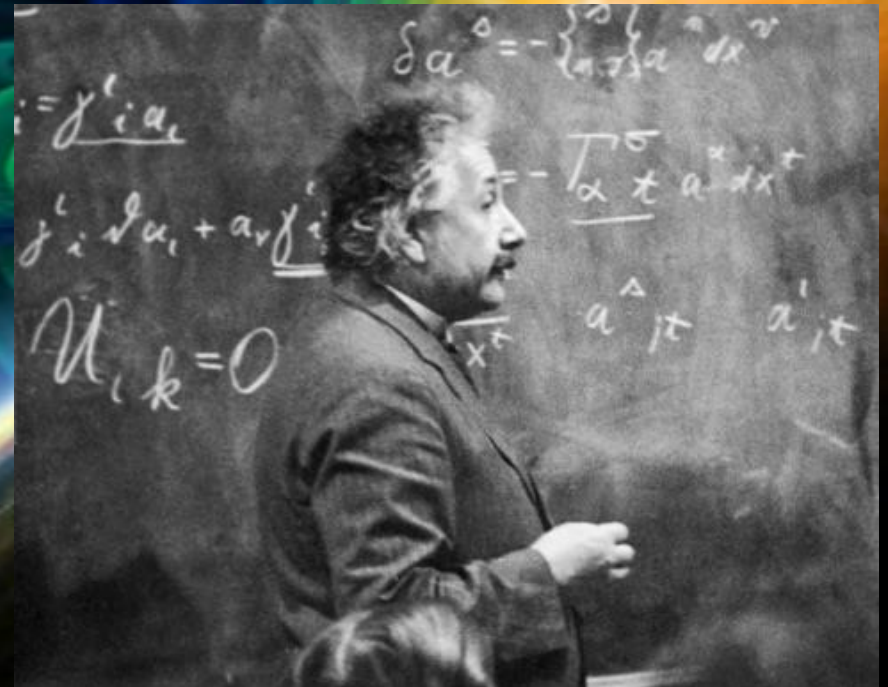


Astrophysical implications



General Relativity tests

- Inspiral, merger and ringdown consistency tests
- Tests of QNMs
- Deviations from GR waveforms
- Graviton Compton length

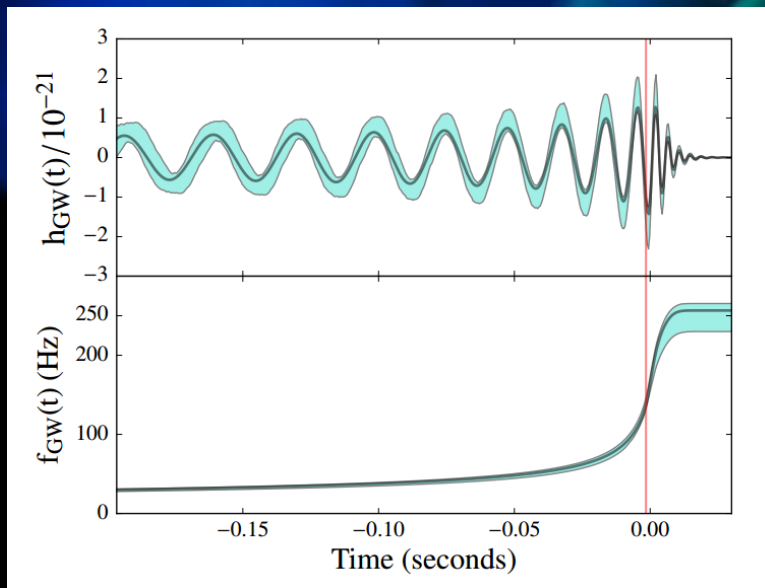


IMR consistency tests

$$h(f) = A(f)e^{i\Phi(f)}$$

$$\Phi(f) = \sum_{k=1}^7 (\varphi_k + \varphi_k^l \log(f)) f^{(5-k)/3} + \sum_{i \neq k} \varphi_i f^i$$

$$\varphi_j \equiv \varphi_j(m_1, m_2, \vec{s}_1, \vec{s}_2) \quad \forall j = k, i$$



