

Solving ill posed problems

Eduard Feireisl
feireisl@math.cas.cz

Institute of Mathematics, Academy of Sciences of the Czech Republic, Prague
Technische Universität Berlin

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Perfect fluids - Euler system

perfect = inviscid, non(heat) conducting

ρ mass density
 $\mathbf{m} = \rho \mathbf{u}$ momentum
 p pressure
 E energy



Leonhard Paul
Euler
1707–1783

Euler system of gas dynamics

$$\partial_t \rho + \operatorname{div}_x \mathbf{m} = 0$$

$$\partial_t \mathbf{m} + \operatorname{div}_x \left(\frac{\mathbf{m} \otimes \mathbf{m}}{\rho} \right) + \nabla_x p = 0$$

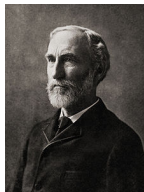
$$\partial_t E + \operatorname{div}_x \left[(E + p) \frac{\mathbf{m}}{\rho} \right] = 0$$

$$E = \frac{1}{2} \frac{|\mathbf{m}|^2}{\rho} + \rho e, \quad e \text{ internal energy}$$

(Incomplete) equation of state (gases)

$$p = (\gamma - 1)\rho e, \quad \gamma \text{ - adiabatic coefficient}$$

Perfect fluids–entropy



Josiah Willard
Gibbs
1839–1903

ϑ (absolute) temperature
 e internal energy

Gibbs' relation

$$\vartheta Ds = De + pD\left(\frac{1}{\rho}\right)$$

Entropy equation

$$\partial_t(\rho s) + \operatorname{div}_x(\mathbf{sm}) = 0$$

“Nobody is perfect”



**Entropy inequality -
Second law of
thermodynamics**

$$\partial_t(\rho s) + \operatorname{div}_x(\mathbf{sm}) \geq 0$$

(Mathematical) troubles with the Euler system



Well posedness

- Solution **exists** for any (physically) admissible data (initial, boundary)
- Solution is **unique** and depends continuously on the data

Known facts concerning well-posedness of the Euler system

- **Local well posedness.** Smooth solutions emanating from smooth data exist on a maximal time interval $T_{\max} > 0$. They are unique and depend continuously on the data
- **Finite time blow up.** There is a “generic” class of data for which $T_{\max} < \infty$. Solutions blow up in a finite time. Typically they become discontinuous shock waves
- **Weak solutions.** Shock waves represent *weak* solutions of the Euler system. Differential operators are understood in the distributional sense. Differential equations replaced by families of integral identities.

Weak solutions to Euler system

$\Omega \subset \mathbb{R}^d$, $d = 1, 2, 3$ fluid domain

Mass conservation

$$\int_0^T \int_{\Omega} [\rho \partial_t \varphi + \mathbf{m} \cdot \nabla_x \varphi] dx = 0$$

Momentum balance

$$\int_0^T \int_{\Omega} \left[\mathbf{m} \cdot \partial \varphi + \mathbf{1}_{\varrho > 0} \frac{\mathbf{m} \otimes \mathbf{m}}{\varrho} : \nabla_x \varphi + p \operatorname{div}_x \varphi \right] dx dt = 0$$

Energy balance

$$\int_0^T \int_{\Omega} \left[E \partial_t \varphi + (E + p) \frac{\mathbf{m}}{\varrho} \cdot \nabla_x \varphi \right] dx dt = 0$$

$$\varphi \in C_c^1((0, T) \times \Omega), \quad \varphi \in C_c^1((0, T) \times \mathbb{T}^d; \mathbb{R}^d)$$

Entropy admissibility condition

$$\int_0^T \int_{\Omega} [\varrho s \partial_t \varphi + \mathbf{s} \mathbf{m} \cdot \nabla_x \varphi] dx dt \leq 0, \quad \varphi \geq 0$$

Ill-posedness of Euler system

Theorem:

Let $\Omega \subset R^d$ be a bounded domain $d = 2, 3$. Let $\varrho_0 > 0$, $\vartheta_0 > 0$ be piecewise constant, arbitrary functions. Then there exists $\mathbf{u}_0 \in L^\infty$ such that the Euler system admits infinitely many *admissible* weak solutions in $(0, T) \times \Omega$ with the initial data $[\varrho_0, \vartheta_0, \mathbf{u}_0]$ satisfying the impermeability boundary condition

$$\mathbf{m} \cdot \mathbf{n}|_{\partial\Omega} = 0$$

Remarks:

- There are examples of Lipschitz (and even C^∞) initial data for which the Euler system admits infinitely many admissible weak solutions
- The result can be extended to the Euler system driven by stochastic forcing

Numerics – consistent approximation

Mass conservation

$$\int_0^T \int_{\Omega} [\varrho_n \partial_t \varphi + \mathbf{m}_n \cdot \nabla_x \varphi] \, dx = E_n^1(\varphi)$$

Momentum balance

$$\int_0^T \int_{\Omega} \left[\mathbf{m}_n \cdot \partial \varphi + \frac{\mathbf{m}_n \otimes \mathbf{m}_n}{\varrho_n} : \nabla_x \varphi + p_n \operatorname{div}_x \varphi \right] \, dx dt = E_n^2(\varphi)$$

Energy balance

$$\int_0^T \int_{\Omega} \left[E_n \partial_t \varphi + (E_n + p_n) \frac{\mathbf{m}_n}{\varrho_n} \cdot \nabla_x \varphi \right] \, dx dt = E_n^3(\varphi)$$

Entropy admissibility condition

$$\int_0^T \int_{\Omega} [\varrho_n s_n \partial_t \varphi + s_n \mathbf{m}_n \cdot \nabla_x \varphi] \, dx dt \leq E_n^4(\varphi), \quad \varphi \geq 0$$

Consistency errors

$$E_n^1(\varphi), E_n^2(\varphi), E_n^3(\varphi), E_n^4(\varphi)$$

Lax equivalence principle

Approximate solutions in conservative entropy variables

ρ mass density
 \mathbf{m} momentum
 $S = \rho s$ total entropy

Formulation for **LINEAR** problems



Peter D. Lax

- **Stability** - uniform bounds on the sequence of approximate solutions $(\rho_n, \mathbf{m}_n, S_n)$
- **Consistency** - vanishing approximation error

$$E_n^i(\varphi) \rightarrow 0 \text{ as } n \rightarrow \infty \text{ for } i = 1, \dots, 4 \text{ and any (smooth) } \varphi$$



- **Convergence** - approximate solutions converge to an exact solution of the Euler system

Lax equivalence principle - nonlinear version ?

Thermodynamics stability

Total energy.

$$E = E(\varrho, \mathbf{m}, S) = \frac{1}{2} \frac{|\mathbf{m}|^2}{\varrho} + \varrho e(\varrho, S)$$

Hypothesis of thermodynamics stability.

$$(\varrho, \mathbf{m}, S) \mapsto E(\varrho, \mathbf{m}, S) \text{ (strictly) convex}$$

Stability estimates for the Euler system

$$\int_{\Omega} E(\varrho_n, \mathbf{m}_n, S_n)(t, \cdot) \, dx \leq \bar{E} \text{ uniformly in } n \text{ and } t \in [0, T]$$

Consistency – vanishing “viscosity”

(Artificial) viscosity approximation

$$\begin{aligned}\partial_t \varrho_n + \operatorname{div}_x \mathbf{m}_n &= \varepsilon_n \Delta_x \varrho_n \\ \partial_t \mathbf{m}_n + \operatorname{div}_x \left(\frac{\mathbf{m}_n \otimes \mathbf{m}_n}{\varrho_n} \right) + \nabla_x p_n &= \varepsilon_n \Delta_x \mathbf{m}_n \\ \partial_t E_n + \operatorname{div}_x \left[(E_n + p_n) \frac{\mathbf{m}_n}{\varrho_n} \right] &= \varepsilon_n \Delta_x E_n\end{aligned}$$

Zero viscosity – turbulent limit

$$\varepsilon_n \searrow 0$$

Weak convergence

- *Pointwise* (ideal) values of functions are replaced by their *integral averages*. This idea is close to the physical concept of *measurement*

$$U \approx \left[\varphi \mapsto \int U\varphi \right]$$

- **Weak convergence**

$$U_n \rightarrow U \text{ weakly} \Leftrightarrow \int U_n \varphi \rightarrow \int U \varphi$$

for any smooth φ

Example

Dirac distribution: $\delta_0 : \varphi \mapsto \varphi(0)$



Paul Adrien Maurice Dirac
[1902-1984]