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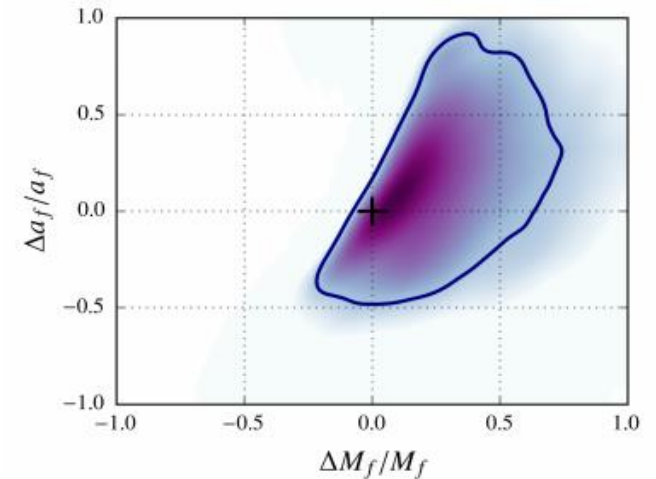
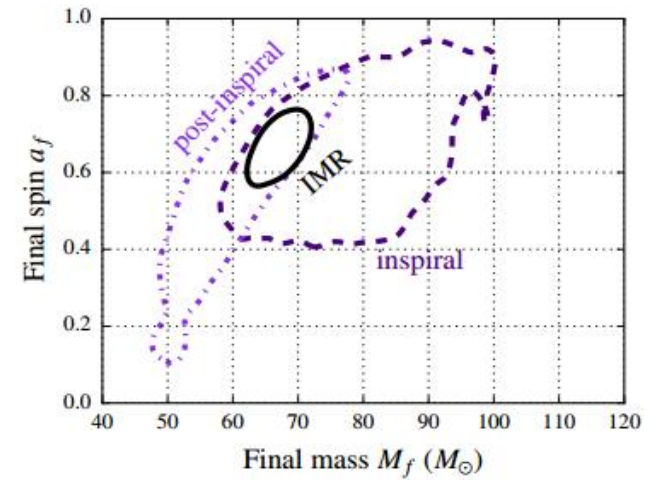
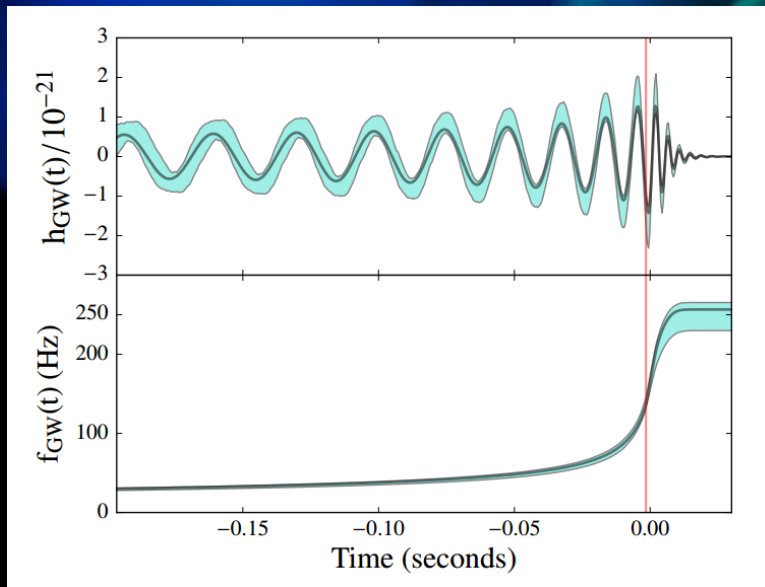


FIG. 3. *Top panel*: 90% confidence regions on the joint posterior distributions for the mass M_f and dimensionless spin a_f of the final compact object predicted from the inspiral (dark violet, dashed) and measured from the post-inspiral (violet, dot-dashed), as well as the result from a full inspiral-merger-ringdown (IMR) analysis (black). *Bottom panel*: Posterior distributions for the parameters $\Delta M_f/M_f$ and $\Delta a_f/a_f$ that describe the fractional difference in the estimates of the final mass and spin from inspiral and post-inspiral parts. The contour shows the 90% confidence region. The plus symbol indicates the expected value (0, 0) in GR.

QNM tests

Can we probe the event horizon
from the ringdown?

QNM tests

Can we probe the event horizon from the ringdown?

- One measured damped mode
- Quality factor can be obtained with different mass and spin, overtones, harmonics.
- Consistent with GR but inconclusive

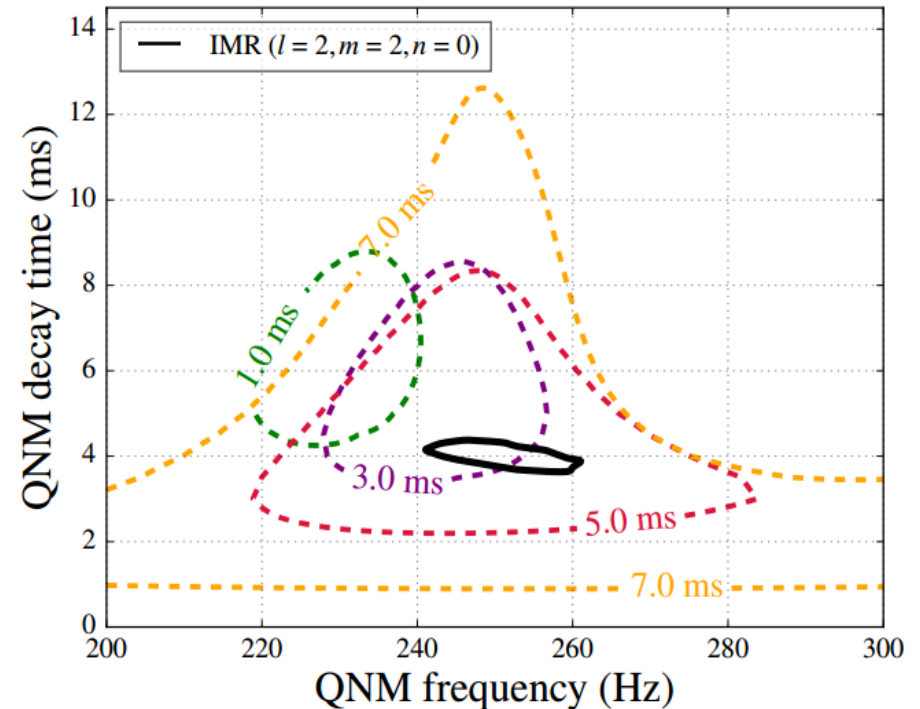


FIG. 4. We show the posterior 90% confidence regions from Bayesian parameter estimation for a damped-sinusoid model, assuming different start-times $t_0 = t_M + 1, 3, 5, 7$ ms, labeled by offset from the merger time t_M of the most-probable waveform from GW150914. The black solid line shows contours of 90% confidence region for the frequency f_0 and decay time τ of the $l = 2, m = 2$ and $n = 0$ (i.e., the least damped) QNM obtained from the inspiral-merger-ringdown waveform for the entire detector's bandwidth.

Deviations from GR waveforms

- Allow for fractional changes with respect to the GR value
- Obtain constraints on possible deviations from GR

$$\hat{\varphi}_j \rightarrow \varphi_j^{\text{GR}} (1 + \delta\hat{\varphi}_j)$$