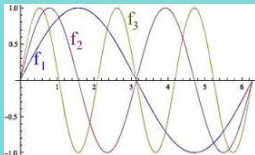
Troubles with weak convergence

Oscillatory solutions - velocity

$U(x) \approx \sin(nx)$, $U \rightarrow 0$ in the sense of averages (weakly)

Oscillatory solutions - kinetic energy

$\frac{1}{2}|U|^2(x) \approx \frac{1}{2}\sin^2(nx) \rightarrow \frac{1}{4} \neq \frac{1}{2}0^2$ in the sense of averages (weakly)



Measure-valued solutions – limits of consistent approximations

$$(t, x) \mapsto \mathbf{U} = (\varrho, \mathbf{m}, S) \approx \langle \mathcal{V}_{t,x}; (\tilde{\varrho}, \tilde{\mathbf{m}}, \tilde{S}) \rangle$$

$\mathcal{V}_{t,x} = \delta_{\mathbf{U}(t,x)}$ – Dirac measure concentrated on $\mathbf{U}_{t,x}$

Young measures

$$\mathbf{U}(t, x) \approx \mathcal{V}_{t,x}$$

\mathcal{V} is a probability measure on the phase space related to the state variables (ϱ, \mathbf{m}, S)



Laurence Chisholm Young
[1905-2000]

$$\mathbf{U}_n \rightarrow \mathbf{U} \text{ weakly, } \int B(\mathbf{U}_n) \varphi \rightarrow \int \overline{B(\mathbf{U})} \varphi$$

$$\overline{B(\mathbf{U})}(t, x) \equiv \langle \mathcal{V}_{t,x}; B(\tilde{\mathbf{U}}) \rangle$$

Lax equivalence principle – nonlinear version

- **Step 1 - stability.**

$(\mathbf{U}_n)_{n=1}^{\infty}$ a stable approximation. Up to a subsequence, \mathbf{U}_n converge (generate) a parametrized (Young) measure $\mathcal{V}_{t,x}$.

- **Step 2 - consistency.** $(\mathbf{U}_n)_{n=1}^{\infty}$ a stable and consistent approximation. Then \mathcal{V} is a measure-valued solution of the Euler system

- **Step 3 - weak-strong uniqueness principle.**

A measure valued and the strong solution emanating from the same data coincide on the life span of the latter. Here coincide means

$$\mathcal{V}_{t,x} = \delta_{\mathbf{u}(t,x)}$$

- **Step 4 - strong convergence**

As the limit is a Dirac mass, the sequence does not oscillate and the convergence is (a.a.) pointwise

Convergence beyond the life span of strong solution

$\mathbf{U}_n = (\varrho_n, \mathbf{m}_n, S_n)$ stable consistent approximation of the Euler system

Convergence to a weak solution



$\mathbf{U}_n \rightarrow \mathbf{U}$ pointwise (a.a.) \Rightarrow the limit $\mathbf{U} = (\varrho, \mathbf{m}, S)$ is a weak solution



the weak limit $\mathbf{U} = (\varrho, \mathbf{m}, S)$ is a weak solution

$\Rightarrow \mathbf{U}_n \rightarrow \mathbf{U}$ pointwise (a.a.)

Conclusion:

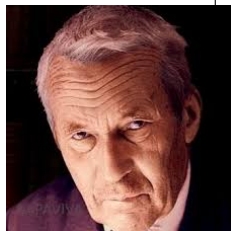
$\mathbf{U}_n \rightarrow \mathbf{U}$ (genuinely) weakly \Rightarrow the limit is not a weak solution of the Euler system

Towards (compressible) turbulence

Scenario I. The viscous approximation $(\varrho_n, \mathbf{m}_n, S_n)$ converges weakly and generates a non-trivial (non monoatomic) Young measure. The weak limit – the barycenter of the Young measure – is not a weak solution of the Euler system

Scenario II - via Kolmogorov hypothesis

The viscous approximation $(\varrho_n, \mathbf{m}_n, S_n)$ is precompact in the strong Lebesgue type topology, in particular contains strongly convergent subsequence. The limit, however, is not unique.