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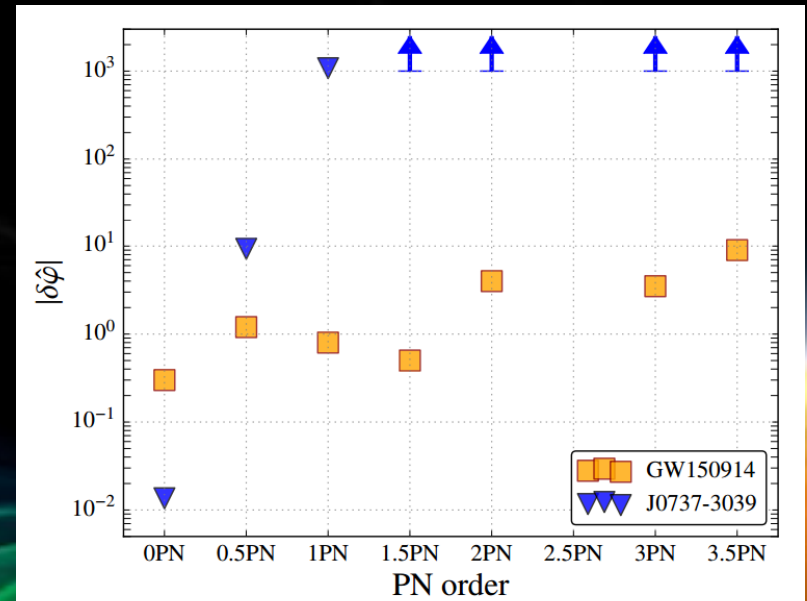
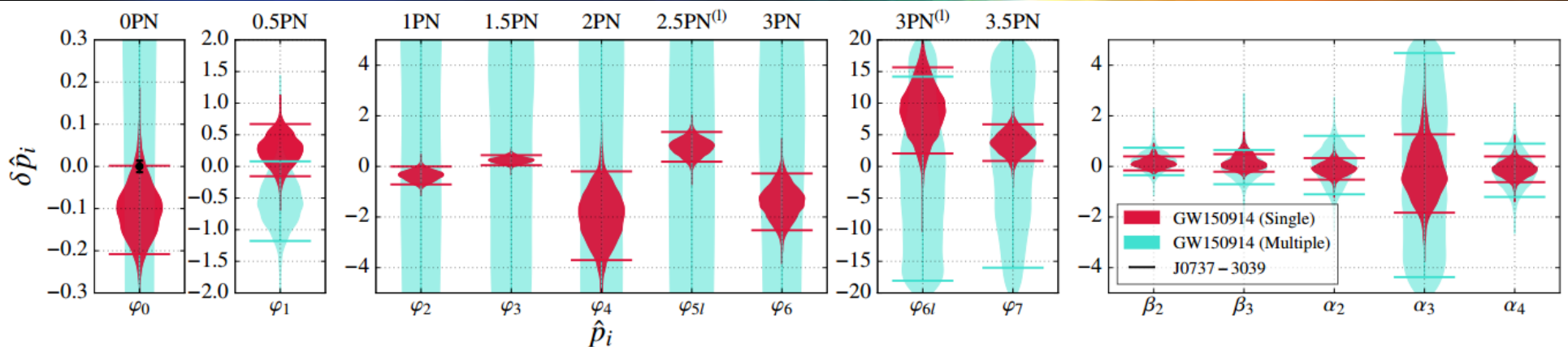


FIG. 6. 90% upper bounds on the fractional variations for the known PN coefficients compared to their known value in GR.



Graviton Compton length

$$\phi(r) = \frac{GM}{r} \left[1 - e^{-r/\lambda_g} \right]$$

$$\lambda_g = \frac{h}{m_g c}$$

$$E^2 = p^2 c^2 + m_g^2 c^4$$

$$\frac{v_g^2}{c^2} = 1 - \frac{h^2 c^2}{\lambda_g^2 E^2}$$

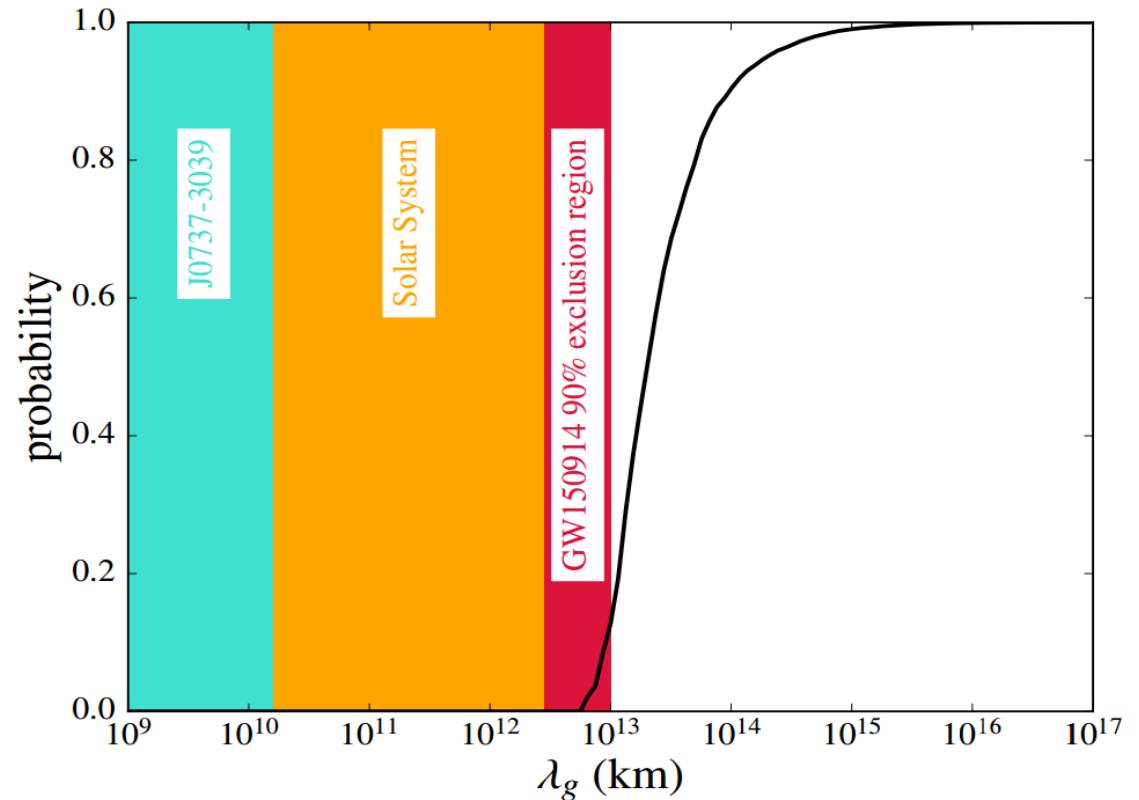