**Andrey
Nikolaevich
Kolmogorov
1903–1987**

Visualizing (computing) the limit object

Scenario I

$$\mathbf{U}_n(t, x) \approx \delta_{\mathbf{U}_n(t, x)}$$

$$\text{Young measure } \langle \mathcal{V}_{t, x}; b(\tilde{\mathbf{U}}) \rangle = \overline{b(\mathbf{U})}(t, x)$$

$$\overline{b(\mathbf{U})} - \text{weak limit of } b(\mathbf{U}_n)$$

S-convergence

$$\frac{1}{N} \sum_{n=1}^N b(\mathbf{U}_n) \rightarrow \overline{b(\mathbf{U})} \text{ strongly}$$

Scenario II - statistical limit

$$\mathbf{U}_n \in C([0, T]; X), X \text{ a suitable Hilbert space}$$

S-convergence

$$\frac{1}{N} \sum_{n=1}^N \delta_{\mathbf{U}_n} \rightarrow \mathbf{U} \text{ strongly}$$

The limit is a statistical solution of the Euler system – a probability measure on the trajectory space

Obstacle problem



**Leonhard Paul
Euler
1707–1783**

Isentropic Euler system driven by temporal white noise

$$d\rho + \operatorname{div}_x \mathbf{m} dt = 0$$

$$d\mathbf{m} + \operatorname{div}_x \left(\frac{\mathbf{m} \otimes \mathbf{m}}{\rho} \right) dt + \nabla_x p(\rho) dt = \sigma dW, \quad p(\rho) = a\rho^\gamma$$

$W = W(t)$ cylindrical Wiener process

$\Omega = \mathbb{R}^d \setminus B$, B compact convex obstacle

Far field conditions

$$\rho \rightarrow \rho_\infty, \quad \mathbf{m} \rightarrow \mathbf{m}_\infty \quad \text{as } |x| \rightarrow \infty.$$

Statistical equivalence

$(\varrho^1, \mathbf{m}^1), (\varrho^2, \mathbf{m}^2)$ fluid motion

$(\varrho^1, \mathbf{m}^1)$ is statistically equivalent to $(\varrho^2, \mathbf{m}^2)$

\Leftrightarrow

■ **Equality of expected values**

$$E[\varrho^1] = E[\varrho^2], E[\mathbf{m}^1] = E[\mathbf{m}^2] \text{ in } (0, T) \times \Omega$$

■ **Equality of expected values of kinetic and internal energy (pressure)**

$$E\left[\frac{|\mathbf{m}^1|^2}{\varrho^1}\right] = E\left[\frac{|\mathbf{m}^2|^2}{\varrho^2}\right], E[p(\varrho^1)] = E[p(\varrho^2)] \text{ in } (0, T) \times \Omega$$

■ **Equality of expected values of angular energy**

$$E\left[\frac{\mathbf{m}^1 \otimes \mathbf{m}^1}{\varrho^1} : (\xi \otimes \xi)\right] = E\left[\frac{\mathbf{m}^2 \otimes \mathbf{m}^2}{\varrho^2} : (\xi \otimes \xi)\right] \text{ in } (0, T) \times \Omega$$

Modeling turbulence by noise ?

Absence of noise in the high Reynolds (low viscosity) limit

- (ϱ, \mathbf{m}) – statistical limit in the vanishing viscosity regime
- (ϱ, \mathbf{m}) statistically equivalent
to a solution of the Euler system driven by the temporal white noise
 \Rightarrow
 (ϱ, \mathbf{m}) is a statistical solution of the *deterministic* Euler system
 \Rightarrow
the noise is not active

Collaborators

D.Breit
(Edinburgh)



M. Hofmanová
(Bielefeld)



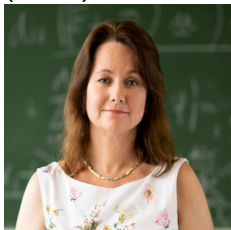
C.Klingenberg
(Wuerzburg)



O.Kreml (Praha)



M.Lukáčová
(Mainz)



S. Markfelder
(Wuerzburg)



H.Mizerová
(Bratislava)



B.She (Praha)