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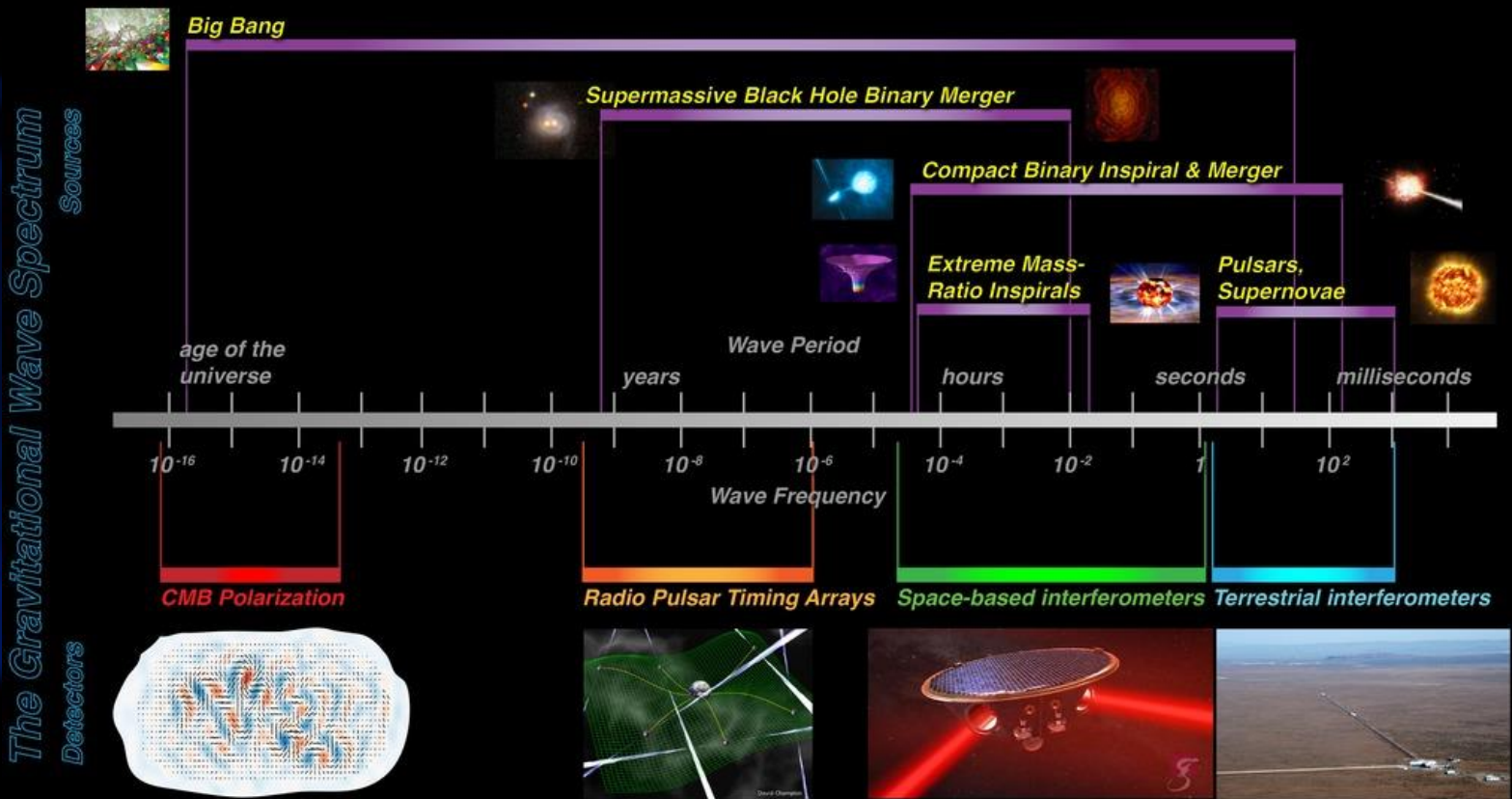
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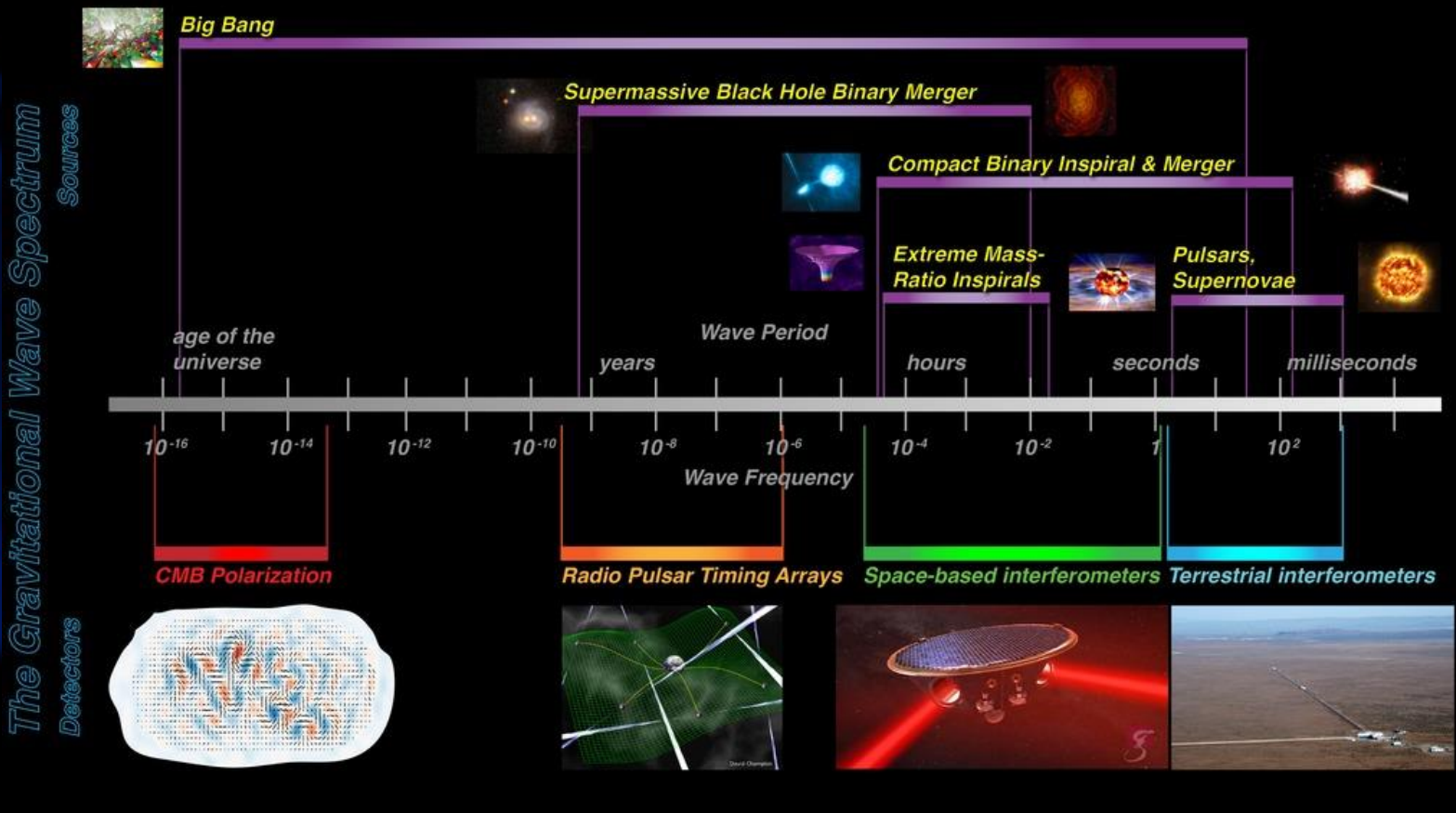
Limit on graviton mass: $m_g \leq 1.2 \times 10^{-22} \text{ eV}/c^2$

B. P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration) Phys.Rev.Lett. 116 (2016) no.22, 221101

LIGO has opened a new window on the universe



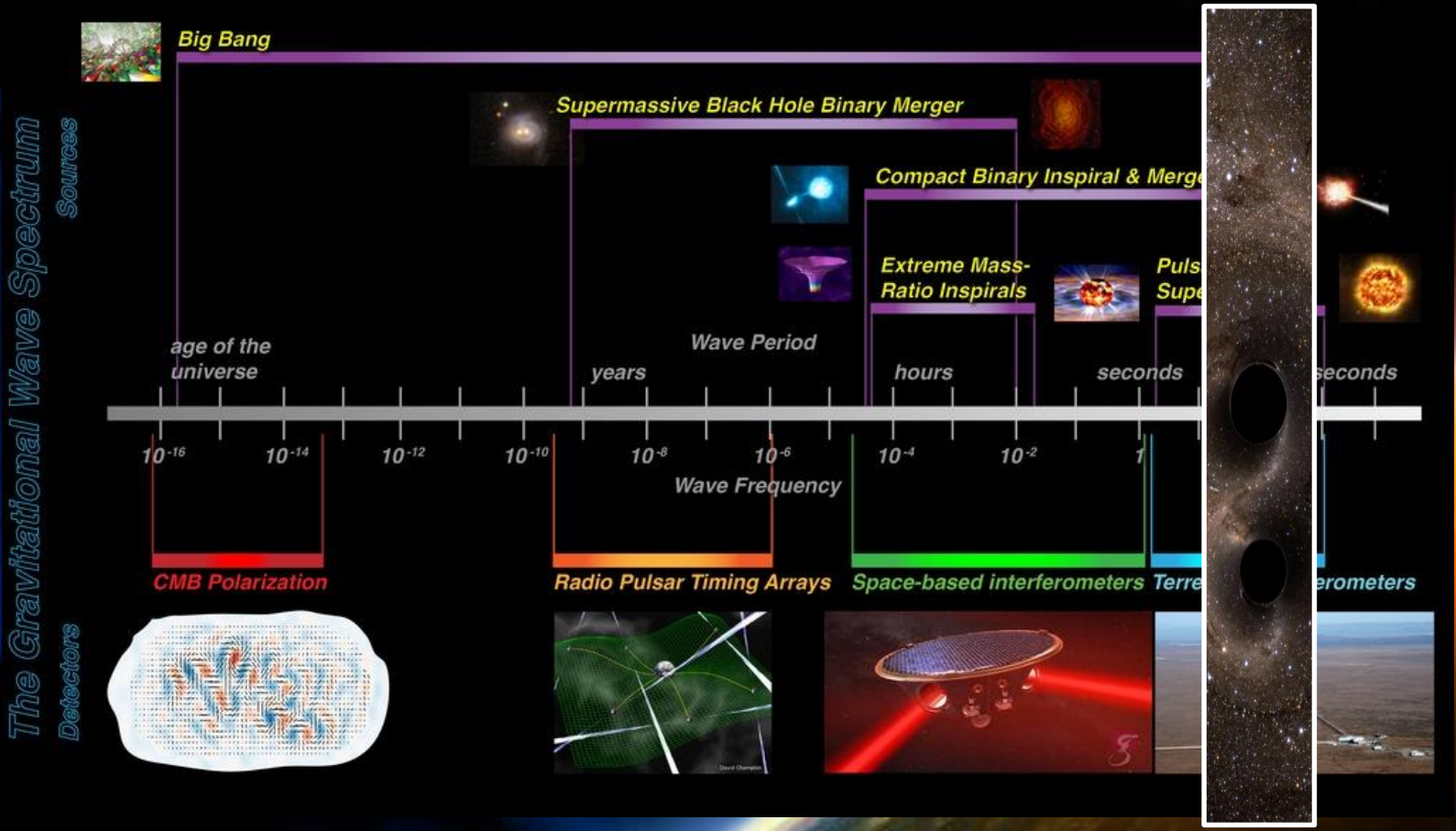
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In the next 5 years, it is likely we will have:

- Hundreds of compact binary coalescence and other source detections
- SNR ~ 100 (GW150914 is ~ 24)
- Observation of fine details of these systems (number, distances, masses, spins, EoS, environment...)

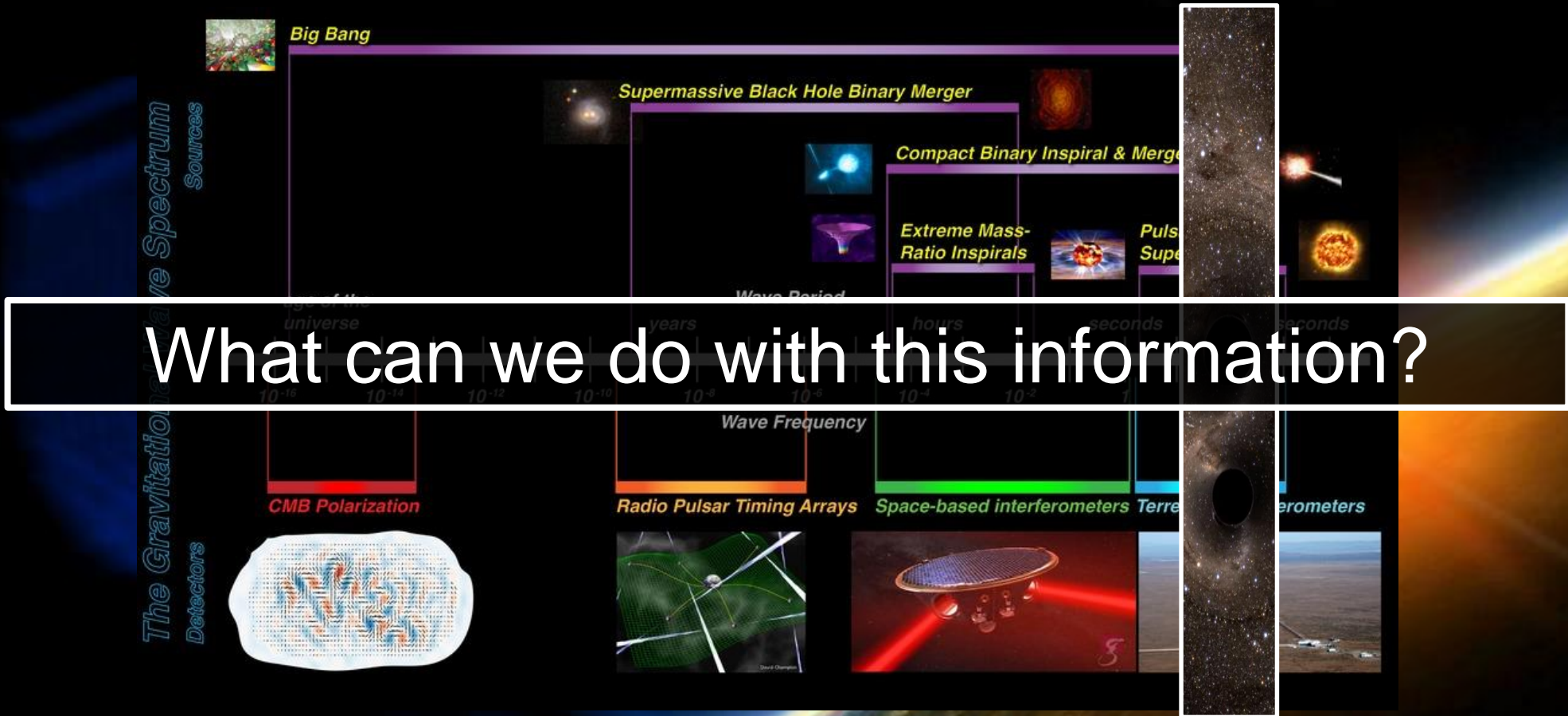
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We have an open window in front of us, let's look what's beyond it!